

1 **Characterizing the depth of cover across South**  
2 **Australia: A simple passive-seismic method for**  
3 **estimating sedimentary thickness**

**Shubham Agrawal<sup>1\*</sup>, Caroline M. Eakin<sup>1</sup>, and John O'Donnell<sup>2</sup>**

<sup>1</sup> Research School of Earth Sciences, Australian National University, Canberra, Australia.

<sup>2</sup> Geological Survey of South Australia, Department for Energy and Mining, Adelaide,  
Australia.

4 Email: shubham.agrawal@anu.edu.au (Shubham Agrawal); caroline.eakin@anu.edu.au  
(Caroline Eakin); John.O'Donnell2@sa.gov.au (John O'Donnell)

*This manuscript is a preprint and has been submitted for publication in Bulletin of the  
Seismological Society of America. Please note that, the manuscript status is under peer-review,  
the manuscript has yet to be formally accepted for publication.*

---

\*Corresponding author: Shubham Agrawal (shubham.agrawal@anu.edu.au)

**Abstract**

5  
6 A blanket of sedimentary and regolith material covers approximately three-quarters of the  
7 Australian continent. This poses a significant exploration challenge, with future mineral and  
8 energy resources discoveries likely confined beneath the sedimentary cover. The most fun-  
9 damental question that can be asked is how thick are the sediments? Borehole drilling and  
10 active seismic experiments provide excellent constraints, but they are limited in geographical  
11 coverage due to their expense, especially when operating in remote areas. On the other hand,  
12 passive-seismic deployments are relatively low-cost and portable, providing a practical alter-  
13 native for initial surveys. Here we introduce a technique utilizing receiver functions for both  
14 temporary and permanent seismic stations in South Australia. We present a straightforward  
15 method to determine the basement depth based on the arrival time of the P-converted-to-S  
16 phase generated at the boundary between crustal basement and sediments. Our results provide  
17 an excellent match with the available borehole data, allowing for a simple predictive relation-  
18 ship between Ps arrival time and basement depth to be established. Our method thus opens  
19 a way to determine the basement depth in unexplored areas requiring only temporary seismic  
20 stations deployed for < 6 months.

21  
22 **Key words:**

- 23 1. Characterizing the depth of sedimentary cover is critically important for future exploration;  
24 however current methods can be prohibitively expensive.
- 25 2. We devise a receiver function based method to quickly and cheaply determine the basement  
26 depth employing temporary seismic stations.
- 27 3. We demonstrate the applicability of our method to South Australia for sedimentary basin thick-  
28 ness ranging up to 2600 m.

## 29 **Introduction**

30 The Australian continent is the flattest and one of the most tectonically stable continents on Earth,  
31 with the last major mountain-building event occurring  $\sim 250$  Ma (Pain et al., 2012). As a result,  
32 around  $\sim 80\%$  of the surface of Australia is masked by sediments, predominately Phanerozoic in  
33 age, obscuring the underlying crustal basement. Onshore sedimentary basins vary in thickness,  
34 ranging from a few hundred meters to up to 15,000 m thick in the Fitzroy Trough within the  
35 Canning Basin, northwest Australia (Yeates et al., 1984). Many of these basins host rich natural  
36 resources, such as base metals (and their subsidiaries), hydrocarbons, and groundwater (e.g. Hitz-  
37 man et al., 2010; Leach et al., 2010). For instance, significant oil and gas reserves can be found in  
38 the Cooper and Eromanga Basins of central Australia. Recent studies focusing on Australia have  
39 suggested that 15% of Australia is a prospective target for sediment-hosted deposits (Hoggard  
40 et al., 2020). The abundance of Archean and Proterozoic blocks in Australia provides a potentially  
41 fruitful avenue for future explorations as sediment-hosted metal minerals systems are largely as-  
42 sociated with Proterozoic basins (e.g. Hitzman et al., 2010; Leach et al., 2010). A known example  
43 would be the Carpentaria Zinc Belt within the North Australian Basin System which contains three  
44 of the ten largest zinc-lead deposits globally (e.g. Southgate et al., 2013).

45

46 Australia's economy is heavily supported by natural resources and mining activity; however the  
47 majority of the easily accessible near-surface resources have already been found and, to a large de-  
48 gree, exploited (Mudd et al., 2019; Mudd, 2007). As such, the UNCOVER Initiative has become a  
49 national priority (Collett and McFadden, 2014), with the aim to improve the discovery potential of  
50 new mineral deposits currently hidden beneath cover and thus keep pace with increasing global de-  
51 mand (Ali et al., 2017). Constraining the sediment thickness (i.e., depth of cover) is a fundamental  
52 component within this UNCOVER Initiative, under the core theme of "characterizing Australia's  
53 cover".

54

55 The most accurate estimates of sediment thickness come from borehole drilling followed by  
56 active-source seismic surveys. While both provide excellent constraints, they are expensive (es-  
57 pecially for deep boreholes) and require heavy machinery, which can be logistically difficult to  
58 dispatch to remote locations. As an alternative, passive-source seismology is non-invasive, rel-

59 atively low-cost, and with portable instrumentation (i.e. seismometers) that is easier to transport  
60 and deploy to remote locations (Rawlinson et al., 2017). Several passive-source seismic techniques  
61 currently exist to study the near-surface structure, such as the horizontal to vertical spectral ratio  
62 (commonly referred to as HVSR), spatial autocorrelation coefficient method, and using the am-  
63 bient seismic wavefield (e.g. Lunedei and Albarello, 2010; Okada, 2006). Here we demonstrate  
64 the feasibility of using passive recording by single-station seismometers to infer the depth to crys-  
65 talline basement directly. We use the arrival information of the P-to-S converted phase from the  
66 top of crustal basement. Utilizing a recent expanse in passive seismic deployments across South  
67 Australia (supplement Figure S1), and a wealth of existing drilling data in regions of exploration,  
68 we show a remarkable relationship between P-to-S arrival time and the depth of basement. This  
69 relationship is calibrated from the existing data and applied to predict the sediment thickness in  
70 under-explored remote regions where recent seismic deployments have occurred. The method we  
71 employ is simple and straightforward to use, typically requiring only a few months of recorded  
72 data and is dependent only on the thickness of the sediments and their Poisson's ratio.

73 [Figure 1 about here.]

#### 74 **Sediment distribution and history of South Australia**

75 South Australia has a rich geological history that stretches as far back as the Archean, with the  
76 origin of the Gawler Craton, the largest crustal province in South Australia (Hand et al., 2007).  
77 Within the craton lies Olympic Dam, the world's largest uranium resource, with significant copper  
78 and gold reserves as well (e.g. Reid, 2019). Much of the mineral rich Gawler Craton, and indeed  
79 South Australia in general, is obscured by sediments, with around a dozen or more major sedimen-  
80 tary basins across the state (Figures 1 & 2). Sediment thickness within these basins varies from  
81 0 to > 3,500 m, owing to a lengthy history of spatially varying sedimentation (Wopfner et al., 1969).

82  
83 The oldest basin within South Australia is the Mesoproterozoic Cariewerloo Basin, with a min-  
84 imum depositional age of  $1,424 \pm 51$  Ma (Fanning et al., 1983). This intracratonic basin holds  
85 dominantly fluvial sandstones with thicknesses up to 1,500 m and holds potential for uranium min-  
86 eralization (Beyer et al., 2018). The second oldest is the Officer Basin, a large intra-cratonic basin  
87 of Neoproterozoic-Devonian age that extends west into Western Australia. Within South Australia,

88 the basement depth exceeds 3,000 m (Wopfner et al., 1969). The Officer Basin is poorly explored,  
89 with limited boreholes and no commercial discoveries. The Warburton Basin is the oldest Paleo-  
90 zoic basin in South Australia (Cambrian-Ordovician age) and overlies the Proterozoic basement in  
91 the north-east part of the state (Gatehouse, 1986). The basin is demarcated into west and east by  
92 uplifted basement under the Birdsville Track Ridge, with the east Warburton studied more exten-  
93 sively as it underlies the oil and gas producing Cooper and Eromanga Basins (Radke, 2009). The  
94 Early to Middle Cambrian Arrowie Basin comprises three separated areas from Flinders Ranges to  
95 the border with New South Wales (Zang et al., 2004). The thickest sediments are in the central part  
96 with values greater than 3,000 m. The early Paleozoic Pedirka and Arckaringa Basins are largely  
97 under-explored, with Permian sediment thickness reaching 1,000 m in certain areas (Drexel and  
98 Preiss, 1995). The Cooper Basin, comprising Middle Triassic to Permian non-marine sequences,  
99 is perhaps the best-studied basin in South Australia owing to prolific oil and gas reserves. The  
100 basin's long axis trends northeast-southwest and the predominantly gas reserves are located 1,250  
101 m below the surface (Radke, 2009). In contrast, the overlying Jurassic-Cretaceous Eromanga Basin  
102 hosts mostly oilfields (Radke, 2009). The Eromanga Basin is more extensive, fully covering the  
103 Cooper and Warburton Basins, extending to the western boundaries of Pedirka and Arckaringa  
104 Basins, and further north and east into central Queensland and New South Wales (Wopfner et al.,  
105 1969). The Mesozoic Bight and Otway Basins have significant offshore sediments and potential for  
106 explorations (Drexel and Preiss, 1995). Otway, in particular, has notable onshore sediments with  
107 thickness of more than 3,000 m, as detected by borehole measurements (Figure 2). The Mesozoic  
108 Berri Basin is completely concealed beneath the Murray Basin with basement depths up to 600 m  
109 (Brown and Stephenson, 1991). Most recently, during the Cenozoic, three large-scale sedimenta-  
110 tion episodes occurred, namely Eucla (extends 500 km offshore), Murray, and Lake Eyre Basin  
111 (Brown and Stephenson, 1991; Drexel and Preiss, 1995).

