

1 **Title: Evidence of Matuyama-Brunhes transition in the cave sediment in Central Europe**

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8

## 9 **Abstract**

10 In this study, we identified the Matuyama-Brunhes magnetic reversal recorded in cave  
11 sediments in Central Europe, Czech Republic. We collected discrete samples from the  
12 homogeneous sedimentary profile in the Za Hajovnou cave located in the eastern part of the  
13 Czech Republic. Novel use of characteristic remanent magnetization (ChRM) directions and  
14 VGP path of the data revealed Matuyama-Brunhes transition boundary within 5 cm of the Za  
15 Hajovnou cave sediment. This result revealed a new more detailed behavior of the polarity  
16 transition from the central European location. Migration of the paleopole between east of  
17 Africa and west of North America is a significant marker in terms of the central European  
18 paleomagnetic record. Also we estimated the sedimentation rate of the cave. In addition, we  
19 discussed our results with a supporting data associated with tektites in the light of a new  
20 hypothesis that a meteorite impact could be a reason of Matuyama-Brunhes reversal.

21 **Keywords:** Matuyama-Brunhes, paleomagnetism, magnetic reversal, cave, sediments,  
22 meteorite impact

## 23 **1. Introduction**

24 Matuyama-Brunhes magnetic reversal occurred approximately 781 kyr ago (Lourens et al.,  
25 2004). Studies in recent years (Channel et al., 2010; Giaccio et al., 2013; Jin and Liu, 2011;

26 Kitaba et al., 2013; Liu et al., 2016; Okada et al., 2017; Pares et al., 2016; Sagnotti et al.,  
27 2010, 2014; Suganuma et al., 2010; Valet et al., 2014; Bella et al., 2019) have shown that this  
28 event is well recorded by respective sediments that had sufficient sedimentation rate and  
29 could be analyzed, in detail, by paleomagnetism.

30 Sediments acquire remanent magnetization during their deposition. The alignment of  
31 ferromagnetic grains occurs in the direction of the earth's magnetic field and acquisition of  
32 primary magnetization due to this sedimentation process is called depositional or detrital  
33 remanent magnetization (DRM) (Gubbins and Herrero, 2017). Remanent magnetization  
34 protected by energy barriers can last over geologic time scales. Nevertheless, due to thermal  
35 and/or chemical processes such as reheating, oxidation and formation of iron hydroxides  
36 during time, barriers may be overcome, the magnetic domains change their arrangement and  
37 rocks can acquire secondary magnetizations. The new secondary magnetization has an  
38 orientation in the direction of the Earth's field. Then rocks can acquire a viscous  
39 magnetization (VRM) long time after their formation due to an exposure to geomagnetic field.  
40 VRM contributes to a noise in paleomagnetic data (Lanza and Meloni, 2005; Butler, 1997).

41 Lock-in-depth affects the nature of the paleomagnetic recording process in sediments. It is  
42 defined as the depth at which the remanent magnetization is stabilized. Lithology, grain-size  
43 distribution of the sediment matrix, sedimentation rate and bioturbation, all have influence on  
44 the position of the lock-in-depth in the sediments (Sagnotti et al., 2005; Bleil and von  
45 Dobeneck, 1999). When assuming the steady sedimentation rate, the result of lock-in-depth  
46 stabilization is younger magnetization than the sediment itself by an amount of time required  
47 to accumulate sediment layer of thickness that equal to the lock-in-depth. For example, if the  
48 sediment has accumulation speed 1 mm per 1000 years, and lock-in-depth is 10 mm, the  
49 magnetization age is 10 000 years younger than the sediment itself (Sagnotti et al., 2005).

