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2 **Chronology of sedimentation and landscape evolution in the Okavango Rift Zone, a** 3 **developing young rift in southern Africa**

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16 **Key Points:**

- 17 • Nascent continental rifting stages are reflected through sedimentological variations
- 18 • Eolian sand that was formed prior to alluvial incision into the rift is preserved on elevated
 19 surfaces
- 20 • Depositional environments in the incised rift have shifted into alluvial-lacustrine
 21 conditions around the Middle Pleistocene Transition
 22

23 **Abstract**

24 The Kalahari Basin in southern Africa, shaped by subsidence and epeirogeny, features the
 25 Okavango Rift Zone (ORZ) as a significant structural element characterized by diffused
 26 extensional deformation forming a prominent depocenter. This study elucidates the Pleistocene
 27 landscape evolution of the ORZ by examining the chronology of sediment formation and filling
 28 this incipient rift and its surroundings.

29 Modeling of cosmogenic nuclide concentrations in surficial eolian sand from distinct structural
 30 blocks around the ORZ provides insights into sand's residence time on the surface. Sand
 31 formation occurred from ~2.2 to 1.1 Ma, coinciding with regional tectonic events. Notably,
 32 provenance analyses of sand within ORZ's lowermost block where large alluvial fans are found
 33 indicate different source rocks and depositional environments than those of the more elevated
 34 eolian sand. This suggests that the major phase of rift subsidence and the following incision of
 35 alluvial systems into the rift occurred after eolian dune formation. Luminescence dating reveals
 36 that deposition in alluvial fan settings in the incised landscape began not later than ~250 ka, and
 37 that a lacustrine environment existed since at least ~140 ka.

38 The established chronological framework constrains the geomorphological effects of the
 39 different tectono-climatic forces that shaped this nascent rifting area. It highlights two
 40 pronounced stages of landscape development, with the most recent major deformation event in

41 the evolving rift probably occurring during the middle Pleistocene transition (1.2-0.75 Ma). This
42 event is reflected as a striking change in the depositional environments due to the configurational
43 changes accompanying rift progression.

44 **Plain Language Summary**

45 Early stages of continental rifting in the Okavango Rift Zone are described from the perspective
46 of sediment dynamics by constructing a time frame for their evolution. Two major types of
47 sediment and their corresponding time scales are studied. The older sediments are eolian sands
48 that were formed between ~2.2 and 1.1 million years ago, lying today on elevated structural
49 surfaces above the incised rift. Within the subsidized rift that was geomorphologically modified
50 not earlier than 1 million years ago, sediments were deposited by alluvial fans at least since 250
51 thousand years ago, and were followed by a lacustrine environment with alternating hydrological
52 conditions, since at least 140 thousand years ago.

53 **1. Introduction**

54 Tectonic geomorphology incorporates various disciplines and is an evolving field with recent
55 advances in geochronological methods (Keller and DeVecchio 2013; Owen, 2022). While
56 numeric age determinations are being widely used to reconstruct and quantify landscape
57 evolution, multiple processes are involved in the buildup of the analyzed proxies such that their
58 interpretation must be consistent with the geomorphologic context (Watchman and Twidale,
59 2002; Le Dortz et al., 2012; Brown, 2020). Among the most studied features for elucidating and
60 evaluating the effects of tectonics and climate on landscape evolution are alluvial and fluvial
61 subaerial fans, which are common also in rift settings and experience a highly dynamic
62 geomorphological history (Gierłowski-Kordesch, 2010; Warren; 2010; Scheinert et al., 2012;
63 Bowman, 2019).

64 Subaerial fans are generally found and best preserved at the base of mountain fronts within
65 tectonically active zones, where changes in base level are induced by tectonics and variations in
66 climate (Harvey, 2002; Blair and McPherson, 2009). Extensive research has been performed to
67 study their morphologies, involved processes and mechanisms, as well as the components within
68 the system (e.g., nature of sediments, vegetation, lithology) and to reconcile the respective roles
69 of climate and tectonics in their formation (Lustig, 1965; Hooke, 1967; Ritter et al., 1995;
70 Viseras et al., 2003; Terrizzano et al., 2017; Harvey et al., 2018; Bowman, 2019). Multiple
71 models for the environmental evolution of fans were formulated based on disparate methods and
72 over a biased global spatial distribution and settings as most of the primary studies were
73 conducted in the American southwest (Lecce, 1990; Scheinert et al., 2012; Stock, 2013).

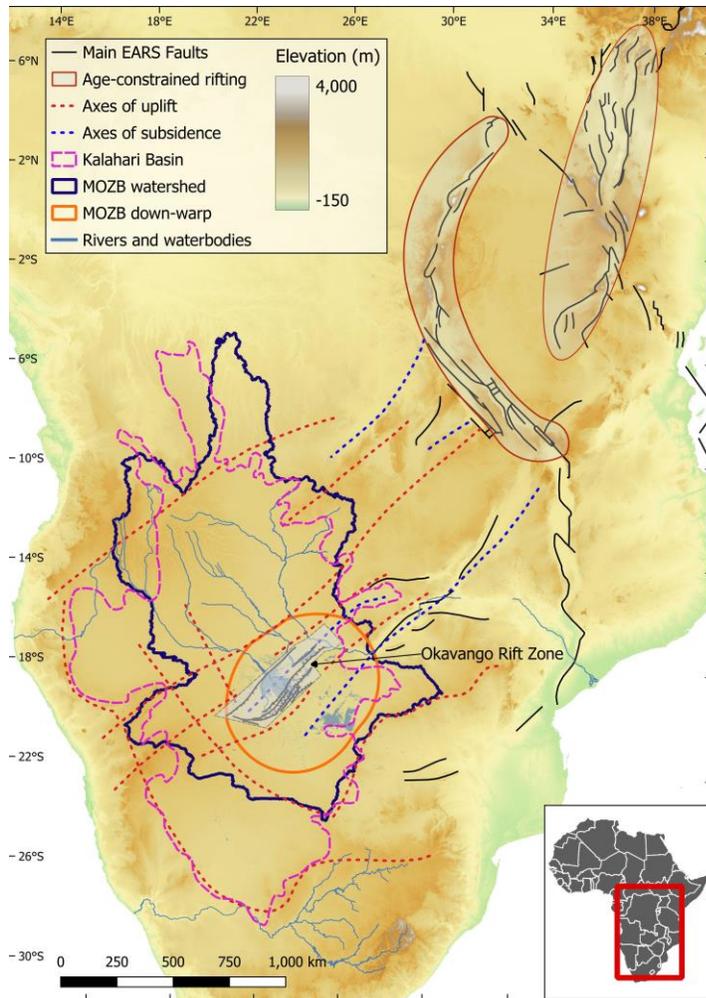
74 The Okavango Rift Zone (ORZ; Figure 1), in interior southern Africa, constitutes an intriguing
75 area to study tectonic geomorphology through the stages involved in the development of alluvial
76 fans and lacustrine/palustrine environments during nascent rifting (Scholz et al., 1976; Kinabo et
77 al., 2007; Wright et al., 2021; Paulssen et al., 2022), where globally unique megafans and paleo-
78 lakes are preserved (Shaw and Thomas, 1992; Burrough and Thomas, 2013). Paleo-lacustrine
79 environments have been thoroughly studied in this area (Moore et al., 2012 and references
80 therein) but, apart from the numerous studies of the Okavango Delta (Podgorski et al., 2013;
81 McCarthy, 2013 and references therein), little attention has been given to the early evolution of
82 fans in central southern Africa (Blair and McPherson, 2009; Wilkinson et al., 2023). Moreover,

83 as the ORZ is bordered by eolian dunes and was subjected to varying zonal climatic interactions
84 (Partridge, 1993; Shaw and Thomas, 1988; Burrough and Thomas, 2013), chronological
85 constraints of landscape evolution that precede the most recent eolian deposition stages are rare
86 (Moore et al., 2012; McCarthy, 2013; Vainer et al., 2021). Therefore, as favored in other regions
87 and settings where recent and earlier fans were studied and compared (DeCelles and Cavazza,
88 1999; Harvey et al., 2005), an investigation of previous phases of landscape development along
89 the ORZ is required.

90 The ORZ lies within the largest continuous sand sheet on Earth and preserves remnants of vast
91 waterbodies (Figure 2) (Grove, 1969; Baillieul 1975; Burrough and Thomas, 2013; McCarthy,
92 2013; Wilkinson et al., 2023). The largest active fan within the ORZ, the Okavango Delta, is
93 characterized by the lowest slope gradient of any other studied subaerial fan and defines one out
94 of three end members of fan types, representing the “losimean” character which is governed by
95 anastomosing meanders (Stanistreet and McCarthy 1993; Bowman, 2019; Wright et al., 2021).
96 Although the Okavango Delta is one of the largest alluvial fans in Africa (McCarthy, 1993) and
97 comprises today the most active depocenter in the Kalahari Basin (Figure 1), isopach maps
98 reveal that, the main depocenter in the Okavango Basin lies ~100 km to the northeast of the
99 Okavango Delta (Figure 3). This area occupies the Linyanti-Chobe Basin within the Chobe
100 Enclave (CE) (Figure 2), which hosts a large alluvial fan that is partially truncated due to
101 tectonic activity (Shaw and Thomas, 1992; McCarthy, 2013; Mokatse et al., 2022a; Wilkinson et
102 al., 2023).