112

113 Given the abundance of sedimentary basins and associated sub-surface natural resources in  
114 South Australia, there have been more than 27,100 onshore boreholes drilled (see Data and Re-  
115 sources Section) that determined the depth to basement (Figure 2). The basement depth values  
116 range between 0-3698 m, with the deepest values in the Cooper-Eromanga Basin area around  
117 Moomba. The vast majority ( $\sim 94\%$ ) of boreholes have been drilled in places with sediment thick-

118 ness less than 200 m (supplement Figure **S2**). The highest concentration of deeper boreholes (>  
119 1400 m) is within the Cooper-Eromanga Basin, with more than 900 boreholes drilled to the base-  
120 ment.

121 [Figure 2 about here.]

## 122 **Data and Methods**

### 123 **Receiver functions in sediments**

124 Receiver functions (RFs) are an effective and commonly used tool to image seismic discontinu-  
125 ities within the Earth's crust and mantle arising from changes in the material properties. The RF  
126 technique focuses on converted phases - typically P to S conversions (Ps) in the P-wave coda -  
127 arising from seismic velocity discontinuities below the receiver (i.e. a seismic station)(Ammon,  
128 1991; Langston, 1979; Vinnik, 1977). To remove the effect of source-side structure and wavefield  
129 propagation, the vertical component (dominated by P-wave energy) is deconvolved from the radial  
130 component seismogram (containing energy from both P and S waves) to obtain the radial RF, here-  
131 after simply referred to as the RF (e.g. Ligorría and Ammon, 1999).

132

133 RFs have been employed extensively to image the Moho (crust-mantle boundary), the lithosphere-  
134 asthenosphere boundary, and the 410 km and the 660 km global seismic discontinuities. (e.g.  
135 Birkey et al., 2021; Kennett et al., 2011; Tauzin et al., 2013). However, only a few have used RFs  
136 to constrain the properties of the sedimentary layers without forward modeling (e.g. Cunningham  
137 and Lekic, 2020; Piana Agostinetti et al., 2018; Srinivas et al., 2013; Yeck et al., 2013; Yu et al.,  
138 2015; Zheng et al., 2005). In the presence of a sedimentary layer, RFs are dominated by strong re-  
139 verberations after the arrival of the direct-P phase. These can last for several seconds and can mask  
140 the signal from deeper crustal discontinuities (e.g. Tao et al., 2014; Zelt and Ellis, 1999). This oc-  
141 curs due to the significant impedance contrast between the low-velocity sedimentary layer and the  
142 basement rock, trapping the P and S waves in the sediment layer. The presence of a sedimentary  
143 layer has some additional customary effects on the RF signal. When the incoming teleseismic P-  
144 wave enters the low-velocity sediments, the angle of incidence steepens (due to Snell's Law), and  
145 the ray becomes nearly vertical (Figure 3). The energy of the direct-P wave is, therefore mostly

146 confined to the vertical component, with minimal energy on the horizontal components (radial  
147 and transverse). The amplitude of the first P-arrival on the radial RF, therefore, decreases as the  
148 incidence angle steepens and the pulse width is broadened (Sheehan et al., 1995; Zelt and Ellis,  
149 1999). Instead, the P-to-S converted phase between the basement and sediments often becomes  
150 the strongest amplitude signal on the radial RF, and can completely mask the low amplitude direct  
151 P-wave (Cunningham and Lekic, 2020; Yeck et al., 2013; Yu et al., 2015). We term this phase  $Ps_b$   
152 and focus on its arrival time on the radial RF relative to the direct P-arrival on the vertical RF to  
153 estimate the depth to basement. Both an increase in Poisson's ratio (larger  $V_p/V_s$  ratio) and the  
154 thickness of the sedimentary layer leads to a later arrival of the  $Ps_b$  phase (Zelt and Ellis, 1999).

### 155 **Seismic stations and receiver function computation**

156 In the last several years, there has been an expansion in the number of passive seismic deployments  
157 across South Australia (supplement Figure **S1**). While coverage is still non-uniform and occasion-  
158 ally sparse, the addition of new networks in increasingly remote locations allows for the investiga-  
159 tion of Earth's structure in under-explored regions within the continental interior that are typically  
160 less accessible to other methods. Most recently, the Marla Line (Liang and Kennett, 2020), Lake  
161 Eyre Basin (Eakin, 2019), and AusArray-SA (O'Donnell et al., 2020) experiments have increased  
162 coverage over the eastern margin of the Gawler Craton, and the various sedimentary sequences  
163 that cover it. The Lake Eyre Basin seismic array, as the name suggests, has increased coverage  
164 surrounding Kati Thanda-Lake Eyre and was also the first to install passive-recording seismome-  
165 ters in the Simpson Desert, where the sediment thicknesses are greatest within South Australia.  
166 This vast sandy desert contains the world's longest parallel sand dunes and was not crossed by  
167 motorized vehicle until 1962 (Wakelin-King and White, 2016).

168

169 Overall data from twelve temporary and two permanent seismic networks (AU and S1), with sta-  
170 tions located across South Australia, were used for analysis in this study (see Data and Resources  
171 Section; supplement Figure **S1**). In addition, some stations situated just beyond the state borders  
172 were included where available, such as the AQT network (1Q) in southwest Queensland. For the  
173 Marla Line experiment (3G), which had dense station spacing of  $< 4\text{km}$ , we selected only every 5<sup>th</sup>  
174 station. Both short-period and broadband seismic stations (channels SH\*, HH\*, BH\*, and EH\*)

175 were utilized for the receiver function analysis, comprising 243 individual seismic stations in total.

176

177 For each station, three-component data for earthquakes of  $M_w \geq 5.5$  and in the distance range  
178 of  $30^\circ - 95^\circ$  was sought (Figure S3 shows an example of earthquakes available for the station  
179 AEB15). For temporary networks, typically, a minimum of four months recording provided more  
180 than 75 earthquakes; however, on average, more than 150 earthquakes were available for each  
181 station. For permanent stations, earthquakes were sought from the previous year (2020), as this  
182 provided sufficient RFs for the analysis. The extracted 200 s of seismograms around the expected  
183 (using iasp91; Kennett and Engdahl, 1991) teleseismic P-arrival (50 s before and 150 s after) were  
184 then demeaned, detrended (linear), cosine tapered, and bandpass filtered between 0.1-1 Hz. Only  
185 earthquakes with a signal-to-noise ratio  $\geq 1.5$  (noise window: 45 s to 15 s prior to predicted P;  
186 signal window: 5 s before to 25 s after predicted P) were kept. The radial RFs were computed  
187 through the 'rf' Python-package (Eulenfeld, 2020), using an iterative time-domain deconvolution  
188 (Ligorria and Ammon, 1999). Each RF was then stretched and compressed on the time axis using  
189 a reference slowness of  $6.4 \text{ s}^\circ$  (moveout correction) since the arrival times of converted phases are  
190 influenced by the slowness of the ray as well. Further, to ascertain the quality of individual RFs and  
191 the subsequent stacks, only radial RFs where the largest arrival is a positive polarity peak within  
192 the first 2 s were kept. Stations with at least ten good RFs were then stacked (an example is shown  
193 in Figure 3c). Out of 243 stations, 231 met this criterion and were used for further analysis. All 231  
194 stacked RFs are presented in supplement Figure S4 and the corresponding list of measurements in  
195 supplement Table S1.

196

[Figure 3 about here.]

197 Using the RF stack for each station, the  $Ps_b$  arrival time (i.e. the P-wave converted to S-wave at  
198 the basement-sediment boundary) was determined with respect to the direct P-wave arrival on the  
199 vertical receiver function hereafter referred to as  $T_{Ps_b}$ . As illustrated in Figure 3, in the presence of  
200 sediments, the direct P-wave has near-vertical incidence beneath the seismic station and thus has  
201 a small amplitude on the radial RF. Instead, the largest positive peak on the radial RF is the  $Ps_b$   
202 phase, which arrives shortly after ( $\sim 1$  s for station AEB07) the direct P on the vertical RF. The  
203 relative time difference between the largest positive peak on the radial RF and the vertical RF is,

204 therefore, simply measured for each station to estimate  $T_{\text{Psb}}$ , which is primarily a function of the  
205 sediment thickness (Figure 3).