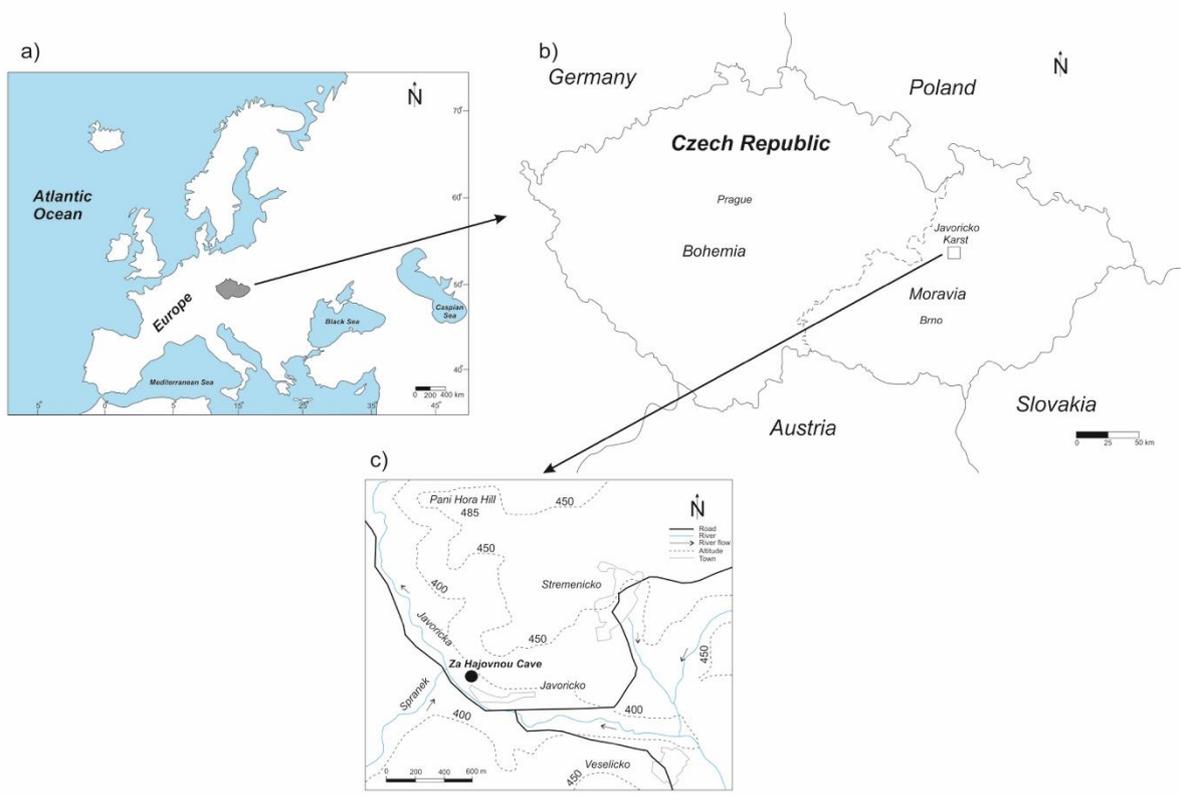
50 Kadlec et al. (2005, 2014) reported that the Central European cave (local name “Za  
51 Hajovnou”), in the Moravia region of the Czech Republic, has a potential record of  
52 Matuyama-Brunhes transition. Here we obtained a new paleomagnetic dataset from three  
53 vertical sediment profiles found in this cave.

#### 54 *1.1 Geology and Sampling*

55 The Za Háčovnou Cave (49° 40' N, 16° 55' E) is a former sinkhole located in Javoricko Karst,  
56 Moravia Region of the Czech Republic (Lundberg et al., 2014; Musil, 2014) (Figure 1). The  
57 Javoricko Karst is formed by light-grey-coloured massive Devonian limestone that overlies  
58 Pre-Cambrian phyllite (Lundberg et al., 2014; Musil, 2014). Spranek and Javoricka are two  
59 rivers that flow through the Jarovicko karst. While Za Háčovnou Cave is situated on the north-  
60 western bank of the Javoříčka river on the southern slope of a Pani Hora hill (Lundberg et al.,  
61 2014; Musil, 2014; Zak et al., 2018;), both Spranek and Javoricka watershed may have  
62 contributed to the sediment development in this cave (Figure 1).

63 The Za Hajovnou cave is approximately 500 m long system (Babek et al., 2015; Musil, 2005).  
64 The cave's corridors were explored previously in a total length of ~200 m (Musil, 2014)  
65 (Figure 2). The cave currently consists of two main parallel corridors with slightly different  
66 sedimentological record (Musil et al., 2014); the first corridor (local name is “Excavated  
67 Corridor”) used to be sinkhole entrance) and the other corridor (local name is “Birthday  
68 Corridor”) has separate entrance and is connected with the Excavated Corridor by the  
69 Connecting Passage Corridor (Figure 2). Sediments from the Excavated Corridor continue to  
70 Birthday Corridor and partially filled the Connecting Passage Corridor (Musil et al., 2014)  
71 (Figure 2).