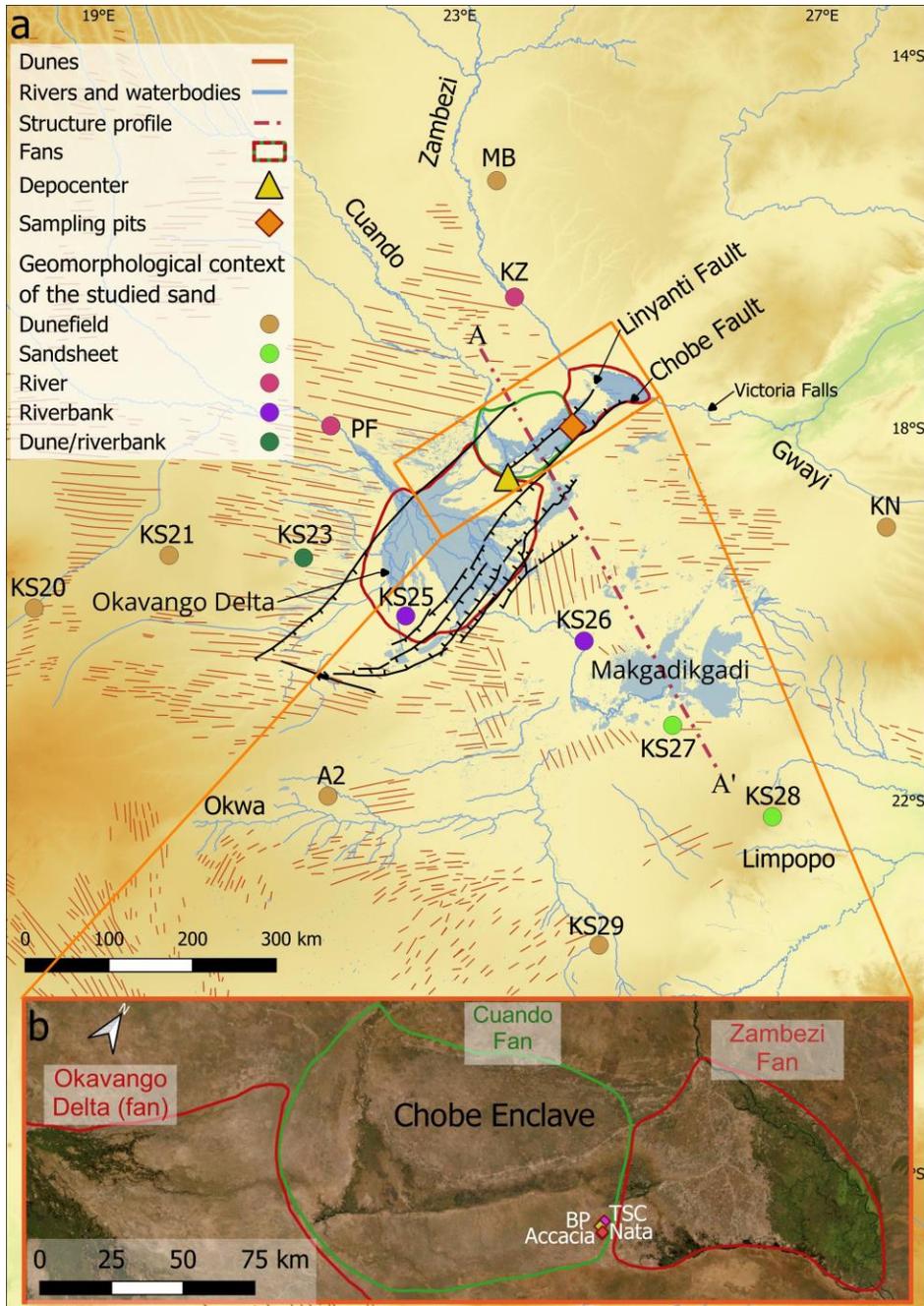
103 While chronological studies of alluvial fans' evolution have shed light on the relationships
104 between their development and tectonics (e.g. Matmon et al., 2006; Placzek et al., 2010; Porat et
105 al., 2010; Terrizzano et al., 2017), the affinity between tectonic settings and eolian accumulation
106 and preservation is poorly constrained and largely unquantified (Cosgrove et al., 2022).
107 Furthermore, it has been postulated that not all fan surfaces are suitable to be dated, particularly
108 at sites where signs of weathering, reworking, and changes in sources are evident (Watchman
109 and Twidale, 2002; Matmon et al., 2005). Due to the prevalence of these processes in the CE
110 (Garzanti et al., 2022) and the uncertainties of available chronological constraints (Moore et al.,
111 2012; McCarthy, 2013), an adjustment of conventional dating methods is needed to construct a
112 chronological framework of this terrane.

113 This study constructs a temporal framework of the geomorphological response to the incipient
114 rifting stages of the ORZ, the southwestern most part of the East African Rift System (EARS).
115 Along the related segments of this rifting system, the latest age constraints for down-warping and
116 faulting are of Pliocene age (Figure 1; Michon et al., 2022). Therefore, rift-related deformation in
117 the ORZ is expected to occur from the Pliocene onwards. Following this assumption, we apply
118 luminescence-based chronologies of buried deposits of the Cuando Megafan, lying in the heart of
119 the ORZ (i.e., CE), with cosmogenic nuclide-based residence time estimates of the surrounding
120 regional eolian sand (Figure 2). These chronometers cover together three relevant temporal
121 orders of magnitude (10^4 - 10^6 yr), providing a time frame for the fluvial-palustrine-lacustrine
122 sediment accumulation in the CE and for sand supply into the central Kalahari. Mineralogical
123 and textural inspections of the sediment are used to characterize the depositional environments
124 and sediment sources. Finally, the data are combined to form a conceptual model of landscape
125 evolution during the early stages of continental rifting.



126

127 Figure 1. Structural elements in southern Africa including the East African Rift Fault System (Chorowicz, 2005),
 128 areas where rifting is chronologically constrained (Michon et al., 2022), and intracratonic structural axes that have
 129 been operated since the Neogene (Haddon and McCarthy, 2005). The shaded area indicates the Okavango Rift Zone
 130 (Bäumle et al., 2019). The background is an SRTM 30-sec DEM (Farr et al., 2007). Inset denotes the extent of the
 131 map on the African continent with its political boundaries.



132

133 Figure 2. (a) Geomorphology of the Makgadikgadi–Okavango–Zambezi Basin including sand dunes (Thomas and
 134 Shaw, 1991), faults (Kinabo et al., 2008; Pastier et al., 2017), hydrogeological features (OpenStreetMap.org;
 135 MapCruzin.com), and the alluvial fans of the Okavango Rift Zone (Wilkinson, 2023). The background is an SRTM
 136 30-sec DEM (Farr et al., 2007) (b) Satellite image of the Chobe enclave (ESRI, 2023) and locations of sampling pits.

137 2. Regional background

138 2.1. Structural Geology

139 The CE is situated between the Congo and Kalahari cratons, overlying extensional and
 140 accretional structures of Proterozoic and Mesozoic age (Dixey, 1956; Doucouré and de

141 Wit, 2003; Oriolo and Becker, 2018). It is bounded by elevated structural arches (Figure 2a;
142 Gumbrecht et al., 2001), and comprises one out of two syntectonic depocenters in the ORZ. It
143 forms part of the Makgadikgadi–Okavango–Zambezi Basin (MOZB), considered as the
144 southwestern-most segment of the EARS (Fairhead and Girdler, 1969; Reeves, 1972; Kinabo et
145 al., 2008). Several tectonic mechanisms were attributed to the sagging of the MOZB including
146 extension resulting from the advancement of the EARS (Modisi et al., 2000; Wright, 2021) inter-
147 cratonic strains causing lithospheric stretching (Pastier et al., 2017; Yu et al., 2017), as well as
148 Internal and peripheral epeirogenic deformation of the Kalahari Basin (Moore, 1999; Vainer et
149 al., 2021; Mokatse et al., 2022a).

150 The structural trough of the ORZ is controlled by NE-SW normal to dextral strike-slip faults
151 forming half-graben structures (Figure 4; Modisi et al., 2000; Kinabo et al., 2007; Kinabo et al.,
152 2008) that accommodate an endorheic hydrological system (Figure 2a). This system contains
153 sediments that were transported into the basin, and then recycled, weathered, and eventually
154 diagenetically altered or cemented by secondary minerals (Huntsman-Mapila et al., 2005; Vainer
155 et al., 2021; Garzanti et al., 2022; Mokatse et al., 2023). A major change in the organization of
156 fluvial systems is assumed to have occurred in the early Pleistocene when the upper part of the
157 Zambezi River was captured by its middle part, diverting flow from the terminal basin into the
158 lower base level of the Indian Ocean (Moore et al., 2012; Vainer et al., 2021). Today, the
159 Okavango Basin is occupied by divide fans that are characterized by hydrological links with
160 neighboring basins (Wilkinson et al., 2023). Within the Okavango Basin, a series of tectonically
161 generated reorganizations of the fluvio-lacustrine system occurred throughout the Quaternary
162 (Moore et al., 2007, Schmidt et al., 2023). These events resulted in the deposition of mixed
163 alluvial, fluvial, palustrine, and evaporite sediments, surrounded by eolian deposits on the
164 elevated basin margins.

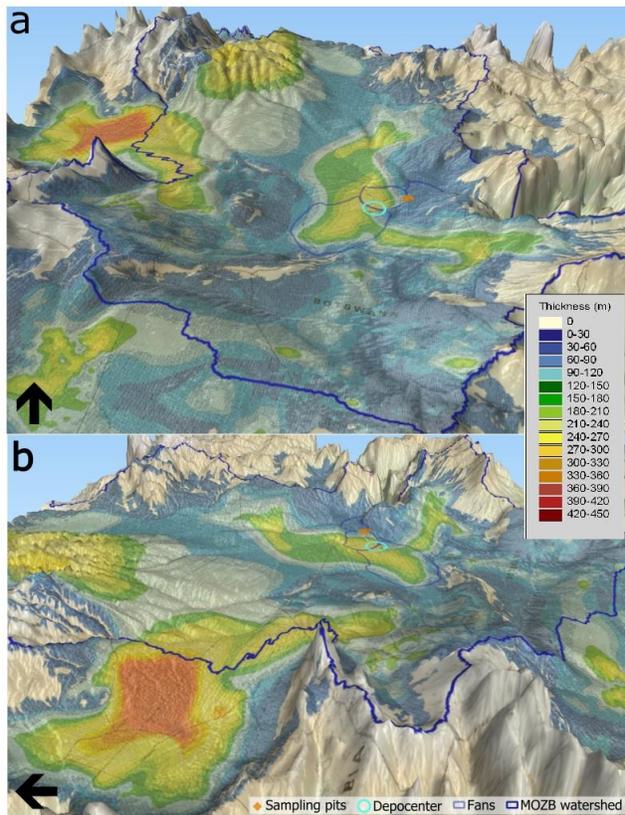
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166 2.2. Chronostratigraphy

167 Cenozoic fluvio-lacustrine and eolian sediments (Kalahari Group) up to 300 m-thick fill the
168 MOZB (Haddon and McCarthy, 2005; Podgorski et al., 2013) that was established to form a
169 similar configuration as today around 2.5 Ma (Du Toit, 1933; Day et al., 2009; Cotterill & De
170 Wit, 2011; Vainer et al., 2021). However, a differential structural geometry probably already
171 existed during the Pliocene (Vainer et al., 2021) comprising three sub-basins at the sub-surface,
172 with the thickest depocenter located between the Okavango and Linyanti-Chobe (Figure 3).
173 Basement rocks, mostly Proterozoic volcanic and metasedimentary rocks, as well as Mesozoic
174 metasediments of the Karoo Supergroup and Lower Jurassic basalts (Modisi, 2000) are rarely
175 exposed in marginal and deformed areas within the MOZB (Figure 4). The lithology and
176 composition of the Kalahari Group within the MOZB are known from limited boreholes mainly
177 drilled in the Okavango sub-basin. They reveal the prevalence of sand derived from both local
178 and distant, mostly northerly, source areas, with variable proportions of silt, clay, and carbonates
179 that underwent in places a high degree of chemical weathering (Huntsman-Mapila et al., 2005;
180 Vainer et al., 2021).