## 206 **Results and Discussion**

### 207 **Spatial variation of $T_{\text{Psb}}$**

208 The obtained  $T_{\text{Psb}}$  values for all 231 stations reveal striking geographical patterns (Figure 4). The  
209  $T_{\text{Psb}}$  values are highest in the northeast of the study area while lowest for stations on the southern  
210 Gawler Craton. The highest  $T_{\text{Psb}}$  value of 1.32 s was recorded at station SB03 (Skippy network;  
211 7B) situated atop the Cooper and Eromanga Basins. The majority (61%) of stations registered  $T_{\text{Psb}}$   
212  $\leq 0.2$  s, while 16 stations ( $\sim 7\%$ ) had  $T_{\text{Psb}} \geq 1$  s.

213  
214 Areas of elevation and outcropping basement - such as the southern Gawler Craton, Flinders  
215 Ranges, and Musgrave Province - have  $T_{\text{Psb}}$  values close to zero, indicating minimal sedimenta-  
216 tion in such regions (gray areas in Figure 4). On a smaller scale, stations OOD, AEB17, SD06,  
217 and AES15 installed on top of the Denison and Peake Inliers (DPI in Figure 2), a small base-  
218 ment outcrop to the west of Kati Thanda-Lake Eyre, display  $T_{\text{Psb}} \sim 0$  s, while surrounded by  
219 stations with  $T_{\text{Psb}} \geq 0.4$  s that lie beyond the basement inlier. Moderate values of  $T_{\text{Psb}}$  ( $\sim 0.5$  s)  
220 are co-located with regions with significant sediment accumulation, such as the Berri, Arrowie,  
221 Bight, and Arckaringa Basins. The highest values of  $T_{\text{Psb}}$  ( $\sim 1$  s or more) are located in regions  
222 with several overlapping sedimentary sequences, such as in the northeast where the Lake Eyre,  
223 Eromanga, Cooper, and Warburton Basins overlap. Surprisingly, for stations situated on top of  
224 the oldest basins such as Officer and Cariewerloo Basins, the  $T_{\text{Psb}}$  values are close to zero ( $\leq 0.2$   
225 s). Nonetheless, it is evident from the above that the  $T_{\text{Psb}}$  variations capture both small-scale and  
226 large-scale sedimentary features, and therefore can be utilized to estimate the depth to basement  
227 beneath each seismic station.

228 [Figure 4 about here.]

## 229 **Calibration of $T_{P_{sb}}$ with borehole basement depth**

230 Due to the resource potential, South Australia has an expansive dataset of borehole drill sites  
231 (Figure 2). This offers a unique opportunity to compare and calibrate the relationship between  
232  $T_{P_{sb}}$  and the basement depth. Of the 243 seismic stations located in South Australia (Figure 4),  
233 85% are located within  $0.5^\circ$  of a borehole site. It is often the case, however, that the seismic  
234 stations are located near multiple boreholes; therefore, in order to directly compare, an average  
235 value of borehole basement depth must be calculated for each station. We chose to interpolate  
236 the borehole basement depth values surrounding each station using the inverse distance weighting  
237 (IDW) method (Shepard, 1968). This method assigns a single basement depth value based on a  
238 weighted average of the borehole values within  $0.5^\circ$  of a station. If no borehole measurement was  
239 found within  $0.5^\circ$  of a station, no basement depth was assigned to it. The weighting given to each  
240 borehole value is based on the inverse of the distance to the station to the power of  $p$ , chosen here  
241 as 2. This takes into account the relative proximity of borehole points to the stations, thus making  
242 sure the interpolated basement depth at a station is dominated by the values which are closest and  
243 is less likely to be skewed by a single aberrant value.

244

245 [Figure 5 about here.]

246 The IDW method returned an interpolated basement depth beneath 182 seismic stations, repre-  
247 sented as circles in Figure 5. We divide the stations in two categories based on the age of sediments  
248 beneath the stations. Stations situated atop Phanerozoic sediments are shown as turquoise circles  
249 while stations on predominantly Proterozoic sediments are pink. Although there is some scatter,  
250 an overall positive correlation between  $T_{P_{sb}}$  and interpolated borehole basement depth can be rec-  
251 ognized, with larger  $T_{P_{sb}}$  arrival times corresponding to deeper borehole basement depths (Figure  
252 5). This positive trend is reinforced when the median values (with standard deviation) are plotted  
253 for each data bin (blue squares with errorbars, Figure 5). The gradient of the trend however is not  
254 constant, with basement depths gradually increasing for arrival times between 0 and  $\sim 0.6$  seconds  
255 and more steeply increasing thereafter. A linear regression on the binned median values (blue  
256 squares), using standard deviations as the weights, is performed to fit two lines that share  $T_{P_{sb}} =$   
257 0.58 s as a common point (yellow square). If a different common point is chosen, then the residual

258 error increases (supplement Figure S5). The fitted straight lines are of the form,

259

$$D = 366T_{P_{sb}}, \quad \text{for } T_{P_{sb}} < 0.58 \text{ s} \quad (1)$$

260 and

$$D = 3206.9T_{P_{sb}} - 1661.2, \quad \text{for } T_{P_{sb}} \geq 0.58 \text{ s}, \quad (2)$$

261 where  $T_{P_{sb}}$  is in seconds and D is the depth to basement in meters. It is to be noted, however  
262 that Equations 1 and 2 represent the best fitting linear trends, but the true relationship between the  
263 basement depth and  $T_{P_{sb}}$  might not be linear. A best-fitting exponential and quadratic relationship  
264 was also explored (supplement Figure S6), but both of these yielded higher RMSE values than the  
265 linear equations. Further, the non-linear curves diverge rapidly for  $T_{P_{sb}} > 1.5$  s. Additional data  
266 from deeper sedimentary basins (basement depth > 4000 m) may help further constrain them in the  
267 future.

268

269 As noted in Figure 3, basement depth and  $T_{P_{sb}}$  are related as,

$$D = \frac{T_{P_{sb}} \times V_{p_{sedi}} \times V_{s_{sedi}}}{V_{p_{sedi}} - V_{s_{sedi}}}, \quad (3)$$

270 where  $V_{p_{sedi}}$  and  $V_{s_{sedi}}$  are the average seismic velocities in the sediment layers. It is evident that  
271 calibrated Equations 1 and 2 are simplified forms of Equation 3 and that the gradient in Figure 5  
272 is a function of the seismic velocities of the sediments. As expected, deeper sedimentary basins  
273 (> 500 m) display a steeper gradient indicating faster seismic velocities that likely result from in-  
274 creased compaction at depth.

275

276 While the seismic velocities within sedimentary basins may vary in other settings, for our study  
277 region, the calibrated equations (1 and 2) seem to provide a reasonable fit to the available borehole  
278 data. We, therefore, use the calibrated equations to estimate the basement depth beneath all seismic  
279 stations (Figure 6), hereafter referred to as the RF basement depth.

### 280 **Comparison of seismically determined basement depth with the pattern of sedimentation** 281 **across South Australia (SA)**

282

[Figure 6 about here.]

283 Depth to basement estimated using the RF analysis correlates strongly with the extent of sed-  
284 imentary basins and spatial trends in sedimentation as suggested by the available borehole data  
285 (Figure 6). In the northeastern part of the study region, the RF estimated depths are highest,  
286 with values up to 2,600 m, due to the superimposition of Permian-Triassic Cooper and Jurassic-  
287 Cretaceous Eromanga Basins. Outside the boundary of Cooper Basin (green colored basin in  
288 Figure 6b), the basement depth decreases as seen in the borehole values. AQT network stations in  
289 Queensland expand the basement depth knowledge in this region. Station AQTK1, which is inside  
290 the Cooper basin, registers a depth of 2,352 m; in comparison, AQT08, located to the northwest  
291 outside of the Cooper Basin, records a shallower depth of 620 m, suggesting decreasing sediment  
292 thickness moving northwards. South of the Cooper Basin along the SA-NSW border, the basement  
293 depth gets progressively shallower (830 m at station CU01 to  $\sim 10$  m at station E1B1), represent-  
294 ing the thinning Eromanga Basin until its southernmost extent near Cockburn (Figure 1). South of  
295 Cockburn, the RF basement depth slightly increases again ( $\sim 150$  m) due to the Cenozoic Murray  
296 Basin bounded to the west by the Flinders Ranges. Notably, three stations (E1F1, CU35, AURMK)  
297 close to the SA-NSW-Victoria border show significantly deeper values ( $\sim 800$  m) due to the pres-  
298 ence of the concealed Mesozoic Berri Basin. This adds further constraints to the boundary of Berri  
299 Basin, which is more sparsely sampled by borehole drilling (Figure 6a).