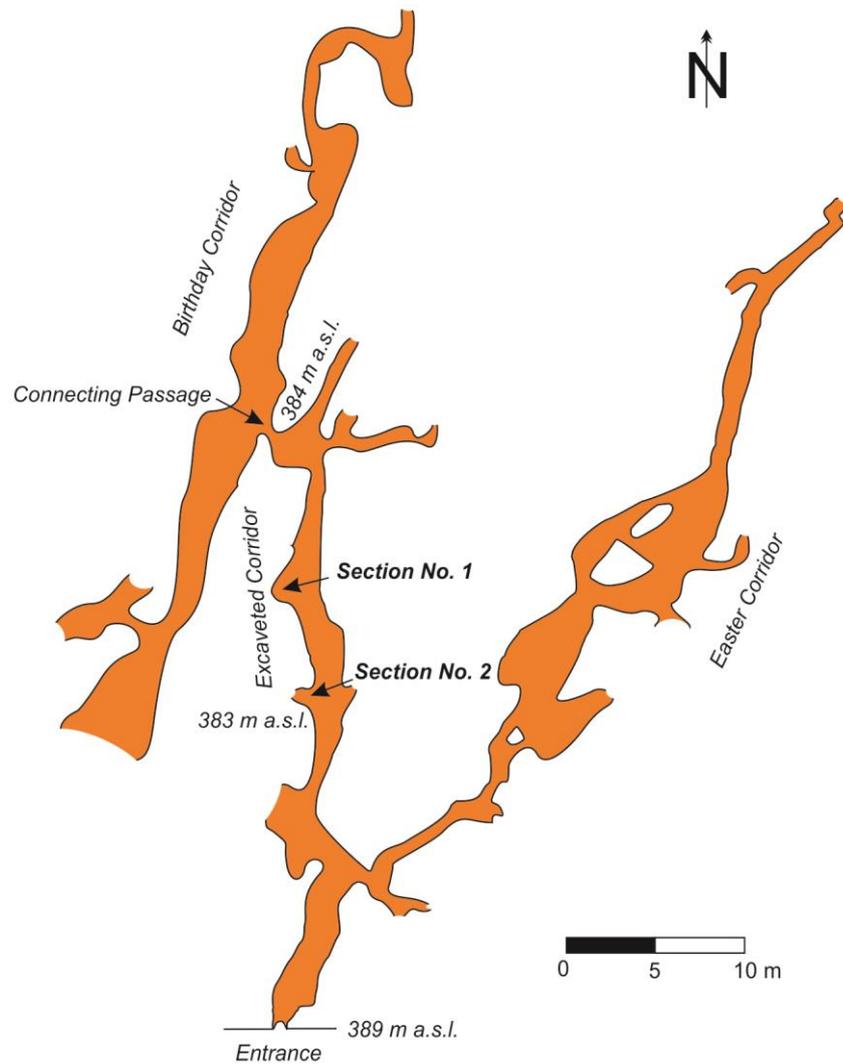
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73

74 **Figure 1:** Location of the study area in a) Central Europe. Dashed lines in b) show more  
 75 detailed placement different regions of the Czech Republic in relation to study area. Map in c)  
 76 shows regional detail of the Za Hajovnou cave placement (modified after Lundberg et al.,  
 77 2014; Musil, 2014).

78



79

80 **Figure 2:** Map of the Za Hajovnou cave (modified after Musil, 2014; Lundberg et al., 2014;  
 81 Kadlec et al., 2014). Locations marked by “Section No. 1” and “Section No. 2” is discussed in  
 82 the text (Kadlec et al., 2005, 2014). Map shows relation between Connecting Passage  
 83 Corridor, Birthday Corridor, and Excavated Corridor (m a.s.l.: meter above sea level).

84

85 Upper sediments of the cave were dated by U/Th dating of flowstones to  $118 \pm 1$  to  $267 \pm 3$   
 86 ka and by inferred Matuyama-Brunhes boundary in Section No. 1 (Figure 3) (Musil, 2005;  
 87 Kadlec et al., 2005, 2014; Lundberg et al., 2014; Babek et al 2015). The sedimentation in the  
 88 cave corridors were active probably from the Early Pleistocene until the beginning of the