181 The only numeric ages for the earliest deposition in the MOZB are derived from cosmogenic
182 nuclide-based burial dating of two depth profiles in the western Okavango Basin. Ages are 3.06

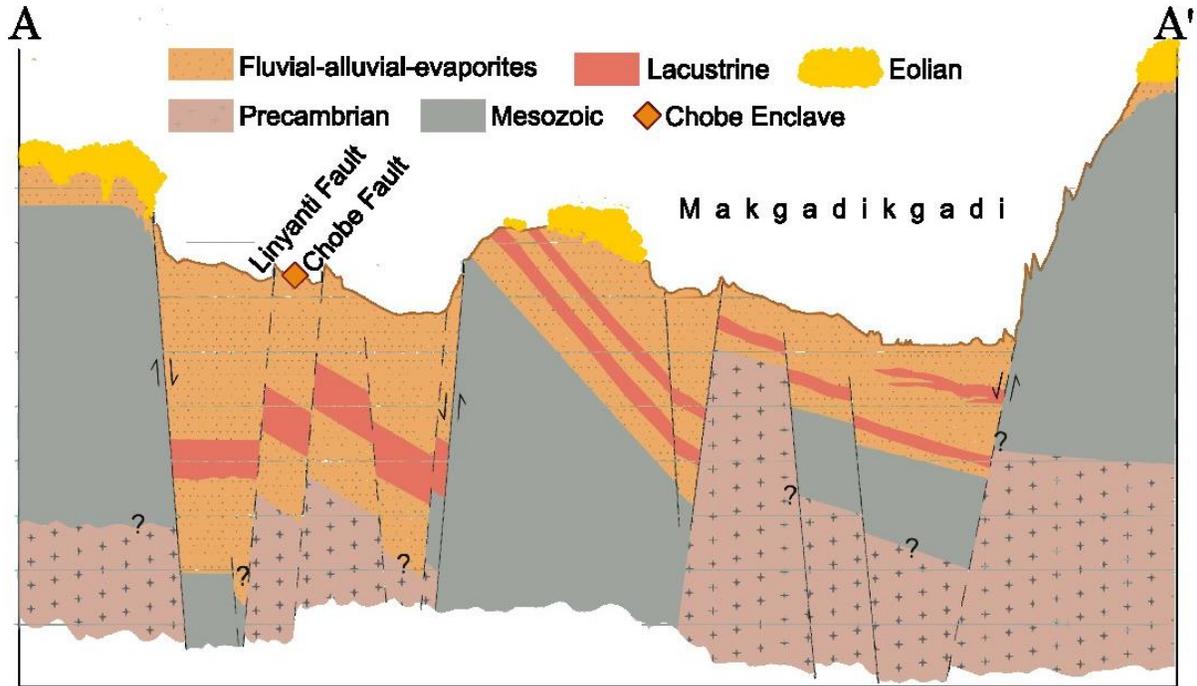
183 $+4.4/-0.46$ Ma at the base of the upthrown block and $3.35^{+0.39}/-0.26$ Ma in the downthrown block,
 184 where basal strata were undatable. The uppermost consolidated sediments at these upthrown and
 185 downthrown sites were buried at $1.12^{+0.13}/-0.12$ and $1.34^{+0.16}/-0.14$ Ma, respectively. These capping
 186 ages were suggested to represent the onset of eolian dominance for sand transport and deposition
 187 (Vainer et al., 2021).



188
 189 Figure 3. Isopach map of the estimated Cenozoic sedimentation (Kalahari Group) (Haddon and McCarthy, 2005)
 190 overlaying 3D elevation model constructed from SRTM version 3.0 Global 1 arc second (Farr et al., 2007) (not to
 191 scale). The blue outline marks the MOZB watershed (a) looking northward. (b) Looking eastward.

192 The surficial fluvio-lacustrine features of the MOZB represent several depositional phases,
 193 resulting from changes in fluvial configuration and deposition on top of older alluvium (Thomas
 194 and Shaw, 1991). Various materials collected mostly from ridges and pan floors were dated by
 195 applying luminescence and ^{14}C dating techniques and were interpreted to represent alternating
 196 wet and dry stages (Burrough et al., 2007; Burrough and Thomas, 2013). Earliest ages, as old as
 197 280 ka, were evoked from a limited number of samples ($n=3$) in these studies, leaving a
 198 noticeable age gap with the ~ 1.1 Ma burial ages at the western MOZB. Successive lacustrine
 199 highstands were inferred to occur between 131 ± 11 and 92 ± 2 ka, with another phase centered
 200 around 64 ka, and fluctuating conditions between 40 ka and the present. The period between 115
 201 and 95 ka coincides with eolian accumulation in dunes at the northeastern MOZB (Stokes et al.,
 202 1998) and the younger inferred highstand stages are coeval with dune buildup in northwestern
 203 MOZB areas (Thomas et al., 2000). Within the CE, ages of depth profiles in elevated ridges
 204 range between 23.4 ± 1.6 and 1.9 ± 0.3 ka (Burrough and Thomas, 2008; Mokatse et al., 2022),
 205 while quartz in carbonate rocks was dated at 48.2 ± 9.6 ka and buried floodplain sands to ~ 50 ka

206 (Diaz et al., 2019). Combined, these ages indicate a dynamic hydrogeological fluvio-lacustrine
 207 environment at least since ~280 Ka.



208
 209 Figure 4. Surface topography (vertical exaggeration ≈ 500) and conceptual geological cross-section (not to scale)
 210 through the Makgadikgadi–Okavango–Zambezi Basin (after Bäuml et al., 2019). The horizontal axis corresponds
 211 to the magenta/red, semi-dashed line in Figures 2a and 6a, respectively.

212 3. Sampling, sample preparation, and analyses

213 3.1 Optically stimulated luminescence (OSL) dating

214 Four pits were excavated to depths of 8-10 m between the Chobe and Linyanti faults within the
215 Cuando alluvial fan. From north to south, they are termed TSC, BP, Nata, and Accacia (Figure
216 2b). Sampling was made via light-sealed tubes hammered into pit walls or by auger downwards
217 from the bottom of the pit for the deepest samples. Additionally, 1-3 surficial quartz-containing
218 carbonate samples from each site were sampled to constrain the timing of diagenetic processes
219 (i.e. carbonate cementation). These carbonate samples were then cut under subdued red lighting
220 to remove any material that was exposed to light and the remaining inner, un-exposed parts were
221 further treated for analyses.

222 Organic matter and carbonate precipitates were removed from the 180-212 μm size fraction of
223 all samples with H_2O_2 and 10% HCl, respectively. Using dense-liquid (sodium polytungstate)
224 separation, the fraction between 2.62 and 2.70 g cm^{-3} was etched for 40 min with 40% HF acid to
225 purify the quartz separates and remove the outer layer affected by alpha-radiation. Subsequently,
226 the samples were treated with 10% HCl for >1 h for dissolution of potential Ca-fluoride
227 precipitates (Diza et al., 2016).

228 The equivalent dose (D_e) of quartz samples was determined using a single-aliquot regeneration
229 dose (SAR) protocol, including four regeneration doses, a zero dose and a repeated dose (Murray
230 and Wintle, 2000). Preheat plateau and dose recovery tests were carried out in the temperature
231 range 180–260 $^{\circ}\text{C}$ for three samples from different profiles to identify optimal preheat
232 conditions. Dose response data constructed from the first 0.7 s of the decay curve (corrected for a
233 background estimated from the last 2 s of this curve) were fitted with an exponential plus linear
234 function in the Analyst software (v4.57; Duller, 2015). Further technical details are given in the
235 Supporting Information. Radioelement concentrations (K, Th, U) were quantified by high-
236 resolution γ -ray spectrometry on ground samples with a mass of ~60–80 g that were stored in a
237 sealed container for at least four weeks prior to analysis.

238 3.2. Cosmogenic nuclides

239 Thirteen surficial sand samples from river sands and eolian dunes were processed to determine
240 their ^{10}Be and ^{26}Al content. Seven samples were collected from within the MOZB, and six from
241 its periphery (Figure 2a). The sand was sieved, and the 250-850 μm size fraction underwent
242 sample leaching by aqua regia solution, magnetic separation, and sequential HF + HNO_3 etching
243 (Kohl and Nishiizumi, 1992). Major elements were measured with inductively coupled plasma
244 optical emission spectrometry to verify low concentrations of elements such as Al or Ti.
245 Following spiking and ion-exchange chromatography, isotopic ratios of oxidized targets were
246 measured by accelerator mass spectrometry at Centre de Recherche et d'Enseignement des
247 Géosciences de l'Environnement (CEREGE), France. Isotopic ratios of the in-house standards
248 used for measurements were 7.40×10^{-12} and 1.91×10^{-11} for $^{26}\text{Al}/^{27}\text{Al}$ and $^{10}\text{Be}/^9\text{Be}$, respectively.
249 Procedural blank values were in the range between 8.6×10^{-16} and 1.66×10^{-15} for $^{26}\text{Al}/^{27}\text{Al}$ and
250 between 2.68×10^{-16} and 6.97×10^{-15} for $^{10}\text{Be}/^9\text{Be}$.

251 3.3. Heavy minerals

252 Two samples from the base of each pit were analyzed for their heavy-mineral assemblage to
253 detect the relative sediment contribution from the Cuando and Zambezi rivers. Heavy minerals
254 were separated by centrifuging in sodium polytungstate (density $\sim 2.90 \text{ g cm}^{-3}$) and recovered by
255 partial freezing with liquid nitrogen. More than 200 transparent heavy-mineral grains were point-
256 counted on grain mounts at suitable regular spacing under the petrographic microscope to
257 minimize the bias caused by grain counting (Garzanti and Andò, 2019). Grains of uncertain
258 identification were checked with Raman spectroscopy (Andò and Garzanti, 2014). Based on the
259 percentage of transparent heavy minerals (tHM), tHM suites are defined as “extremely poor”
260 (tHMC < 0.1) and “very poor” (tHMC 0.1-0.5; Garzanti and Andò, 2007). The ZTR index is the
261 sum of zircon, tourmaline, and rutile over total tHM (Hubert, 1962) and is classically used to
262 estimate sediment “durability” (i.e., the extent of recycling; Garzanti, 2017).

263

264 **4. Modelling**

265 4.1. OSL dating

266 Due to the low radioactivity of the sand (Supporting information, Table S1), the cosmic dose rate
267 makes up a significant contribution to the total dose rate ($\sim 25\text{--}65\%$, depending on the sample).
268 Therefore, an assessment of the time-dependent cosmic dose rate has been made (Supporting
269 information, Figure S1). This has been estimated step by step by first accounting for the
270 youngest samples taken from the carbonate unit, and hereafter calculating the cosmic dose rates
271 of the samples taken from the sand unit below the carbonate layers, considering the age of the
272 younger samples. Also, carbonate precipitation that can influence the dose rate calculation was
273 considered (Supporting information, Table S2).