300

301 Near the border between SA and the Northern Territory ( $26^\circ$  S), there is a dramatic change in  
302 the RF basement depth around longitude  $134^\circ$  E. East of this location, within the Simpson Desert,  
303 the RF estimated depth to the Proterozoic crystalline basement reaches up to 2,500 m due to the  
304 combined sediments from the Early Palaeozoic Warburton, Permian-Triassic Pedrika, Eromanga,  
305 and Cenozoic Lake Eyre Basins (Gatehouse, 1986; Wopfner et al., 1969). These stations in the  
306 Simpson Desert, which belong to the Lake Eyre Basin array, are the first to be deployed in the  
307 region, thus providing new basement depth constraints for hitherto under-explored parts of Aus-  
308 tralia. West of  $134^\circ$  E marks the low-sediment region of the exhumed Musgrave province, the  
309 result of the Petermann Orogeny around 570–530 Ma (e.g. Wade et al., 2008). Stations located  
310 within the Musgrave province record typically shallow RF basement depths of  $< 25$  m. The Marla  
311 Line experiment (3G), a dense east-west linear transect of seismic stations, further illustrates the  
312 decreasing sediment thickness from the east (870 m) to the west (10 m) due to thinning of the Ero-

313 manga Basin. South of Oodnadatta, four stations (OOD, AEB17, SD06, AES15) show shallower  
314 basement depth (< 70 m) compared to the surrounding stations (~200 m). These stations sit on  
315 Neoproterozoic Denison and Peake Inliers surrounded by sediments from Eromanga Basin with  
316 Permian Arckaringa Basin sediments to the west (Drexel and Preiss, 1995).

317

318 The southern part of the Gawler Craton has areas of exposed Paleoproterozoic-Archean crust  
319 with little to no sediment cover, which is well represented by estimated basement depths of < 50 m.  
320 However, for stations atop Proterozoic Basins like the Officer, Amadeus, and Cariewerloo Basins  
321 the RF basement depths do not match the borehole values (pink circles in Figure 5). Within the  
322 boundaries of Mesoproterozoic Cariewerloo Basin, this is especially evident (Fanning et al., 1983),  
323 with borehole basement depths of up to 1,500 m. Contrastingly, the RF estimated basement depth  
324 is < 100 m. This can be explained as older Proterozoic sediments are often heavily metamorphosed  
325 with increased seismic velocities and  $V_p$  values potentially greater than 6 km/s (Wang et al., 2016),  
326 therefore inhibiting a strong impedance contrast with the crustal basement. Given the propensity of  
327 sediment-hosted base metals to be hosted in such Proterozoic basins (Hitzman et al., 2010; Leach  
328 et al., 2010), our method may be useful in determining the top of Proterozoic sediments instead.  
329 These regions of older and seismically faster Proterozoic basins produce the anomalous cluster of  
330 data points in Figure 5 (pink circles) with borehole depth values between 500-1000 m for  $T_{Ps_b} <$   
331 0.2 s.

### 332 **Uncertainties in RF basement depth estimation**

333 The RF estimated basement depths appear to match the sedimentary basin and borehole values  
334 reasonably well. The typical error expected for any measurement is  $\pm 134$  m for the shallower  
335 (excluding Proterozoic Basins) and  $\pm 360$  m for the deeper sedimentary basins, as estimated from  
336 the size of the error-bars in Figure 5 (i.e. the average standard deviations of the binned data).  
337 These typical error ranges are likely influenced by some of the inherent uncertainties underlying  
338 our chosen method. One of our primary assumptions is that in the presence of sediments, there  
339 is insignificant P-wave energy on the horizontal components; thus, the highest amplitude on the  
340 radial RF is due to the  $Ps_b$  phase. However, in regions of thin or well-compacted sediments, this  
341 might not always be the case, and there could be a potential overlap of direct P and  $Ps_b$  phase,

342 which may slightly underestimate the RF derived basement depth.

343

344 Another assumption is that the largest impedance contrast, representing the first and largest am-  
345 plitude peak in the RF, is generated by the boundary between the crustal basement and overlying  
346 sediments. However, as previously discussed, this appears to be limited by age, and is only ap-  
347 plicable for Phanerozoic sediments, not older Proterozoic sedimentary basins. Older sediments  
348 are more likely to be metamorphosed, resulting in higher seismic velocities, with less distinc-  
349 tion between ‘sediments’ and the underlying basement. In such cases, the  $T_{P_{Sb}}$  values can appear  
350 anomalously low ( $< 0.2$  s), which may be misinterpreted as a thin layer of sediments ( $< 200$  m) if  
351 the geological context is not considered or prior geological information is lacking.

## 352 **Conclusion and future implications**

353 We present a novel yet simple method to estimate the depth to basement in the presence of varying  
354 sediment thickness using receiver functions. Without prior knowledge of the sub-surface velocity  
355 structure, the depth to the basement is estimated using the relative arrival time of the P-converted-  
356 to-S phase generated at the base of the sediments. The method was demonstrated using data from  
357 seismic stations across South Australia and is able to capture the known variations in Phanerozoic  
358 sediments across the region. Despite an extensive existing borehole database for South Australia,  
359 we are able to improve the basement depth database, particularly in areas of deep sediments and  
360 remote areas such as the Simpson, Strzelecki, and Sturt Stony Deserts.

361

362 Despite the uncertainties, the methodology showcased here has three main advantages. Firstly, it  
363 can be employed to get a quick and credible assessment of sediment thickness using only temporary  
364 seismic stations deployed for less than 6 months, with little or no knowledge of the sub-surface  
365 velocity structure. Thus, it could serve as a pre-drill strategy to estimate the basement depth before  
366 physically deploying drilling equipment. This is particularly beneficial for deep basins, where  
367 drilling becomes increasingly expensive for deeper boreholes. Secondly, using temporary seismic  
368 stations for a few months is relatively inexpensive, and the acquired passive-seismic data will  
369 additionally be useful for many other applications to image the Earth’s interior. Thirdly, while

370 our method can't determine the base of Proterozoic basins, it may help determine the top of such  
371 metamorphosed sediments, which is desirable for future sediment-hosted base metal exploration  
372 projects, especially in places like Australia. Therefore, this method has immense potential for  
373 under-explored regions where rich mineral resources may currently be hidden under cover.

## 374 **Data and Resources**

375 The supplemental figures referenced in the text are included in the accompanying Supplemental  
376 Material file. Figure S1 shows the seismic stations available for the study and Figure S4 depicts  
377 the individual stacked receiver functions for those stations. An example for earthquakes used for a  
378 typical station is shown in Figure S3. Figure S5 and S6 contain four cases of regression analysis  
379 where RMSE error was found to be higher than Figure 5. The borehole basement depth vari-  
380 ation in South Australia is portrayed in Figure S2 via histograms. Further, Table S1 (uploaded  
381 as a separate text file) contains the individual station measurements for  $T_{P_{sb}}$  and the estimated  
382 basement depth. The earthquakes used in this study were collected from Incorporated Research  
383 Institutions for Seismology (IRIS) Data Services (<http://ds.iris.edu/ds/>, last accessed May 2021).  
384 The seismic waveform data from 14 seismic networks was used (last accessed May 2021): 5G  
385 (doi:10.7914/SN/5G\_2018); 6K (doi:10.7914/SN/6K\_2020); 5J ([http://www.fdsn.org/networks/detail/5J\\_2017/](http://www.fdsn.org/networks/detail/5J_2017/));  
386 7B (doi:10.7914/SN/7B\_1993); 1F (doi:10.7914/SN/1F\_2009); 6F (doi:10.7914/SN/6F\_2008); YJ  
387 ([http://www.fdsn.org/networks/detail/YJ\\_2009/](http://www.fdsn.org/networks/detail/YJ_2009/)); 7K (doi:10.7914/SN/7K\_2007); 7I (doi:10.7914/SN/7I\_2003);  
388 1G (doi:10.7914/SN/1G\_2008); 1Q (doi:10.7914/SN/1Q\_2016); 3G (doi:10.7914/SN/3G\_2018);  
389 S1 (doi:10.7914/SN/S1); AU (<https://www.fdsn.org/networks/detail/AU/>). The Python package  
390 'rf' version 1.0.0 (Eulenfeld, 2020) was used to compute the receiver functions. All the seismic  
391 data was handled using the Python package Obspy (<https://docs.obspy.org/>; Krischer et al., 2015).  
392 Borehole data used in the study was obtained from the South Australian Resources Information  
393 Gateway and is titled 'Crystalline basement intersecting drillholes' (<https://map.sarig.sa.gov.au/>;  
394 last accessed May 2021). All the geological data used in the study was acquired from the Geo-  
395 science Australia Portal (<https://portal.ga.gov.au/>; last accessed July 2021). Information about the  
396 active/operating mines in South Australia was obtained from the Department of Energy and Min-  
397 ing, Government of South Australia (<https://energymining.sa.gov.au/>; last accessed July 2021).

398 Plots were made using the Generic Mapping Tools, Version 6.1.1 (<https://www.generic-mapping->  
399 [tools.org/](https://www.generic-mapping-tools.org/); Wessel et al., 2019) and matplotlib version 3.4.2 (<https://matplotlib.org/>).