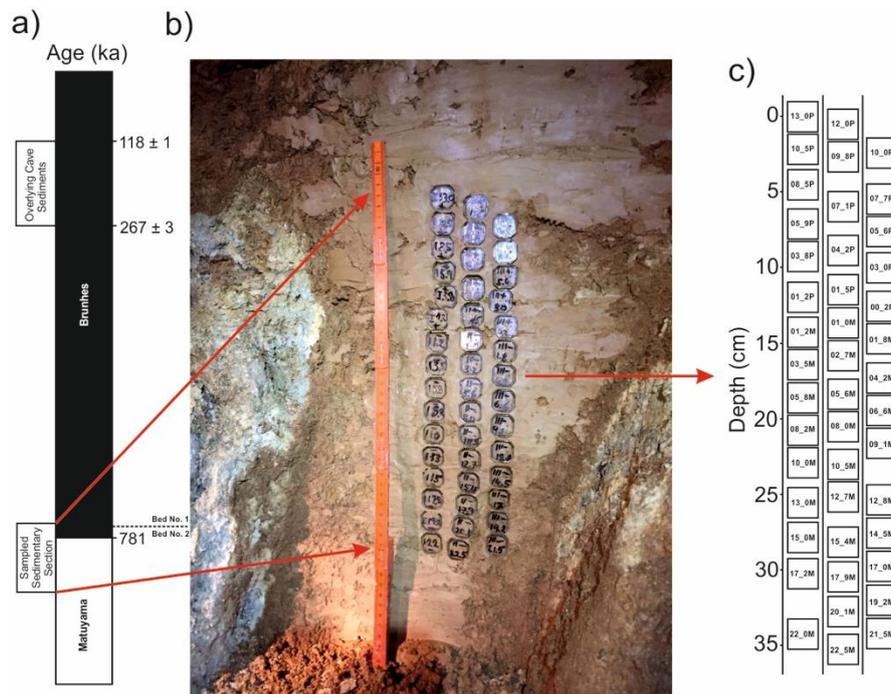
89 Middle Pleistocene. The sediment then consists of Pleistocene glacial time in north-western  
90 Europe called Cromerian Interglacials complex (Muller, 1992), MIS19 (marine isotope stage)  
91 which is interglacial period ~780 ka (Pol et al., 2019) and Matuyama-Brunhes reversal  
92 (Kadlec et al. 2005, 2014; Musil et al., 2014; Musil, 2014; Zak et al., 2018; Lundberg et al.,  
93 2014).

94 The Matuyama-Brunhes boundary (781 ka) was suggested in the upper part of the backwater  
95 fine sediments, deposited from suspension in the flooded cave. These sediments underlay the  
96 mostly non-fluvial deposits which entered the cave through a steep passage and filled the  
97 Connecting Passage Corridor. (Kadlec et al., 2014; Lundberg et al., 2014; Musil et al., 2014).

98 Sedimentary sections retrieved in the Za Hajovnou cave by Kadlec et al. (2005, 2014) were  
99 composed of two parts. The first part (Section No. 1, in Figure 2) was situated in the  
100 Excavated Corridor about 28 m from the cave entrance (Kadlec et al., 2005) (Figure 2). It was  
101 interpreted to contain the magnetic transition from a reversed to a normal polarity and inferred  
102 from the age dates of the overlying non-fluvial sediments, that it could be the Matuyama-  
103 Brunhes reversal (Kadlec et al., 2005). The second sedimentary section (Section No. 2)  
104 partially overlapped the Section No. 1, and was located in the Excavated Corridor (Kadlec et  
105 al., 2014) (Figure 2). Kadlec et al. (2014) indicated that this section had sediment with just  
106 reversed polarity except upper part of the sediment where the magnetization was difficult to  
107 interpret because the sediments had weak magnetization for which the sensitivity of the Agico  
108 JR-6A spinner magnetometer was insufficient. Section No. 2 underlies the Section No. 1 and  
109 contained older backwater sediment with reversed magnetic polarity (age > 781 ka, (Kadlec et  
110 al., 2014)).

111 The difficulties in interpretation of the primary study by Kadlec et al. (2014) was the  
112 motivation for the presented research. Here we collected 44 oriented discrete sedimentary

113 samples from the Excavated Corridor near the upper backwater sedimentary Section No. 1  
114 (Figure 2, 3).



115  
116 **Figure 3:** a) Age diagram of Za Hajovnou cave (dashed lines show the boundary between  
117 Bed No. 1 and 2 b) Sampled sedimentary Section No. 1 and c) discrete samples (numbers  
118 show the sample name).

119  
120 **2. Materials and Methods**

121 *2.1. Preparation of the Samples*

122 The samples were placed in plastic boxes (2x2x2 cm; 8cc) for paleomagnetism studies. The  
123 thickness of the sampled part of the sedimentary section was 35.1 cm. Lithology of the  
124 sampled sedimentary section; The upper backwater fine sediment part was brown clayey silt  
125 with white angular clasts of weathered limestone and bone fragments (Bed No. 1) (Kadlec et  
126 al., 2014). Lower part of the section consisted of the brown silty clay without white clasts  
127 (Bed No. 2) which is presented in Figure 3 (Kadlec et al., 2014).