274 Two models were considered for carbonate units (Supporting information, Table S2). One
275 assumes a short time between sand accumulation, carbonate precipitation, and pore filling, thus
276 no modelling of carbonate emplacement over time and its influence on dose rate evolution is
277 carried out (cf. Nathan and Mauz, 2008; Mauz and Hoffmann, 2014; Kreutzer et al., 2019). In
278 case this assumption does not apply, the alternative approach was to perform sensitivity tests by
279 contrasting the conventional OSL ages with those resulting from modelling the time-dependent
280 dose rate using the RCarb model (Mauz and Hoffmann, 2014; Kreutzer et al., 2019). Also, the
281 possibility of U uptake during carbonate precipitation was considered, but given that ^{226}Ra and
282 daughter nuclides contribute $>70\%$ of the total β - and γ -dose rate it was concluded that
283 modelling the time-dependent dose rate with reference to poorly constrained assumptions would
284 probably not result in substantially changed ages, necessitating a revision of the environmental
285 interpretation (Degering and Degering, 2020).

286 The age information obtained for the samples from the carbonate units was considered for
287 estimating the cosmic dose rate applicable to the samples extracted from the sand units below.
288 This approach of individually modelling the cosmic dose rate for each sample based on age
289 information from stratigraphically younger samples was contrasted with the simple (and more
290 common) approach of assuming a constant sedimentation rate (Supporting information, Table
291 S3). This comparison reveals that the age estimates in both ways are indistinguishable at the 1σ
292 confidence level. Therefore, the ages derived from a constant sedimentation rate are used as the

293 TSC profile shows an almost linear increase in age with depth and because adopting one
294 criterion consistently across the entire profile is simpler and more straightforward whenever age
295 inversions occur.

296 Another factor causing potential OSL age inaccuracy is the internal dose rate of quartz grains.
297 Especially in low-dose-rate environments, such as the CE, the contribution from the internal dose
298 rate to the total dose rate can be significant. There are only a few previous studies on measured
299 values of internal radioelement concentrations of quartz, and these yielded variable results (e.g.,
300 Vandenberghe et al., 2008; Steup, 2015). As the U and Th content of quartz seems to scatter to a
301 much larger extent than, e.g., the K content of K-feldspar, it may not seem reasonable to assume
302 a universal value for the internal quartz dose rate. Within the scope of this study, it was not
303 possible to quantify the internal dose rate of the samples. A previous publication including OSL
304 ages of comparable samples from the Okavango Basin does not report analytical values for the
305 internal quartz dose rate but states that this dose rate contribution does not change the
306 interpretation of the results (Burrough et al., 2009). Thus, zero internal quartz dose rate was
307 assumed, with the implication that age estimates might be younger, should there be a
308 significantly large internal dose rate from quartz grains.

309 The D_e used to estimate the burial age was derived from applying the Central Age Model (CAM;
310 Galbraith et al., 1999). Ages were calculated with the DRAC software (v1.2; Durcan et al.,
311 2015).

312

313 4.2. Surficial residence time

314 The surficial residence time of the sand was assessed through numerical modelling simulating
315 the accumulation of cosmogenic nuclides under eolian, fluvial, or lacustrine settings by applying
316 the Cosmolian model (Vainer et al., 2018a; Vainer and Ben Dor, 2021; Vainer et al., 2022).
317 Simulations commence with the build-up of cosmogenic nuclides during erosion of source areas
318 that are represented by the coordinates of the headwater of the sample specific-sub basin (Table
319 S4, Supporting Information). Values of 3, 9, and 20 m Ma^{-1} were considered, following Regard
320 et al. (2016) and references therein for erosion rates in the source areas of the sand (Garzanti et
321 al., 2022). Simulations then reproduce the vertical component of sand grains during transport by
322 randomly changing the overburden by 20 cm increments, with twenty-four combinations of
323 possibilities of boundary conditions. The average latitude and altitude values of each sub-basin
324 of a sample define the parameters for cosmogenic nuclides' production rates during transport and
325 are changing only as a function of changing depth. Three amplitudes of 1, 10, and 25 m are used
326 as different boundary conditions to encompass the range of dune heights (Lancaster, 1981;
327 Stokes et al., 1998), and shallow waterbodies in the MOZB (Moore et al., 2012). The retention
328 time at each depth increment is based on a probability function, constructed from dated eolian
329 (Lancaster et al., 2016 and references therein), lacustrine (Huntsman-Mapila et al., 2006;
330 Burrough and Thomas, 2008; Burrough and Thomas, 2013), fluvial (Shaw et al., 1992; Brook et
331 al., 2008), and pluvial (Nash et al., 1981) sediments. Two datasets were constructed to form two
332 probability functions by dividing the OSL/TL/ ^{14}C ages of buried sediments with the
333 corresponding depth of each dated sample. This conversion from age to vertical displacement
334 rates was applied to 54 eolian samples and 35 fluvial, lacustrine, and palustrine (FLP) samples.

335 These two datasets were further modified to account for the possible bias stemming from
336 oversampling shallow deposits by removing the fastest 10% from each dataset.

337 Each simulation during which the build-up of cosmogenic nuclides occurred lasted for 5 My and
338 was repeated with the same conditions 10,000 times. The duration of various successful
339 simulations in which convergence between the simulated and measured concentrations of both
340 ^{26}Al and ^{10}Be occurred have been summarized and are interpreted as the most probable timing
341 since the modelled sand was introduced into the landscape (Vainer et al., 2022).

342

343 5. Results

344 5.1 OSL dating

345 The results of D_e estimates are shown in Table S1 in Supporting Information and calculated ages
346 are shown in Table 1 and illustrated in Figure 5. The assumption that the ^{238}U decay chain is in
347 secular equilibrium was followed for samples that were taken from carbonate-cemented units.
348 Ages of 29 ± 3 ka and 81 ± 7 ka for samples TSC 0.8 and TSC 2.0 were obtained, respectively.
349 Samples NATA 0.8 and NATA 0.8B, taken from the same depth of 0.8 m, yielded ages of 74 ± 7
350 ka and 58 ± 6 ka, respectively, not overlapping at the 1σ confidence level. A sample taken 1.2 m
351 below (NATA 2.0), however, produced an age of 64 ± 6 ka, consistent with the dating results of
352 both overlying samples and with an age of 55 ± 6 ka obtained at 0.6m depth by Mokatse et al.,
353 2022a. At the ACA site, the sample from 0.8 m depth yielded an age of 54 ± 5 ka, and the
354 deepest carbonate bed at 4.6 m depth is dated to 163 ± 16 ka, which is synchronous (within
355 uncertainty) with the deposition of the sand at 5.8 m depth at 137 ± 11 .

356 Table 1. Dose rate assessment and age calculation. A value of 10 ± 3 wt% water was assigned to all samples expect
357 from waterlogged samples where 21 ± 3 wt% values were assigned. The cosmic dose rate of the sand samples
358 (below the carbonate layers) was modelled according to the thickness of sand and carbonate units and the time of
359 emplacement of overlying layers. See Supporting Information for further details.

Sample	Generalized content	U [ppm]	Th [ppm]	K [%]	Cosmic \dot{D} [Gy ka ⁻¹]	Total \dot{D} [Gy ka ⁻¹]	CAM D_e [Gy]	Age [ka]
ACA 0.8	Carbonate + sand	0.35 ± 0.09	0.29 ± 0.18	0.045 ± 0.005	0.219 ± 0.020	0.348 ± 0.027	18.71 ± 0.82	54 ± 5
ACA 4.6	Carbonate + sand	0.83 ± 0.22	1.08 ± 0.33	0.100 ± 0.014	0.159 ± 0.016	0.482 ± 0.040	78.51 ± 4.44	163 ± 16
ACA 5.8	Sand	0.85 ± 0.10	1.59 ± 0.09	0.158 ± 0.021	0.142 ± 0.014	0.551 ± 0.027	75.27 ± 4.77	137 ± 11
ACA 6.3	Sand	0.80 ± 0.11	1.53 ± 0.28	0.088 ± 0.013	0.138 ± 0.014	0.472 ± 0.027	71.92 ± 4.04	152 ± 12
ACA 7.0	Waterlogged sand	0.50 ± 0.10	0.93 ± 0.12	0.049 ± 0.009	0.133 ± 0.013	0.314 ± 0.020	25.46 ± 1.46	81 ± 7
BP 5.5	Sand	0.44 ± 0.12	0.80 ± 0.12	0.116 ± 0.016	0.142 ± 0.014	0.383 ± 0.026	57.21 ± 2.53	150 ± 12
BP 6.2	Sand	0.36 ± 0.13	0.47 ± 0.19	0.061 ± 0.009	0.136 ± 0.014	0.292 ± 0.026	51.32 ± 3.58	150 ± 16
BP 6.8	Sand	0.34 ± 0.04	0.49 ± 0.04	0.053 ± 0.005	0.132 ± 0.013	0.278 ± 0.023	43.86 ± 2.01	201 ± 20

		0.11	0.15	0.008			2.40	
BP 8.3	Sand	0.16 ± 0.08	0.31 ± 0.12	0.019 ± 0.004	0.122 ± 0.012	0.190 ± 0.018	55.99 ± 3.39	252 ± 27
BP 10	Waterlogged sand	0.71 ± 0.06	1.34 ± 0.18	0.094 ± 0.012	0.112 ± 0.011	0.389 ± 0.018	47.87 ± 2.42	203 ± 14
NATA 0.8	Carbonate + sand	0.22 ± 0.10	0.16 ± 0.27	0.052 ± 0.006	0.212 ± 0.021	0.312 ± 0.029	78.87 ± 3.99	74 ± 7
NATA 0.8 B	Carbonate + sand	0.36 ± 0.11	0.07 ± 0.26	0.055 ± 0.010	0.212 ± 0.021	0.338 ± 0.030	23.14 ± 0.87	58 ± 6
NATA 2.0	Carbonate + sand	0.51 ± 0.12	0.34 ± 0.08	0.050 ± 0.008	0.191 ± 0.019	0.360 ± 0.027	19.49 ± 0.87	64 ± 6
NATA 3.5	Sand	0.31 ± 0.07	0.95 ± 0.16	0.095 ± 0.015	0.171 ± 0.017	0.376 ± 0.024	23.05 ± 0.99	143 ± 11
NATA 4.5	Waterlogged sand	0.31 ± 0.10	0.78 ± 0.02	0.048 ± 0.007	0.161 ± 0.016	0.298 ± 0.021	53.74 ± 2.46	180 ± 16
NATA 5.0	Waterlogged sand	0.35 ± 0.10	0.48 ± 0.26	0.036 ± 0.006	0.157 ± 0.016	0.275 ± 0.023	53.53 ± 2.77	94 ± 10
TSC 0.8	Carbonate + sand	0.52 ± 0.09	0.54 ± 0.26	0.099 ± 0.017	0.212 ± 0.021	0.438 ± 0.029	25.81 ± 1.43	29 ± 3
TSC 2.0	Carbonate + sand	0.58 ± 0.14	0.40 ± 0.27	0.062 ± 0.010	0.191 ± 0.019	0.388 ± 0.032	12.55 ± 0.79	81 ± 7
TSC 3.1	Sand	0.69 ± 0.11	1.44 ± 0.10	0.136 ± 0.021	0.176 ± 0.018	0.524 ± 0.029	31.49 ± 1.31	116 ± 9
TSC 4.0	Sand	0.23 ± 0.09	0.56 ± 0.17	0.036 ± 0.006	0.166 ± 0.017	0.279 ± 0.023	60.89 ± 3.19	139 ± 14
TSC 6.0	Sand	0.27 ± 0.11	0.58 ± 0.12	0.029 ± 0.005	0.149 ± 0.015	0.265 ± 0.023	38.76 ± 2.30	199 ± 19

360 The ages of sand that underlie the carbonate layers were calculated assuming cosmic dose rate
361 production during constant sedimentation rate and range between 252 ± 27 and 116 ± 9 ka (Table
362 1). Ages generally follow a stratigraphic order, apart from samples ACA 7.0, BP 10, and NATA
363 5.0. These outliers were saturated in water when sampled (discussed in Supporting Information).