## 400 **Declaration of Competing Interests**

401 The authors declare no competing interests.

## 402 **Acknowledgements**

403 The work was supported by Australian Research Council Grant DE190100062. The Lake Eyre  
404 Basin seismic array and many other previous temporary deployments were made possible via fund-  
405 ing from AuScope ([auscope.org.au](https://auscope.org.au)), instrumentation from the Australian National Seismic Imag-  
406 ing Resource (ANSIR), and contributions from staff at the Research School of Earth Sciences,  
407 Australian National University. The AusArray SA deployment is supported by the Geological Sur-  
408 vey of South Australia, with instrumentation from ANSIR and Geoscience Australia. The GSSA  
409 thanks Bruce Goleby, Ann Goleby, Isaac Axford, Kate Selway, John Stephenson, Alexei Gorbatov,  
410 Michelle Salmon, Robert Pickle, and Colin Telfer for their contribution to AusArray SA. We are  
411 incredibly grateful to landholders, traditional owners, and the Department of Defence for grant-  
412 ing land access for the seismic arrays. JPOD publishes with the permission of the Director of the  
413 Geological Survey of South Australia.

## 414 **References**

- 415 Ali, S. H., Giurco, D., Arndt, N., Nickless, E., Brown, G., Demetriades, A., Durrheim, R., En-  
416 rriquez, M. A., Kinnaird, J., Littleboy, A., et al. (2017). Mineral supply for sustainable develop-  
417 ment requires resource governance. *Nature*, 543(7645):367–372.
- 418 Ammon, C. J. (1991). The isolation of receiver effects from teleseismic p waveforms. *Bulletin of*  
419 *the seismological Society of America*, 81(6):2504–2510.
- 420 Beyer, S., Kyser, K., Polito, P., and Fraser, G. (2018). Mesoproterozoic rift sedimentation, fluid

- 421 events and uranium prospectivity in the cariewerloo basin, gawler craton, south australia. *Aus-*  
422 *tralian Journal of Earth Sciences*, 65(3):409–426.
- 423 Birkey, A., Ford, H. A., Dabney, P., and Goldhagen, G. (2021). The lithospheric architecture  
424 of australia from seismic receiver functions. *Journal of Geophysical Research: Solid Earth*,  
425 126(4):e2020JB020999.
- 426 Brown, C. M. and Stephenson, A. E. (1991). Geology of the murray basin, southeastern australia.  
427 *Bulletin-Australia, Bureau of Mineral Resources, Geology and Geophysics*, (235).
- 428 Collett, D. and McFadden, P. (2014). Uncover initiative: ushering in a new era of exploration in  
429 australia.
- 430 Cunningham, E. and Lekic, V. (2020). Constraining properties of sedimentary strata using receiver  
431 functions: An example from the atlantic coastal plain of the southeastern united states. *Bulletin*  
432 *of the Seismological Society of America*, 110(2):519–533.
- 433 Drexel, J. and Preiss, W. (1995). The geology of south australia, vol. 2. the phanerozoic: South  
434 australia geological survey. *Bulletin*, 55.
- 435 Eakin, C. (2019). Seismicity, minerals, and craton margins: The lake eyre basin seismic deploy-  
436 ment. *ASEG Extended Abstracts*, 2019(1):1–2.
- 437 Eulenfeld, T. (2020). rf: Receiver function calculation in seismology. *Journal of Open Source*  
438 *Software*, 5(48):1808.
- 439 Fanning, C., Flint, R., and Preiss, W. (1983). Geochronology of the pandurra formation. *Geologi-*  
440 *cal Survey of South Australia Quarterly Geological Notes*, 88:11–16.
- 441 Gatehouse, C. G. (1986). The geology of the warburton basin in south australia. *Australian Journal*  
442 *of Earth Sciences*, 33(2):161–180.
- 443 Hand, M., Reid, A., and Jagodzinski, L. (2007). Tectonic framework and evolution of the gawler  
444 craton, southern australia. *Economic Geology*, 102(8):1377–1395.
- 445 Hitzman, M. W., Selley, D., and Bull, S. (2010). Formation of sedimentary rock-hosted stratiform  
446 copper deposits through earth history. *Economic Geology*, 105(3):627–639.

- 447 Hoggard, M. J., Czarnota, K., Richards, F. D., Huston, D. L., Jaques, A. L., and Ghelichkhan,  
448 S. (2020). Global distribution of sediment-hosted metals controlled by craton edge stability.  
449 *Nature Geoscience*, 13(7):504–510.
- 450 Kennett, B. and Engdahl, E. (1991). Traveltimes for global earthquake location and phase identi-  
451 fication. *Geophysical Journal International*, 105(2):429–465.
- 452 Kennett, B., Salmon, M., Saygin, E., and Group, A. W. (2011). Ausmoho: the variation of moho  
453 depth in australia. *Geophysical Journal International*, 187(2):946–958.
- 454 Krischer, L., Megies, T., Barsch, R., Beyreuther, M., Lecocq, T., Caudron, C., and Wassermann, J.  
455 (2015). Obspy: A bridge for seismology into the scientific python ecosystem. *Computational*  
456 *Science & Discovery*, 8(1):014003.
- 457 Langston, C. A. (1979). Structure under mount rainier, washington, inferred from teleseismic body  
458 waves. *Journal of Geophysical Research: Solid Earth*, 84(B9):4749–4762.
- 459 Leach, D. L., Bradley, D. C., Huston, D., Pisarevsky, S. A., Taylor, R. D., and Gardoll, S. J. (2010).  
460 Sediment-hosted lead-zinc deposits in earth history. *Economic Geology*, 105(3):593–625.
- 461 Liang, S. and Kennett, B. L. (2020). Passive seismic imaging of a craton edge–central australia.  
462 *Tectonophysics*, 797:228662.
- 463 Ligorria, J. P. and Ammon, C. J. (1999). Iterative deconvolution and receiver-function estimation.  
464 *Bulletin of the seismological Society of America*, 89(5):1395–1400.
- 465 Lunedei, E. and Albarello, D. (2010). Theoretical hvsr curves from full wavefield modelling of  
466 ambient vibrations in a weakly dissipative layered earth. *Geophysical Journal International*,  
467 181(2):1093–1108.
- 468 Mudd, G., Czarnota, K., Skirrow, R. G., McAlpine, S., Yuan, Y., Yellishetty, M., Weng, Z.-H., and  
469 Werner, T. (2019). *Critical Minerals in Australia: A review of opportunities and research needs*.  
470 Geoscience Australia.
- 471 Mudd, G. M. (2007). *The sustainability of mining in Australia: key production trends and their*  
472 *environmental implications for the future*. Department of Civil Engineering, Monash University.

- 473 Okada, H. (2006). Theory of efficient array observations of microtremors with special reference to  
474 the spac method. *Exploration Geophysics*, 37(1):73–85.
- 475 O’Donnell, J., Thiel, S., Robertson, K., Gorbatov, A., and Eakin, C. (2020). Using seismic tomog-  
476 raphy to inform mineral exploration in south australia: the ausarray sa broadband seismic array.  
477 *MESA Journal*, 93:24–31.
- 478 Pain, C. F., Pillans, B. J., Roach, I. C., Worrall, L., and Wilford, J. R. (2012). Old, flat and red–  
479 australia’s distinctive landscape. *Shaping a nation: A geology of Australia*, pages 227–275.
- 480 Piana Agostinetti, N., Martini, F., and Mongan, J. (2018). Sedimentary basin investigation using re-  
481 ceiver function: an east african rift case study. *Geophysical Journal International*, 215(3):2105–  
482 2113.
- 483 Radke, B. (2009). *Hydrocarbon and geothermal prospectivity of sedimentary basins in central*  
484 *Australia: Warburton, Cooper, Pedirka, Galilee, Simpson and Eromanga basins*. Geoscience  
485 Australia.
- 486 Rawlinson, N., Stephenson, R., and Carbonell, R. (2017). Seismic imaging at the cross-roads:  
487 Active, passive, exploration and solid earth. *Tectonophysics*, 718:1–8. Seismix 2016: Advances  
488 in active and passive seismic imaging of continents and their margins.
- 489 Raymond, O., Totterdell, J., Woods, M., and AJ, S. (2018). Australian geological provinces  
490 2018.01 edition. *Geoscience Australia, Canberra*.
- 491 Reid, A. (2019). The olympic cu-au province, gawler craton: a review of the lithospheric archi-  
492 tecture, geodynamic setting, alteration systems, cover successions and prospectivity. *Minerals*,  
493 9(6):371.
- 494 Sheehan, A. F., Abers, G. A., Jones, C. H., and Lerner-Lam, A. L. (1995). Crustal thickness  
495 variations across the colorado rocky mountains from teleseismic receiver functions. *Journal of*  
496 *Geophysical Research: Solid Earth*, 100(B10):20391–20404.
- 497 Shepard, D. (1968). A two-dimensional interpolation function for irregularly-spaced data. In  
498 *Proceedings of the 1968 23rd ACM national conference*, pages 517–524.