128 *2.2. Demagnetization Measurements*

129 To clean the secondary magnetizations from the sedimentary samples, we applied a stepwise  
130 alternative field (AF) demagnetization method in Pruhonice Paleomagnetism Laboratory of  
131 Czech Academy of Sciences. This method was carried out using a 2G Enterprises Cryogenic  
132 Magnetometer on 44 samples divided into 3 different sequences. The first sequence was 17  
133 samples (13\_0P, 10\_5P, 09\_8P, 08\_5P, 05\_9P, 03\_8P, 01\_2P, 01\_2M, 03\_5M, 05\_8M,  
134 08\_2M, 10\_0M, 13\_0M, 15\_0M, 17\_2M, 19\_2M, 22\_0M) where we demagnetized at 1 mT  
135 intervals between 0-49 mT and 10 mT intervals between 50-100 mT. The second sequence  
136 was 14 samples (12\_0P, 07\_1P, 04\_2P, 01\_5P, 01\_0M, 02\_7M, 05\_6M, 08\_0M, 10\_5M,  
137 12\_7M, 15\_4M, 17\_9M, 20\_1M, 22\_5M) where we demagnetized at 2 mT intervals between  
138 0-48 mT and 10 mT intervals between 50-100 mT. The third sequence was 13 samples  
139 (10\_0P, 07\_7M, 05\_6P, 03\_0P, 00\_2P, 01\_8M, 04\_2M, 06\_2M, 09\_1M, 12\_8M, 14\_5M,  
140 17\_0M, 21\_5M) and we demagnetized at 0.5 mT intervals between 0-39.5 mT and 10 mT  
141 intervals between 40-100 mT. Demagnetization data was interpreted by using Remasoft  
142 software which was written by Agico Company (Chadima and Hroudá, 2009).

143 Characteristic remanent magnetization (ChRM) directions and maximum angular deviation  
144 (MAD) values were determined from principal component analysis (PCA) (Kirschvink 1980)  
145 on Zijdeveld diagram (Zijdeveld, 1967). Virtual geomagnetic pole's (VGP's) latitudes and  
146 longitudes were calculated using PMGSC software by Randy Enkin. Table 1 shows the data  
147 of alternative field demagnetization for each individual sample. Two examples of alternative  
148 field demagnetization method for Matuyama and Brunhes sections of the samples with  
149 changes of declination, inclination angles and remanent magnetization intensity step by step  
150 are given in Figure 4. Rest of the samples are in the supplementary table S1 and table S2  
151 containing supplementary Figures. Maximum angular deviation (MAD) changes of this study  
152 and comparisons with previous studies (Okada et al., 2017; Sagnotti et al., 2014) are shown in  
153 Figure 5.