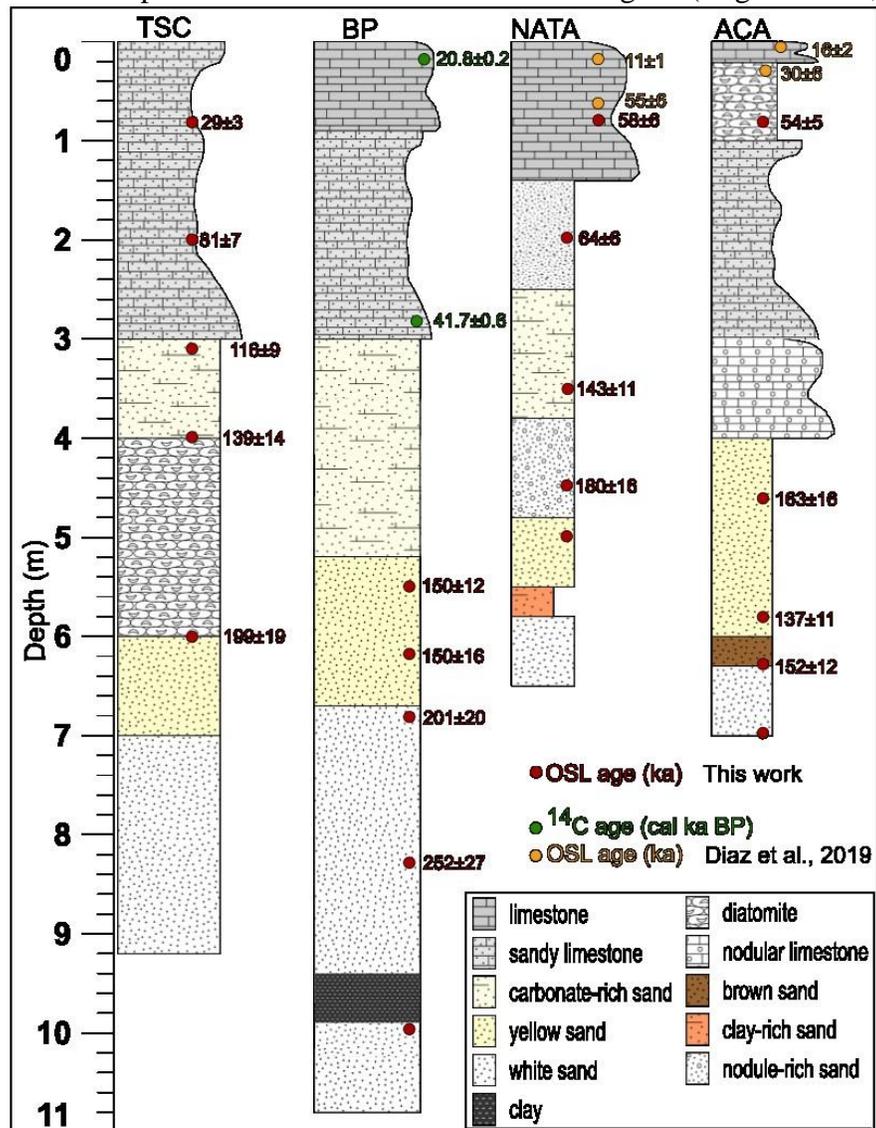
364 5.2 Cosmogenic nuclides

365 Blank corrected concentrations of ^{26}Al and ^{10}Be of sand samples range from 1.18×10^6 to
366 14.03×10^6 , and from 0.33×10^6 to 3.67×10^6 atoms g^{-1} , respectively (Table 2). Although
367 distributed over a noticeable concentration range, $^{26}\text{Al}/^{10}\text{Be}$ ratios are clustered in a narrow
368 spectrum between 3.5 and 4.7, not correlated with nuclides' concentrations, which are not
369 correlated in turn with distance from the CE.

370 Kernel density estimates resulting from the Cosmolian model produce overall log-normal
371 distributions and their weighted average value reflects the most probable surface residence time
372 of each sand sample (Figure 6). While the different combinations of scenarios generally
373 converge into a distinct peak, several samples present a bi-modal distribution with a relatively
374 narrow combined range or positive skewness. These are probably the outcome of grains within
375 the same sample with different sources and transportation histories (Vainer and Ben Dor, 2021),
376 and this variance is reflected in the uncertainty estimation.

377 Residence time estimates for all analyzed samples span the time range between $0.91^{+0.24}/_{-0.22}$ and
378 $2.22^{+0.96}/_{-0.69}$ Myr (Table 2). Their ages display a correlation (with $r = 0.57$) with distance from

379 the CE, with ages being overall younger with proximity to the depocenter on the western margins
 380 of the CE (Figure 6). Simulations that were carried out with an overburden of 1 m (density of 1.7
 381 g cm^{-3}) did not reach convergence with the measured values. This result agrees with the evoked
 382 mean value of the overburden required for simulations to converge with measured
 383 concentrations. This value is 4.7 ± 3.9 m if all samples are considered, or 5.0 ± 3.3 m if the two
 384 thinnest and two thickest simulated values of overburden are excluded (Table 2). Furthermore,
 385 the least likely assigned erosion rate during production in source areas is 3 m Ma^{-1} , in accordance
 386 with major sources in northern provinces where erosion rates are higher (Regard et al., 2016;



387 Garzanti et al., 2022).

388 Figure 5. Columnar sections and chronology of deposits recovered from sampling pits dug in the Chobe Enclave.
 389 Red circles with no age specification indicate water-logged samples where estimated ages are out of stratigraphic
 390 order. The locations of pits are shown in Figure 2b.

391 5.3 Mineralogy

392 All analyzed sands are pure quartzose, with quartz representing 98-100% of the grain
 393 framework, with a few feldspars (almost exclusively K-feldspar) and rare mica. The very poor to

394 extremely poor tHM suite consists mainly of tourmaline, associated with kyanite, zircon, and
 395 staurolite (Table 3). Rutile, epidote, titanite, hornblende, anatase, and brookite, are minor,
 396 sillimanite sporadic, and garnet and apatite are rare. The mineralogical suite significantly differs
 397 (e.g., has much more kyanite; Table 3) from the assemblage that characterizes the regional
 398 Kalahari sand dunes. The ZTR index is 58 ± 10 , and the staurolite/kyanite ratio ranges between
 399 0.4 and 1.2. Simple forward mixing calculations (Garzanti et al., 2012) and similarity analysis
 400 (Vezzoli and Garzanti, 2009) suggest subequal contributions from the Cuando and Uppermost
 401 Zambezi.

402 Table 2. Cosmogenic nuclides concentrations of sand samples from the Okavango Rift Zone (ORZ) and its vicinity
 403 and their simulated results given by the Cosmolian Model. The results shown include (1) the most probable average
 404 overburden (2) the ratio of successful simulations of vertical quartz grain displacement rates evoked from using
 405 fluvial, lacustrine, and palustrine (FLP) datasets for simulating vertical displacement rates vs. those from eolian
 406 datasets, and (3) the surficial residence time that represents the most probable time since the sediment was
 407 introduced into the landscape. Parameters used as input are detailed in Table S4, Supporting Information.

Sample	Current landform	Elevation [m]	$^{26}\text{Al} \times 10^6$ [at g $^{-1}$]	$^{10}\text{Be} \times 10^6$ [at g $^{-1}$]	$^{26}\text{Al}/^{10}\text{Be}$	Simulated overburden height [m]	Residence time [Myr]
PF	River	1016	1.44 \pm 0.18	0.34 \pm 0.01	4.2	7.4	1.27 $^{+0.43}/_{-0.32}$
A2	Dunefield	1127	3.93 \pm 0.18	1.03 \pm 0.04	3.8	4.5	1.43 $^{+0.34}/_{-0.25}$
KN	Dunefield	1064	2.63 \pm 0.14	0.70 \pm 0.02	3.7	5.2	1.53 $^{+0.32}/_{-0.27}$
MB	Dunefield	1072	1.18 \pm 0.15	0.33 \pm 0.02	3.5	7.1	2.22 $^{+0.96}/_{-0.69}$
KZ	River	977	3.14 \pm 0.19	0.68 \pm 0.02	4.6	4.8	0.98 $^{+0.25}/_{-0.26}$
KS20	Dunefield	1250	12.9 \pm 0.58	3.67 \pm 0.11	3.5	2.4	2.17 $^{+1.02}/_{-0.42}$
KS21	Dunefield	1217	3.0 \pm 0.15	0.77 \pm 0.03	3.9	5.6	1.39 $^{+1.48}/_{-0.24}$
KS23	Dune/river bank	1027	5.48 \pm 0.25	1.31 \pm 0.05	4.2	4.1	1.14 $^{+0.39}/_{-0.25}$
KS25	Riverbank	941	2.75 \pm 0.14	0.62 \pm 0.02	4.5	5.4	1.04 $^{+0.24}/_{-0.22}$
KS26	Riverbank	931	4.23 \pm 0.20	0.90 \pm 0.03	4.7	4.2	0.91 $^{+0.24}/_{-0.22}$
KS27	Sandsheet	962	14.03 \pm 0.61	3.28 \pm 0.10	4.3	2.0	1.59 $^{+0.8}/_{-0.37}$
KS28	Sandsheet	1249	4.38 \pm 0.20	1.18 \pm 0.05	3.7	4.3	1.46 $^{+0.35}/_{-0.26}$
KS29	Dunefield	1058	3.80 \pm 0.19	0.93 \pm 0.04	4.1	4.6	1.27 $^{+0.81}/_{-0.24}$

408

409 Table 3. Heavy mineral assemblages. Data of regional fluvial and eolian sand are after Garzanti et al. (2021) and
 410 (2022), respectively.