- 499 Southgate, P., Neumann, N., and Gibson, G. (2013). Depositional systems in the Mt Isa Inlier from  
500 1800 Ma to 1640 Ma: Implications for Zn–Pb–Ag mineralisation. *Australian Journal of Earth*  
501 *Sciences*, 60(2):157–173.
- 502 Srinivas, D., Srinagesh, D., Chadha, R., and Ravi Kumar, M. (2013). Sedimentary thickness  
503 variations in the Indo-Gangetic foredeep from inversion of receiver functions. *Bulletin of the*  
504 *Seismological Society of America*, 103(4):2257–2265.
- 505 Tao, K., Liu, T., Ning, J., and Niu, F. (2014). Estimating sedimentary and crustal structure using  
506 wavefield continuation: theory, techniques and applications. *Geophysical Journal International*,  
507 197(1):443–457.
- 508 Tauzin, B., van der Hilst, R. D., Wittlinger, G., and Ricard, Y. (2013). Multiple transition zone seis-  
509 mic discontinuities and low velocity layers below western United States. *Journal of Geophysical*  
510 *Research: Solid Earth*, 118(5):2307–2322.
- 511 Vinnik, L. (1977). Detection of waves converted from P to SV in the mantle. *Physics of the Earth*  
512 *and Planetary Interiors*, 15(1):39–45.
- 513 Wade, B., Kelsey, D., Hand, M., and Barovich, K. (2008). The Musgrave Province: stitching north,  
514 west and south Australia. *Precambrian Research*, 166(1-4):370–386.
- 515 Wakelin-King, G. and White, S. (2016). The national heritage potential of landscapes within the  
516 Australian drylands. *Geoheritage*, 8(2):105–118.
- 517 Wang, M., Hubbard, J., Plesch, A., Shaw, J. H., and Wang, L. (2016). Three-dimensional seismic  
518 velocity structure in the Sichuan Basin, China. *Journal of Geophysical Research: Solid Earth*,  
519 121(2):1007–1022.
- 520 Wessel, P., Luis, J., Uieda, L., Scharroo, R., Wobbe, F., Smith, W., and Tian, D. (2019). The  
521 generic mapping tools version 6. *Geochemistry, Geophysics, Geosystems*, 20(11):5556–5564.
- 522 Wopfner, H. et al. (1969). *Depositional history and tectonics of South Australian sedimentary*  
523 *basins*. Economic Commission for Asia and the Far East Committee on Industry and . . . .

- 524 Yeates, A., Gibson, D., Towner, R., and Crowe, R. (1984). Regional geology of the onshore  
525 canning basin, wa.
- 526 Yeck, W. L., Sheehan, A. F., and Schulte-Pelkum, V. (2013). Sequential h- $\kappa$  stacking to obtain  
527 accurate crustal thicknesses beneath sedimentary basins. *Bulletin of the Seismological Society*  
528 *of America*, 103(3):2142–2150.
- 529 Yu, Y., Song, J., Liu, K. H., and Gao, S. S. (2015). Determining crustal structure beneath seis-  
530 mic stations overlying a low-velocity sedimentary layer using receiver functions. *Journal of*  
531 *Geophysical Research: Solid Earth*, 120(5):3208–3218.
- 532 Zang, W.-L., Jago, J., Alexander, E., and Paraschivoiu, E. (2004). A review of basin evolution,  
533 sequence analysis and petroleum potential of the frontier arrowie basin, south australia.
- 534 Zelt, B. and Ellis, R. (1999). Receiver-function studies in the trans-hudson orogen, saskatchewan.  
535 *Canadian Journal of Earth Sciences*, 36(4):585–603.
- 536 Zheng, T., Zhao, L., and Chen, L. (2005). A detailed receiver function image of the sedimentary  
537 structure in the bohai bay basin. *Physics of the Earth and Planetary Interiors*, 152(3):129–143.

538 **Author postal address**

539 **Shubham Agrawal**

540 142 Mills Road, ANU, Acton, ACT 0200, Australia.

541 **Caroline Eakin**

542 142 Mills Road, ANU, Acton, ACT 0200, Australia.

543 **John O'Donnell**

544 Level 4, 11 Waymouth Street, Adelaide, South Australia 5000, Australia.

545 **List of Figures**

546 1 A comprehensive on-shore sedimentary basin map of South Australia. Basin infor-  
547 mation is acquired from the Geoscience Australia database (Raymond et al., 2018).  
548 Gray circles with white crosses are major approved/operating mines (see Data and  
549 Resources Section). The dashed blue line is the inferred boundary of the Gawler  
550 Craton. Abbreviations are as follows - Cr.B: Cariewerloo Basin, DPI: Denison and  
551 Peake inliers, KT-LE: Kati Thanda-Lake Eyre, LT: Lake Torrens, WA: Western  
552 Australia, NT: Northern Territory, QLD: Queensland, NSW: New South Wales,  
553 VIC: Victoria. For the color version of this figure, refer to the electronic edition. . . . . 24

554 2 Individual borehole basement depth measurements (colored squares) for the state  
555 of South Australia (see Data and Resources Section). The black dashed lines are  
556 geological provinces (MP: Musgrave Province; GC: Gawler Craton; ARC: Ade-  
557 laide Rift Complex) with little or no sediment cover (Raymond et al., 2018). The  
558 color scale is adjusted in the upper 200 km to encapsulate the variations in depth.  
559 For the color version of this figure, refer to the electronic edition. . . . . 25

560 3 a) Schematic ray paths of P and S waves traversing a sediment layer (of thickness  
561 D) in response to an incoming P wave from an epicentral distance of  $60^\circ$  for the  
562 shown velocity structure. b) An example of  $T_{Ps_b}$  estimation from receiver functions  
563 at typical station AEB07. The station is located on thick sediment cover (see Figure  
564 4 for location), hence the significant delay in the  $Ps_b$  phase relative to direct P. The  
565 formula for  $T_{Ps_b}$  is derived assuming vertical incidence of P and S waves beneath  
566 the station. c) Individual receiver functions that passed the quality control for  
567 station AEB07, which were used to calculate the stacked receiver function. The  
568 right panel provides epicentral distance (red dots) and backazimuth (blue dots)  
569 values for each receiver function. For the color version of this figure, refer to the  
570 electronic edition. . . . . 26

571 4 Values for times of arrivals of P-to-S converted phase ( $T_{Ps_b}$ ) at the basement for  
572 seismic stations in South Australia.  $T_{Ps_b}$  values are estimated from receiver func-  
573 tions, as illustrated in Figure 3. Transparent triangles are stations where receiver  
574 functions did not pass quality control. Stations names highlighted with green are  
575 stations discussed in the text. Gray shaded areas are the same as the dashed geolog-  
576 ical provinces in Figure 2. KT-LE: Kati Thanda-Lake Eyre. For the color version  
577 of this figure, refer to the electronic edition. . . . . 27

578 5 Statistical comparison of the receiver function estimated  $T_{Ps_b}$  and borehole base-  
579 ment depth beneath 182 seismic stations, plotted as circles. Turquoise circles are  
580 stations situated atop Phanerozoic Basins while pink on Proterozoic Basins. Dark  
581 blue squares are the binned median values for every 0.9 s, with vertical solid blue  
582 lines representing the standard deviation. The yellow square denotes the point of  
583 inflection in the data, about which two linear equations (maroon dotted lines with  
584 equations at top) are regressed. RMSE: Root Mean Square Error. For the color  
585 version of this figure, refer to the electronic edition. . . . . 28

586        6     Receiver function estimated basement depth (colored triangles) juxtaposed with  
587            (a) borehole measurements (same as Figure 2) and (b) sedimentary basins (same  
588            as Figure 1). The black dots are stations that didn't pass quality control. It is to be  
589            noted that in (b), the basins are colored by the age of sediments, while the stations  
590            are colored according to the RF basement depth, given by the legend below the  
591            figure. For the color version of this figure, refer to the electronic edition. . . . . 29