Sample	HMC w%	tHM w%	zircon	tourmaline	rutile	Ti Oxides	titanite	apatite	epidote	garnet	staurolite	andalusite	kyanite	sillimanite	amphibole	clinopyroxene	others	ZTR	St/Ky
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TSC6.0	0.08	0.05	8	43	7	0	0	0	5	0	20	0	17	0	0	0	0	58	1.2
TSC6.8	0.10	0.06	17	35	4	0	0	0	2	0	19	0	21	1	0	0	0	56	0.9
BP5.5	0.83	0.47	17	32	7	0	0	0	6	0	13	0	24	0	0	0	0	57	0.5
BP10.0	0.22	0.13	12	33	3	2	5	0	4	0	10	0	29	0	0	0	0	48	0.4
ACA6.3	0.13	0.08	14	38	4	0	1	0	4	0	11	0	25	0	2	0	0	55	0.5
ACA7.0	0.19	0.09	17	37	7	0	2	0	5	0	13	0	22	0	1	0	0	60	0.6
NATA3.5	0.19	0.11	17	40	6	0	3	0	2	0	13	0	16	0	0	0	2	63	0.8
NATA5.0	0.21	0.11	18	41	10	0	0	1	6	1	9	0	16	0	1	0	0	68	0.6
Fluvial sand																			
Cuando	0.07	0.03	24	27	6	0	1	0	4	0	19	1	11	0	5	0	0	57	1.8
Linyanti	0.10	0.06	9	51	6	0	0	0	0	0	20	1	13	0	0	0	0	66	1.5
Chobe	0.27	0.14	29	24	6	0	0	0	5	0	10	0	23	0	2	0	0	69	0.4
Zambezi (Kazungula)	0.39	0.25	22	15	6	0	0	0	8	1	7	0	36	1	3	0	0	43	0.2
Zambezi (Livingstone)	0.52	0.23	18	16	5	2	0	1	12	0	4	1	27	0	4	14	0	39	0.1
Zambazi (Sheshka)	0.2	0.1	12	18	14.0	0	1	0.5	9	0	8	1	21	0.5	2	12	0	44	0.4
Zambezi (Kazungo)	0.2	0.1	24	25	8	0.5	0	0	3	0.5	7	0	30	0.5	1	0	0	57	0.2
Kalahari Sand Dunes																			
KS27	0.3	0.3	33	18	7	0	0.5	0.5	28	5	2	0	1	0	1	6	0	58	1.5
KS28	0.5	0.1	47	34	5.5	0	0.5	0.5	0.5	0	10	0	1	0	0	1	0	87	10.5
KS29	0.3	0.1	40	43	5.0	0	0	0	1	0.5	8	0	2	0	0	1	0	88	4.5
A2	0.2	0.2	8	37	1	0	0	0	1	0	43	0	8	0	3	0	0	46	5.4
KS30	0.6	0.3	36	15	6	0	0.5	0	40	1	0.5	0	1	0	0	0	0	57	0.3

411

412 **6. Discussion**

413 6.1 Deposits of the Cuando-Zambezi alluvial fans

414 The studied pits are spread over an area smaller than 5 km² (Figure 2b), and their sedimentary
415 sequences display spatio-temporal variations described below and illustrated in Figure 5.

416 (1) The basal sediments from all four pits comprise regionally continuous white sand,
417 intercalated by muds with a phyllosilicate content ranging between 13-35% (Mokatse et al.,
418 2023). The occurrence of > 70% clay sub-unit (most of which is a mixture of kaolinite and
419 sepiolite) within the white sand at BP (Figure S8, Supporting Information) confirms deposition
420 in a composite environment with markedly changing fluvial energy through time, as typical of
421 alluvial fans (Stock, 2013). The ages of the strata above the white sand constrain the oldest
422 deposits as not later than ~200-150 Ka, as also indicated by two OSL ages of the lower sand unit
423 at BP that were buried at 252 ± 27 and 201 ± 20 Ka.

424 The deepest samples from BP, NATA, and ACA pits were taken from water-logged units using
425 an auger drill and yielded ages out of stratigraphic order. Possible reasons for this age
426 underestimation are discussed in the Supporting Information but this point remains unclear.

427 Therefore, the earliest deposition at the studied sites is constrained to have occurred before 250
428 Ka.

429 (2) Yellow/brown sand overlying the white sand is observed at all sites, displaying some lateral
430 variations and variations in carbonate content. Overall, the clay content in this unit ranges
431 between 19 and 40 %; higher sepiolite abundance at the expense of kaolinite in comparison with
432 the white sand below may imply some evaporitic conditions (Mokatse et al., 2023). The yellow
433 sand facies was deposited between ~200 and 140 ka (199 ± 19 and 137 ± 11 ka; $n = 5$). At the
434 ACA site, the yellow sand is different in nature than in the other sites as it is noticeably rich in
435 carbonate and iron. This could indicate a reworked paleosol that may explain the age inversion
436 observed at ACA (even though ages overlap within their analytical uncertainty), which could be
437 also explained by bioturbation. At the NATA site, highly siliceous, bioturbated deposits
438 accumulated at 180 ± 16 Ka. Between ~200 and 139 ± 14 Ka, a diatomite unit accumulated at
439 TSC. These observations point to multiple depositional environments with a relatively large
440 range of water depths and depositional energies, composition of solutes, and precipitation-to-
441 evaporation ratios. These sub-environments were found in close proximity inside a dynamic
442 alluvial fan setting.

443 (3) The diatomite and carbonate deposits that lie in unconformity above the sand below (Figure
444 S8, Supporting Information) mark the initial deposition in a lacustrine environment that took
445 place during the regionally wet MIS 5 (Burrough et al., 2009). Their deposition is constrained by
446 three samples from two sites to have occurred after ~200 ka, with depositional ages of 143 ± 11
447 and 116 ± 9 Ka. Lateral and vertical calcite content changes (Mokatse et al., 2023), ranging from
448 0 to 4% at the NATA and TSC sites (where ages were determined), to ~50% at BP and ACA
449 where age is defined only by correlation. The change from primary siliceous deposits that
450 contain no carbonate to the deposition of carbonate implies a noticeable change in the chemistry
451 of the precipitating solution that could have resulted from an adjustment to morphotectonic or
452 climatic shifts, as discussed below.

453 (4) A change in the environment occurred at 81 ± 7 ka and is synchronous with the global
454 climatic perturbations and regional environmental changes of MIS 3 (e.g. Agostaand and
455 Compagnucci, 2016; Stewart and Jones, 2016). A carbonate-rich palustrine/lacustrine
456 environment is inferred due to the abundance of calcite at all sites, commonly representing the
457 most abundant mineral (Mokatse et al., 2023). The upper units at the NATA and ACA sites,
458 which lie ~0.5 km from each other, are constrained by six OSL ages ranging from 58 ± 6 and 11
459 ± 1 ka. While carbonate is the main precipitate at NATA around 50 ka, diatomites and clays
460 (with high sepiolite content) were deposited at ACA, pointing to less alkaline conditions locally,
461 possibly related to pluvial lake settings. These sediments resemble surficial deposits of the
462 Okavango Delta that originate from semi-continuous flood events under semi-arid conditions and
463 desiccation. In the Okavango Delta, silicious and carbonate-rich precipitates are discretely
464 deposited, and while carbonate minerals are present, they are far less common in the Okavango
465 Delta than in the CE (McCarthy and Ellery, 1995; Ringrose et al., 2008; Dauteuil et al., 2021).
466 These differences raise the question of the composition and origin of the parent solutions of the
467 water flows in the CE during the later Pleistocene.

468

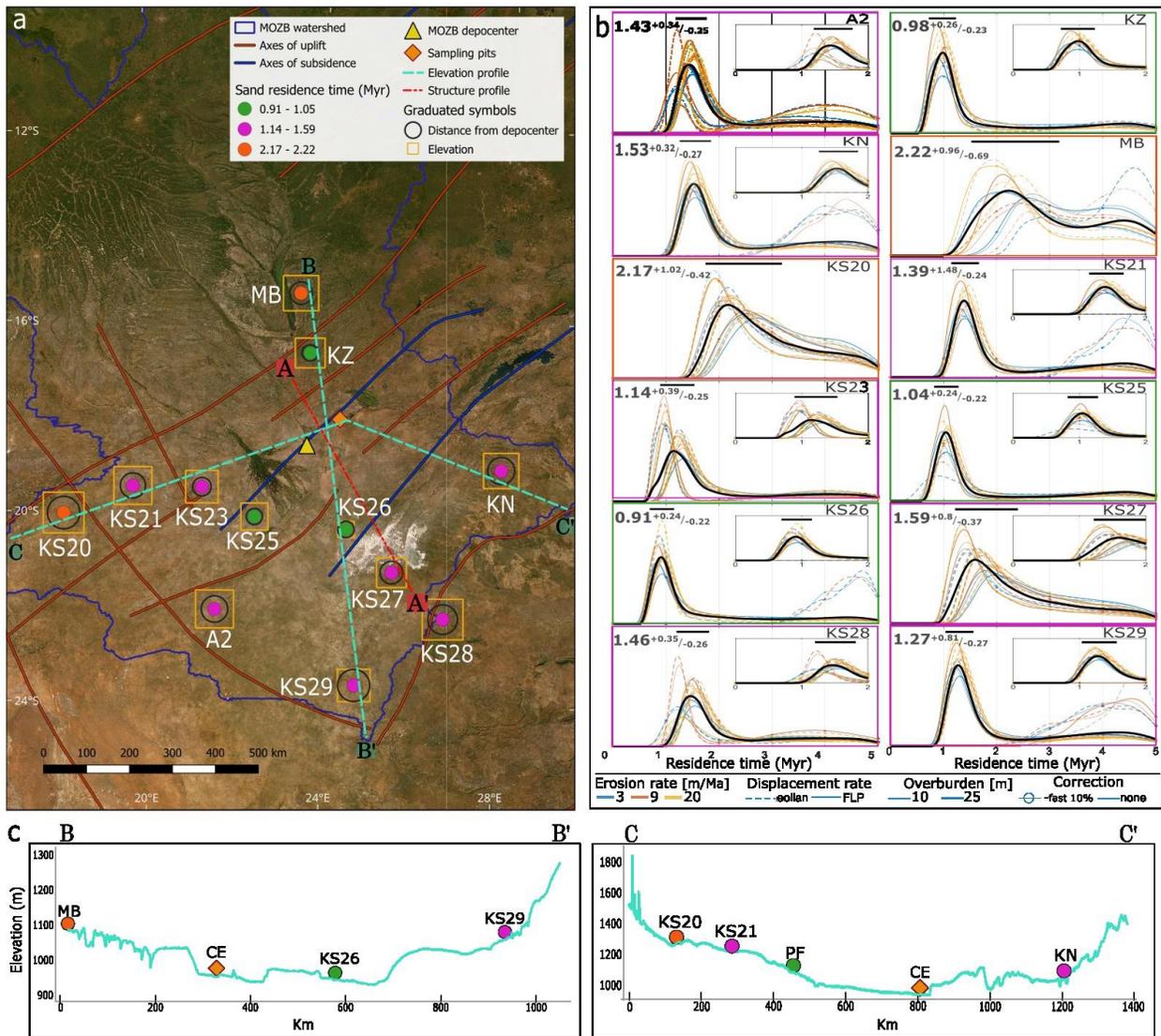
469 6.2 Provenance

470 The mineralogical assemblages of all buried CE samples imply similar sources, represented by a
471 mixture of sediments presently carried by the Zambezi and Cuando rivers that drain northern
472 terrains (Garzanti et al., 2021). This could be the result of the inter-basin hydrological
473 connectivity with the Zambezi River that changes naturally as drainages are separated or
474 combined through avulsion and due to external forces, such as climate change and tectonic
475 activity (Shaw and Thomas, 1992; Wilkinson et al., 2023). Furthermore, XRD patterns of the
476 studied samples reveal that non-carbonate mud samples (Figure S8, Supporting Information)
477 contain 14-46 % phyllosilicates (Mokatse et al., 2023), congruent with a primary fluvial/alluvial
478 transporting agent. Moreover, kyanite enrichment in sediments carried by the Chobe River (the
479 spill of the Cuando into the CE) across the CE, relative to the upper reaches of the Cuando,
480 points to the incorporation of fluvial sediments from the Upper Zambezi by the Chobe, and their
481 reworking from deposition in alluvial fan settings (Garzanti et al., 2022; Mokatse et al., 2022b).
482 A northern source is also suggested for the surficial sand that is carried by rivers into the MOZB
483 as it presents a significantly higher success rate of Cosmolian simulations by applying
484 displacement rates constructed from the FLP rather than the eolian dataset rates (Table 2). This
485 sand (samples KS25, KS26) has $\leq 2\%$ success in Cosmolian convergence events for scenarios
486 with an erosion rate of 3 m Ma^{-1} that characterizes southern source areas, while higher erosion
487 rates that characterize northern areas yield higher successful scenarios. This pattern of
488 simulations resembles the simulations of northern dune sand and river samples (MB, KZ, PF,
489 KS21) that arrive from areas with higher erosion rates (Garzanti et al., 2022) and differ from the
490 rest of the sand samples to the south that present noticeable convergence also for scenarios with
491 slower eroding source areas (Figure 6). Furthermore, a coupled fluvial-eolian transport agent is
492 deduced also for currently eolian dune samples (Garzanti et al., 2022), as all modelled samples
493 experienced successful simulations by applying rates from both FLP and eolian datasets.

494 Although some mineralogical similarity exists with sand dunes located on the upper reaches of
495 the Zambezi, the mixed source for the buried CE sediments differs from sources that
496 predominate the Kalahari sand dunes as well as from their diagenetic history (Table 3). Their
497 differences in nature and age are reflected in their colors. The sand in the CE is mostly white and
498 yellow (Figure 5; Supporting Information S8) resulting from secondary iron oxyhydroxides
499 coating, likely due to hydration under alkaline conditions during fluvial transportation.
500 Conversely, the eolian Kalahari Sand is red (Wang et al., 2007) due to longer pedogenesis with
501 rubificating edaphic conditions (Walker and McKee, 1979). Hence, whereas the modes of the
502 grain size distribution of CE sands and Kalahari sand lie within the same range (Mokatse et al.,
503 2022a), they do not share a genetic link and do not represent re-deposition of Kalahari dunes.
504 Additional observations point to fluvial incision and transport of material from elsewhere that
505 postdates the establishment of the dunes. These observations include (1) the offset and truncation
506 of dunes west of the Okavango Delta by faults and the lowering of base level associated with the
507 subsidence of the MOZB (McFarlane and Eckardt, 2007); and (2) the flow of the Gwayi River
508 parallel to the crests of linear dunes in the eastern MOZB (Figure 2a) (Thomas and Shaw, 1991;
509 Moore et al., 2012). Thus, the reason for the different mineralogical and textural signatures
510 between the Kalahari Sand and the CE buried sediments is probably the subsidence of the CE
511 and the incision of rivers into it (after the fixation of the eolian sand), carrying sediment from
512 their headwaters in a significantly greater proportion than recycled eolian sand from their
513 riverbanks.

514 6.3 Sand chronology

515 Eolian sand within the MOZB was exhumed between $2.22^{+0.96}_{-0.69}$ and $1.14^{+0.39}_{-0.25}$ Ma,
 516 marking the upper age limit for the last significant geomorphologically evident subsidence event
 517 in the CE, as no sand with similar sources and diagenetic history is found in the CE. The
 518 distribution of the mean residence times of all sand samples correlates moderately ($r = 0.6$) with
 519 elevation and increases with distance from the Linyanti-Chobe Basin ($r = 0.57$; Figure 6).
 520 Collectively, this points to the preservation of the older sediments on the higher margins of the
 521 tectonic trough of the CE and the incorporation of more recently eroded material downwards into
 522 the evolving basin.



523
 524 Figure 6. (a) Sands of the Makgadikgadi–Okavango–Zambezi Basin are categorized into three groups based on their
 525 sedimentary residence time. Simplified structural axes highlight the multi-block configuration (Haddon and
 526 McCarthy, 2005) overlying a satellite image of central southern Africa (ESRI, 2023). The size of the graduated
 527 symbols increases with larger values. (b) Kernel density estimates of the sedimentary residence time of sands,
 528 simulated with the Cosmolian Model (Vainer et al., 2018; Vainer and Ben Dor, 2021). The probability plots show
 529 successful runs in which simulated concentrations of ²⁶Al and ¹⁰Be simultaneously matched with their analytical

530 values. 10,000 iterations were applied for each combination of the boundary conditions. The various scenarios
531 include three values of erosion rate at the source areas, vertical displacement rates constructed from accumulation
532 ages of either eolian or Fluvial-Lacustrine-Pluvial (FLP) datasets, and three values representing different
533 transportation agents that resolve in changeable overburden thickness. The weighted average of the matching
534 simulations is shown with a black solid line with uncertainty marked by the horizontal line on top, calculated with
535 full-width at the half-maximum approach. The simulations performed for the PF sample are not shown due to a <1%
536 success rate. (c) Elevation profiles of two nearly perpendicular cross sections across the basin, passing through the
537 Chobe Enclave, constructed from SRTM 30 sec DEM (Farr et al., 2007).

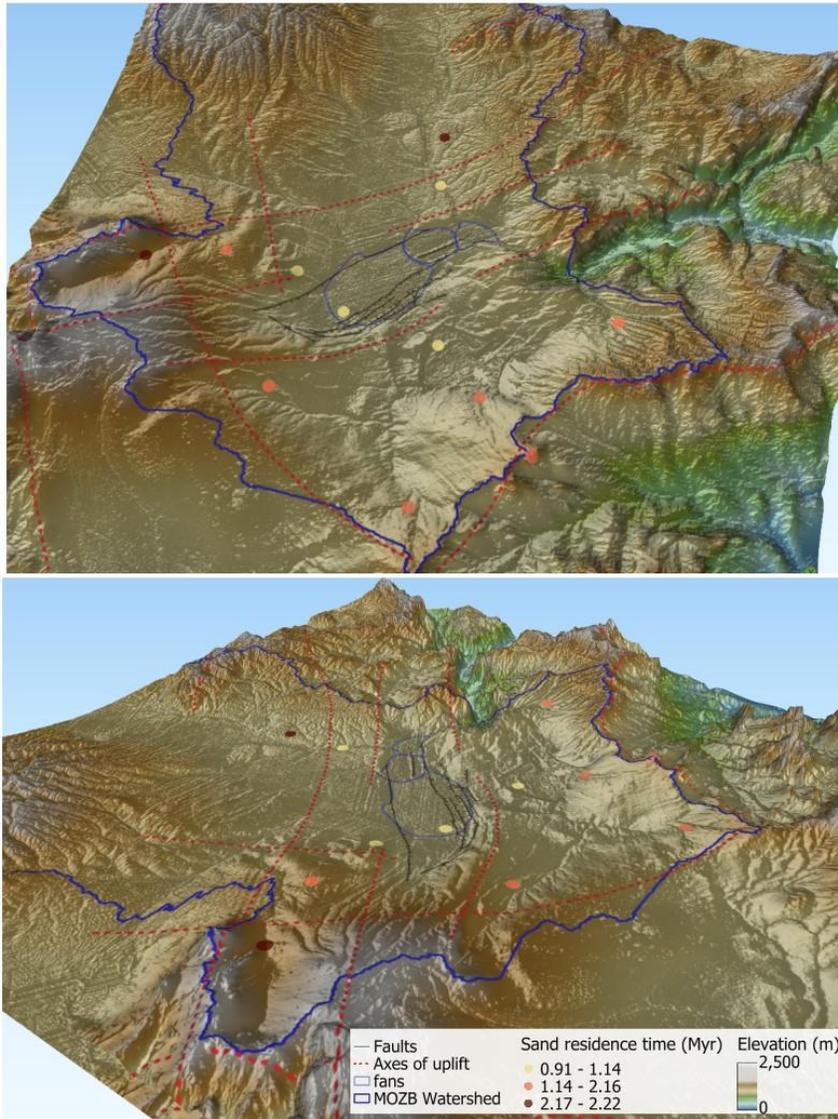
538 The Jenks natural breaks optimization highlights three periods of sand introduction that also
539 roughly correspond to their structural position with respect to the CE and their geomorphological
540 context (fluvial/eolian) at present (Figure 6). (1) Sand collected in fluvial settings near the CE
541 depocenter belongs to the youngest age group with mean ages in the range of 1.05-0.91 Ma ($n =$
542 3). This excludes additional input from the Okavango River (PF) that yielded less than 1% of
543 successful simulations with a mean age of $1.27^{+0.43}_{-0.32}$ Ma. In the subsurface of the western
544 MOZB, sediments younger than ~ 1.1 and 1.4 Ma are not preserved in the downthrown and
545 upthrown blocks, respectively (Vainer et al., 2021). The absence of buried deposits younger than
546 ~ 1 Ma and the lack of sand production since that time suggest a re-organization affecting the
547 interconnection between the fluvial and eolian systems around 1 Ma. The paucity in sediment
548 burial after ~ 1 Ma is observed throughout the southern Kalahari, suggesting the beginning of a
549 primary regional eolian phase (excluding the ORZ), following tectonic uplift of the Kalahari
550 margins (Matmon et al., 2015; Vainer et al., 2018b). (2) Eolian sand located on the surface that is
551 just above the CE yielded mean residence time ages of 1.59-1.14 Ma ($n = 7$). During this period,
552 sand from eolian landforms located ~ 50 -300 km to the south and southwest of the MOZB water
553 divide was extensively formed (Vainer et al., 2022), pointing to a regional (over the MOZB
554 limits) phase of sand production. (3) The most distal to the CE eolian sand was exhumed around
555 2.2 Ma ($n=2$; Figure 6). The initial sand supply into the MOZB coincides with the deposition of
556 basal eolian sand characterized by eolian grain size distribution in the southwestern Kalahari
557 between $2.2^{+0.18}_{-0.17}$ and $1.74^{+0.15}_{-0.15}$ Ma (Vainer et al., 2022) and with distinct hydrological
558 changes in the western Kalahari at ~ 2 Ma (Miller et al., 2010). Together, this chronology points
559 to Kalahari Basin-scale changes that resulted in the initiation of sand cover and its eolian
560 distribution around 2.2 Ma.