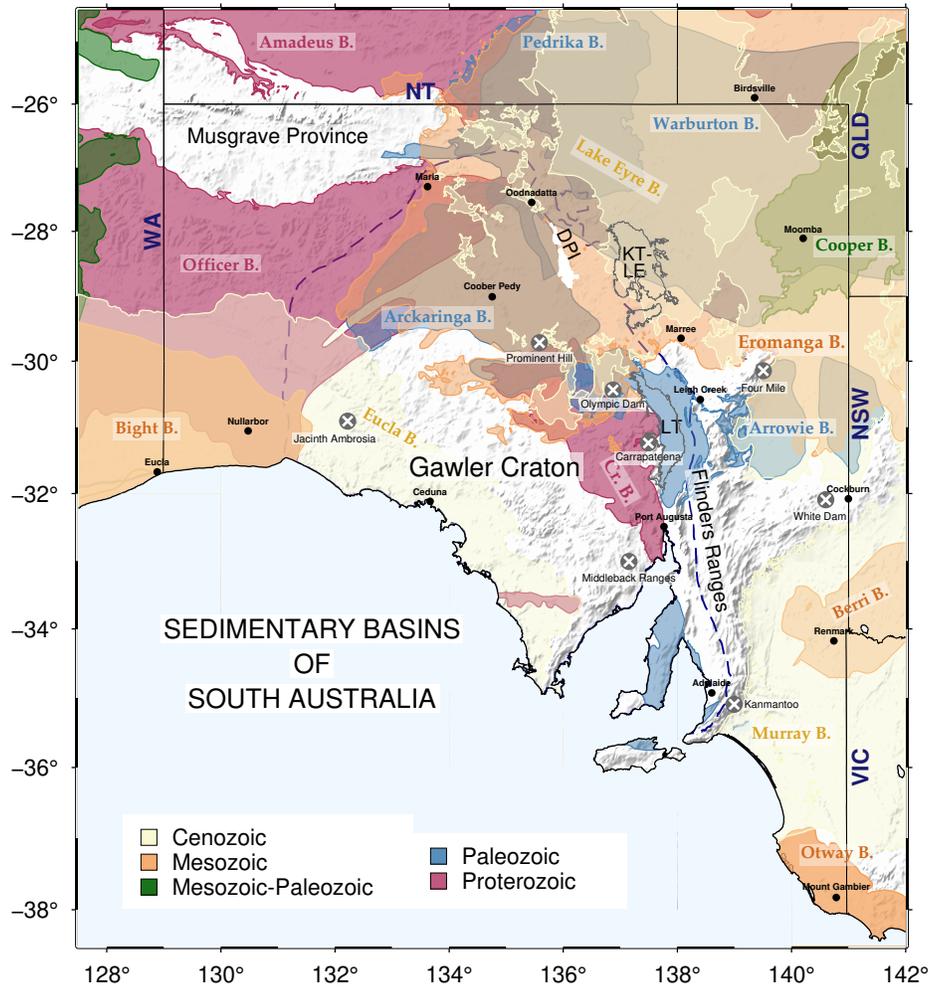


Figure 1. A comprehensive on-shore sedimentary basin map of South Australia. Basin information is acquired from the Geoscience Australia database (Raymond et al., 2018). Gray circles with white crosses are major approved/operating mines (see Data and Resources Section). The dashed blue line is the inferred boundary of the Gawler Craton. Abbreviations are as follows - Cr.B: Cariewerloo Basin, DPI: Denison and Peake inliers, KT-LE: Kati Thanda-Lake Eyre, LT: Lake Torrens, WA: Western Australia, NT: Northern Territory, QLD: Queensland, NSW: New South Wales, VIC: Victoria. For the color version of this figure, refer to the electronic edition.

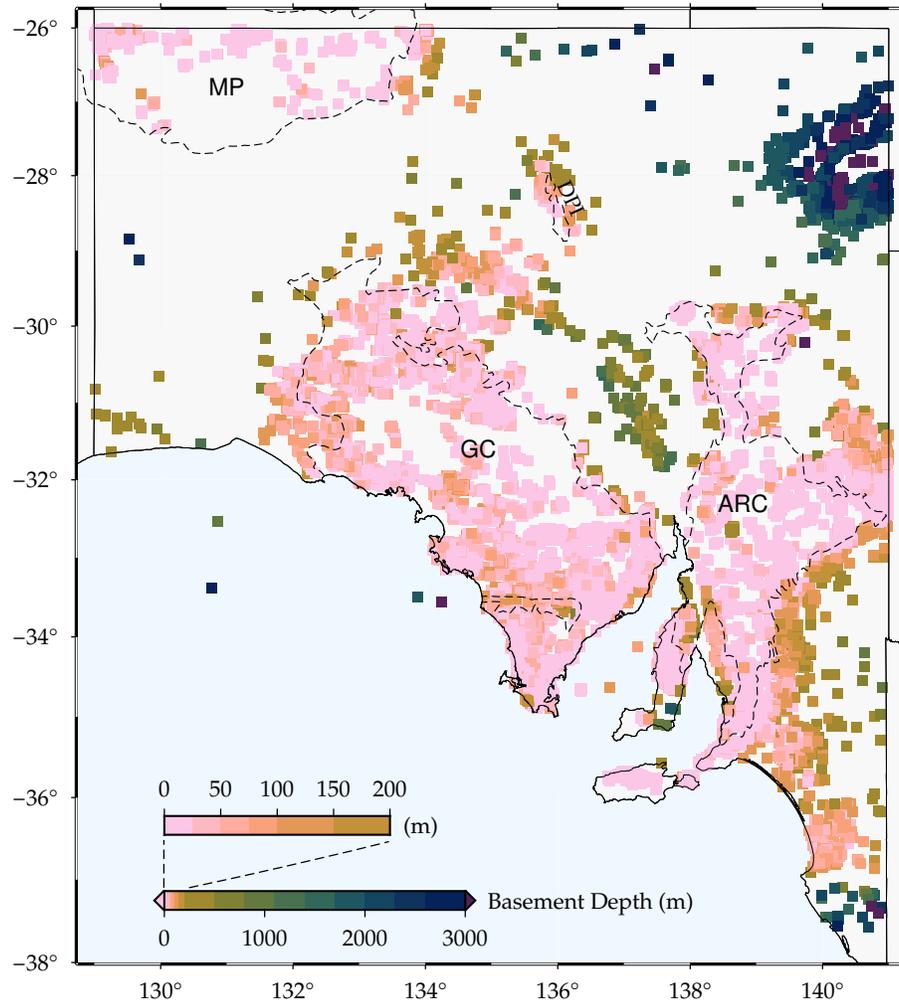


Figure 2. Individual borehole basement depth measurements (colored squares) for the state of South Australia (see Data and Resources Section). The black dashed lines are geological provinces (MP: Musgrave Province; GC: Gawler Craton; ARC: Adelaide Rift Complex) with little or no sediment cover (Raymond et al., 2018). The color scale is adjusted in the upper 200 km to encapsulate the variations in depth. For the color version of this figure, refer to the electronic edition.

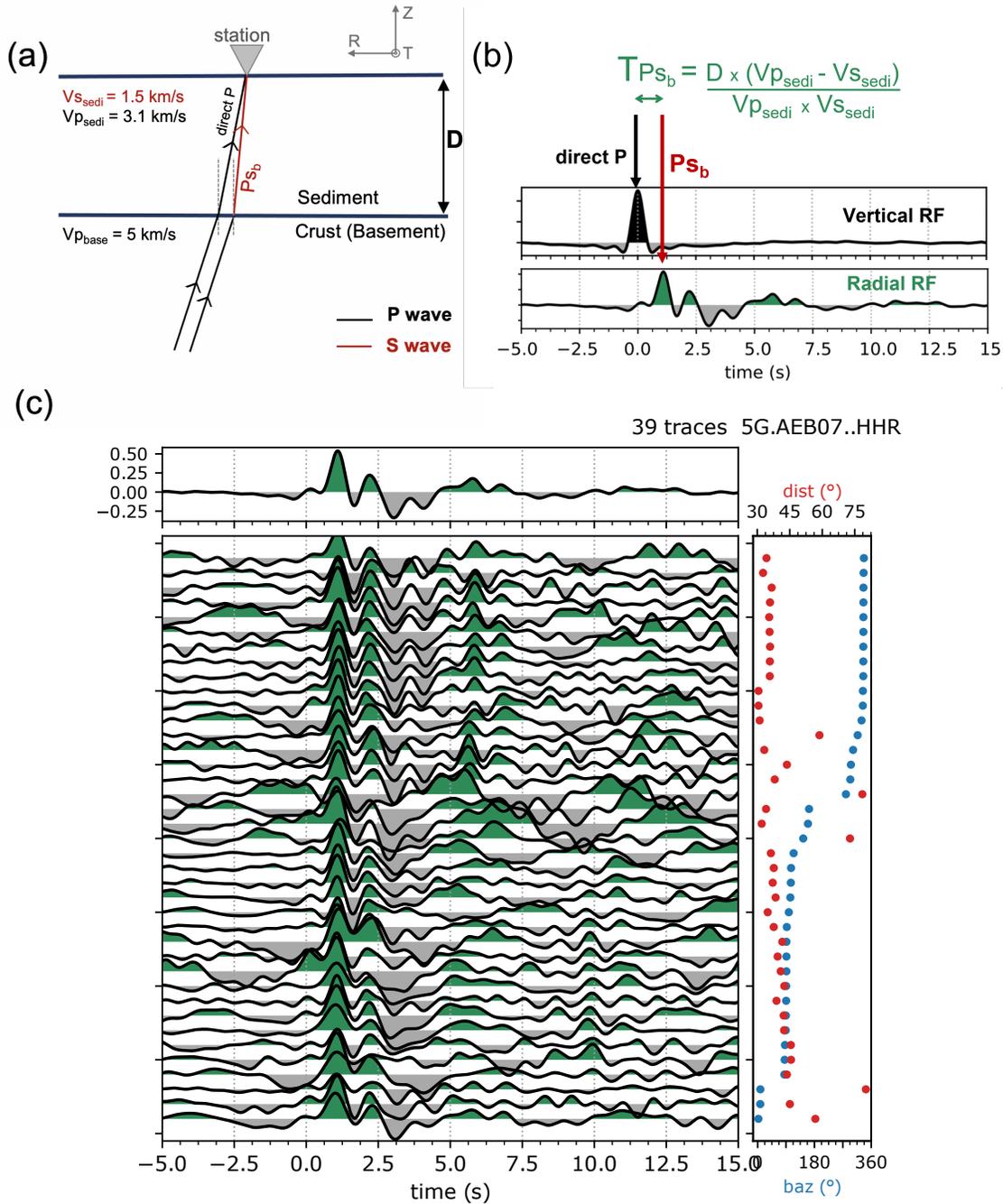


Figure 3. a) Schematic ray paths of P and S waves traversing a sediment layer (of thickness  $D$ ) in response to an incoming P wave from an epicentral distance of  $60^\circ$  for the shown velocity structure. b) An example of  $T_{Ps_b}$  estimation from receiver functions at typical station AEB07. The station is located on thick sediment cover (see Figure 4 for location), hence the significant delay in the  $P_{sb}$  phase relative to direct P. The formula for  $T_{Ps_b}$  is derived assuming vertical incidence of P and S waves beneath the station. c) Individual receiver functions that passed the quality control for station AEB07, which were used to calculate the stacked receiver function. The right panel provides epicentral distance (red dots) and backazimuth (blue dots) values for each receiver function. For the color version of this figure, refer to the electronic edition.