561 6.4 Landscape evolution of the Okavango Rift Zone since the Pleistocene

562 The chronology of landscape evolution in the ORZ is addressed via two dating methods that
563 differ by an order of magnitude in their dating capabilities, allowing temporally constraining the
564 rifting before and after the last significant phase of subsidence in the CE. The chrono-structural
565 development of the ORZ can be tracked through the relationships between sedimentation and
566 geomorphology, hinting at the stages of morphodynamical evolution of the nascent rifting zone
567 (Figures 6, 7). Two elevation profiles that pass through the CE illustrate a symmetric (N-S, E-W)
568 structural-block development during continental rifting (Holz et al., 2017), with the oldest
569 sediments deposited at ~ 2.2 Ma. The two sites, where sand of this age is present, are located on
570 elevated landforms on the outer-most structural blocks with respect to the CE (Figure 7). These
571 sands could have been generated due to erosion following relief forming in the MOZB by virtue
572 of tectonics at ~ 2.5 Ma (Thomas and Shaw, 1990; McCarthy et al., 2002; Moore et al., 2012;
573 Vainer et al., 2021). A more recent tectonic activity resulted in the formation of a lower base
574 level, enabling the preservation of older sand on the surface of the elevated landforms.

575 The inner lower blocks that lie above the CE accommodate sand that was formed at ~ 1.6 - 1.1 Ma,
576 representing a second phase in landscape lowering and deformation. Tectonism at ~ 1.4 Ma was
577 biochronologically inferred from the lacustrine radiation of tigerfish in the MOZB (Goodier et al.,
578 2011) and was claimed by Moore et al. (2012) to cause changes in the configuration of the
579 MOZB hydrological system. Such a change is also observed in the chrono-stratigraphy of the
580 sand, as no sand that is found today in eolian settings has been produced since. Therefore, this
581 timing signifies the earliest date for subsidence and formation of accommodation space in the

582 CE. Finally, the successful modelling efforts of the fluvial sand indicate that it was exhumed
 583 around ~ 1 Ma, pointing to a change in the depositional environments that possibly resulted from
 584 a new structural configuration occurring around the same time (Matmon et al., 2015; Vainer et
 585 al., 2018b), concurrent with the Middle Pleistocene Transition (MPT, 1.2-0.75 Ma; Herbert,
 586 2023).



587
 588 Figure 7. Sand residence time and structural/geomorphological elements on top of a 3D elevation model of the
 589 Makgadikgadi–Okavango–Zambezi Basin (purple outline), constructed from SRTM version 3.0 Global 1 arc second
 590 DEM and 30 sec DEM (Farr et al., 2007) (not to scale). (a) looking northward. (b) Looking eastward.

591 Burial ages of basal deposits in the CE could indicate the timing of the last significant rifting
 592 stage. However, such deposits were not reached in this work, and thus direct dating of the
 593 earliest sedimentation in the CE could not be achieved. We could only determine that
 594 accommodation space in the CE was available after 1.1 Ma and before 0.25 Ma. Sedimentary
 595 sequences deposited during continental rifting may overlie volcanic or basement rocks and
 596 typically consist of gravel, followed, or intercalated by fluvio-deltaic sands, overlain in turn by

597 lacustrine and evaporitic deposits (Olsen et al., 1996; Young et al., 2000; Nielsen et al., 2007).
598 The syn-rift sequence described in this study lacks basal conglomerate and begins with alluvial
599 sand. In accordance with the isopach map of Haddon and McCarthy (2005; Figure 3), the
600 unreached sedimentary suite below the studied pits in the CE is possibly 20-50 m thick.
601 Assuming similar accumulation rates to those of the dated sediments, and accounting for the
602 estimated missing thickness range, it is speculated that sedimentation in the CE may have started
603 closer to 1 Ma than to 0.25 Ma.

604 Several observations point to the existence of a topographic depression since ~1 Ma in the
605 MOZB, where waterbodies were sustained and linked to tectonic-induced landscape evolution
606 (Grove, 1969; Moore et al., 2012). (1) Phylogeographic records of catfishes point to their
607 radiation in a lacustrine environment at 0.9 ± 0.5 Ma. (Day et al., 2009; Cotterill and De Wit,
608 2011) (2) Early Stone Age (ESA) artifacts with a minimum age of 0.5 Ma were found in paleo
609 lacustrine settings (McFarlane and Segadika, 2001; McFarlane and Eckardt, 2006; Moore and
610 Cotterill, 2007) (3) Gravels containing ESA artifacts were found <10 km downstream the
611 Victoria Falls (Figure 2), indicating the initiation of gorge incision due to lacustrine overtop
612 from the MOZB into the Zambezi River at 1.1-0.65 Ma (Clark, 1950; Moore and Cotterill,
613 2010). The existence of this waterbody (or waterbodies) in the CE cannot be determined with the
614 findings of this study.

615 The earliest dated waterbody deposits in the CE are diatomite and carbonate which accumulated
616 at ~140 ka (Figure 5). This waterbody could have extended some 300 km to the southwest,
617 where partially cemented lakebed deposits and coarse sand interpreted to represent a beach ridge
618 accumulated at 133 ± 12 and 140 ± 11 ka, respectively (Shaw et al., 2003; Burrough et al., 2007).
619 These sediments were interpreted to have formed under wet environmental conditions, as their
620 deposition is synchronous with cave deposits on the western elevated margins of the MOZB
621 (Brook et al., 1998; Burrough et al., 2007). This agrees with our observations of synchronous
622 deposition of silica nodule-rich sands, diatomites, and carbonate-rich sands, reflecting stability in
623 water flux. However, structural displacement along the northern MOZB flanks was also
624 proposed to occur between 300 and 100 Ka, based on the preservation of archaeological artifacts
625 of this age on paleo-Makgadikgadi lakebeds. This was interpreted to cause the deflection of the
626 Cuando River from the Makgadikgadi Basin into the CE, forming a waterbody (Moore and
627 Larkin, 2001; Moore et al., 2012). Furthermore, drying conditions that commenced at 110 ka
628 following a wet period were inferred based on thermoluminescence dating and geochemical
629 study of duricrusts in the Makgadikgadi Basin (Ringrose et al., 2005; Ringrose et al., 2009).
630 Accordingly, a transition from a clast-dominated to chemical-dominated accumulation took place
631 in the CE between 116 ± 9 and 81 ± 7 ka.

632 The last phase of carbonate precipitation and diatomite deposition in the CE documents an
633 enduring waterbody that existed between 54 ± 5 and 11 ± 1 ka (Table 1, Figure 5). Lacustrine
634 deposits from this period, centered at ~40 ka, were reported from all MOZ basins (summarized
635 in Burrough et al., 2009), hinting at a vast expansion of the lacustrine/palustrine system.
636 Holocene sediments have not been observed in the studied sections, and their absence is
637 consistent with the climatically driven desiccation of a waterbody in the Makgadikgadi Basin in
638 the early Holocene (Partridge et al., 1997; Burrough et al., 2009), possibly resulting in their
639 erosion. Alternatively, such deposits could have also been removed due to the activation of faults

640 that occurred at ~6 ka in the CE, causing the diversion of drainage networks and inverted relief
641 (Mokatse et al., 2022).

642

643 **7. Conclusions**

644 A combination of a mineralogical provenance study, optically stimulated luminescence (OSL)
645 dating of alluvial and lacustrine deposits, and cosmogenic nuclide-based estimation of sand
646 residence time was applied to chronologically constrain the landscape evolution in the Chobe
647 Enclave, a tectonically active sector of the Okavango Rift Zone. The Chobe Enclave adjoins the
648 thickest depocenter in the Makgadikgadi–Okavango–Zambezi Basin, which experienced
649 significant down-warping at ~2.5 Ma. Cosmogenic nuclide-based modelling indicates that sand
650 that was formed following this event is preserved on the elevated margins of the Makgadikgadi–
651 Okavango–Zambezi Basin. Model results further suggest an additional event of landscape
652 lowering occurring around 1.5 Ma, which probably corresponds to regional tectonism as most of
653 the eolian Kalahari Sand was formed around this time and has been recycled since then in the
654 semi-endorheic Kalahari Basin. This stage was followed by the accommodation of waterbodies
655 within the Makgadikgadi–Okavango–Zambezi Basin, where their deposits of upper Pleistocene
656 age are preserved. The last estimated episode of sand formation at 1.1 Ma marks the older limit
657 for localized rifting in the Chobe Enclave, which probably occurred during the Middle
658 Pleistocene Transition that took place between 1.2 and 0.75 Ma (Herbert, 2023).

659 Alluvial fans and waterbodies evolved within the depressed landscape of the Chobe Enclave, and
660 their mineralogical signature suggests supply from both Zambezi and Cuando rivers, influenced
661 by hydrological connectivity, climate, and tectonic activity. This alluvial system carried sand of
662 different origin and diagenetic history than the older eolian sand that is structurally placed above
663 the Chobe Enclave. Dating the alluvial sediments that were deposited in the evolved rift via OSL
664 provided a younger time constraint for the incision. The earliest documented sedimentation in
665 alluvial fan settings is dated as 252 ± 27 ka, representing the youngest age limit for a rifting
666 episode in the Chobe Enclave. Finally, the subsidizing trough of the Chobe Enclave hosted
667 waterbodies for at least ~140 ka, which were possibly connected with other waterbodies within
668 the Makgadikgadi–Okavango–Zambezi Basin.

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678

679 **Open Research**

680 Unless mentioned otherwise, the data presented in this paper is original. Mixing and similarity
681 analyses based on mineralogical assemblages were performed after Garzanti et al. (2012) and
682 Vezzoli and Garzanti (2009), respectively. OSL modelling for estimating dose rate and
683 depositional age was performed via Mauz and Hoffmann (2014) Kreutzer et al. (2012, 2019,
684 2022), and Duller (2015). Cosmogenic nuclides modelling for residence time estimation was
685 done using Cosmolian (Vainer and Ben Dor, 2021).

686

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