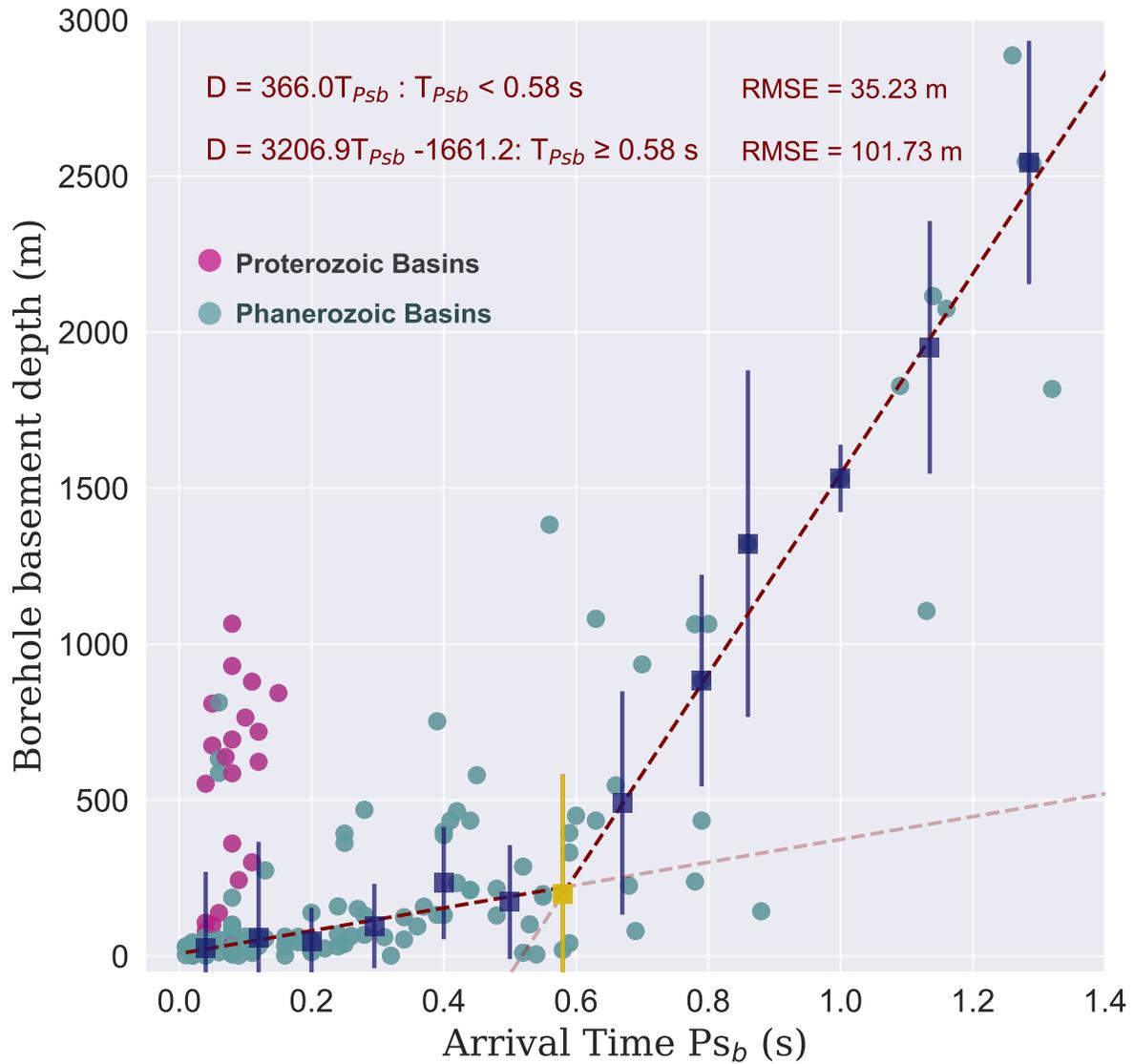


Figure 5. Statistical comparison of the receiver function estimated  $T_{Ps_b}$  and borehole basement depth beneath 182 seismic stations, plotted as circles. Turquoise circles are stations situated atop Phanerozoic Basins while pink on Proterozoic Basins. Dark blue squares are the binned median values for every 0.9 s, with vertical solid blue lines representing the standard deviation. The yellow square denotes the point of inflection in the data, about which two linear equations (maroon dotted lines with equations at top) are regressed. RMSE: Root Mean Square Error. For the color version of this figure, refer to the electronic edition.

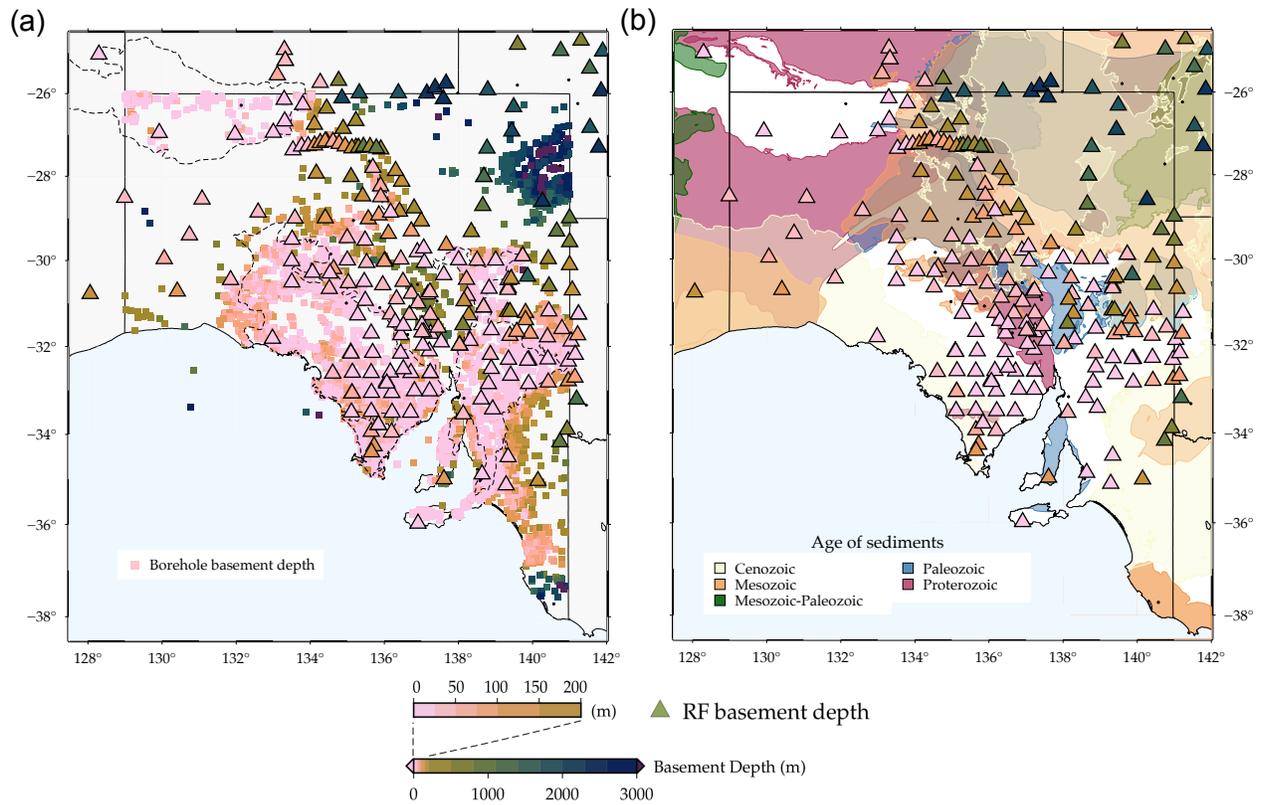


Figure 6. Receiver function estimated basement depth (colored triangles) juxtaposed with (a) borehole measurements (same as Figure 2) and (b) sedimentary basins (same as Figure 1). The black dots are stations that didn't pass quality control. It is to be noted that in (b), the basins are colored by the age of sediments, while the stations are colored according to the RF basement depth, given by the legend below the figure. For the color version of this figure, refer to the electronic edition.