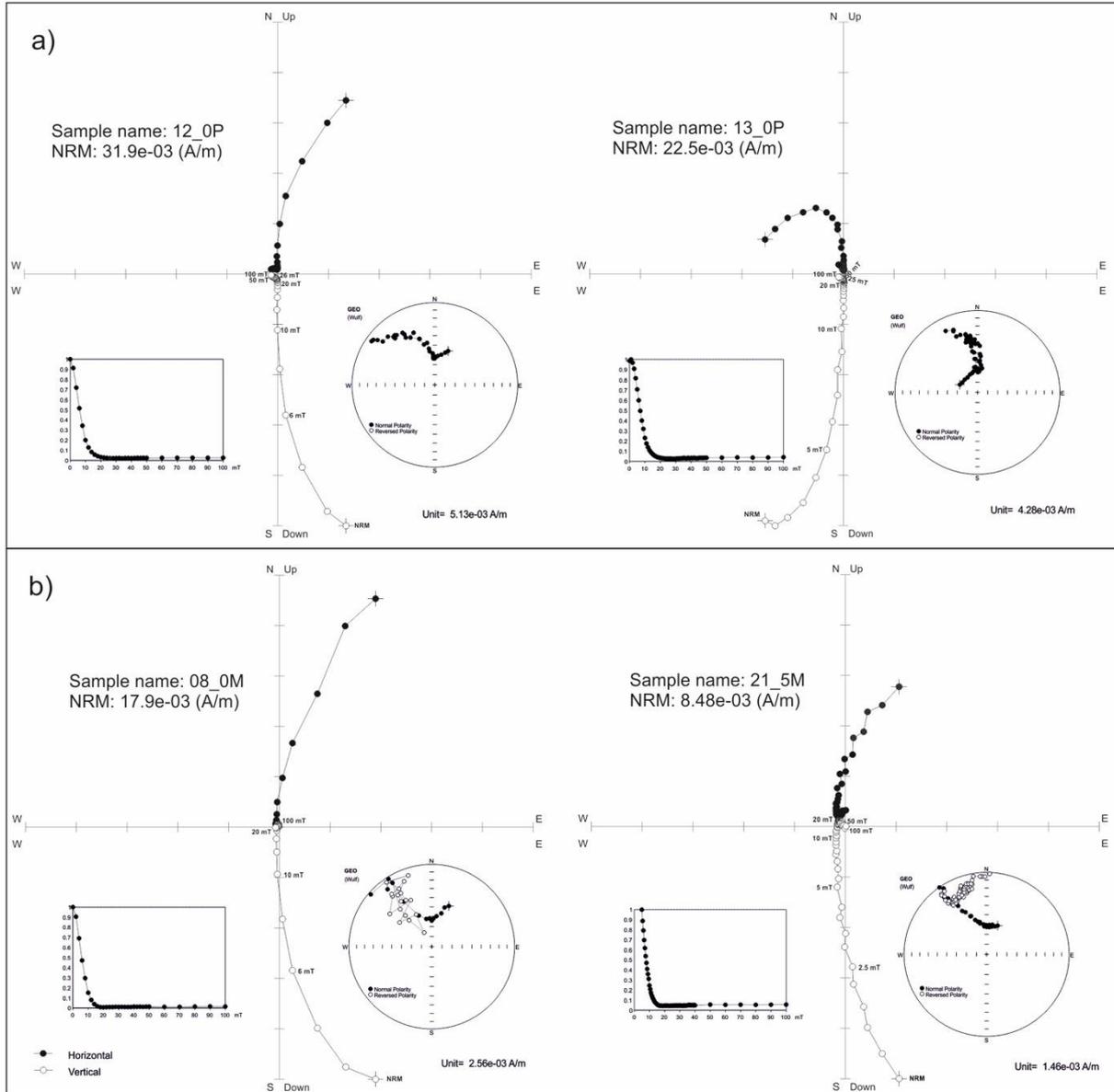
154 **Table 1:** Alternative field demagnetization, virtual geomagnetic pole (VGP) data and  
 155 Matuyama-Brunhes magnetic reversal scale for this study. Minus (-) values for virtual  
 156 geomagnetic pole (VGP) latitudes and longitudes indicate southern and western hemisphere.  
 157 MAD: maximum angular deviation,  $\phi_p$ : VGP latitude,  $\lambda_p$ : VGP longitude.

Sample Name	Depth (cm)	Dec <sup>(o)</sup>	Inc <sup>(o)</sup>	Intensity (A/m)	MAD <sup>(o)</sup>	$\phi_p$ (N <sup>o</sup> /S <sup>o</sup> )	$\lambda_p$ (E <sup>o</sup> /W <sup>o</sup> )	
13_0P	0	369.2	65.9	7.56E-04	0.4	83.8	114.5	Brunhes
12_0P	-1	356.1	58.5	1.06E-03	1.8	79.4	-146.7	
10.5P	-2.5	379.7	37.8	8.03E-04	1.4	57.8	160.5	
10_0P	-3	407.1	44.3	7.96E-04	1.2	47	121.8	
09_8P	-3.2	401.1	33	1.04E-03	1	44.5	135.4	
08_5P	-4.5	378.5	43.2	7.02E-04	1.6	61.8	159.2	
07_7P	-5.3	378.7	25.2	7.47E-04	1.4	50.7	167.1	
07_1P	-5.9	356.1	37.3	9.54E-04	1.4	61.3	-155.8	
05_9P	-7.1	342.5	57.7	4.15E-04	1.4	73.3	-108.3	Transition
05_6P	-7.4	369.1	30.4	4.48E-04	1.1	56.1	-179.3	
04_2P	-8.8	338.7	2.1	1.01E-03	3.8	38.3	-135.9	
03_8P	-9.2	308.2	-8.4	7.45E-04	2.3	-20.2	73.2	
03_0P	-10	327.2	-32.7	1.26E-03	0.7	-16.8	49.1	
01_5P	-11.5	339.3	-15.8	6.87E-04	5.3	-29.8	40.3	
01_2P	-11.8	338.1	3.7	6.12E-04	2.6	38.9	-134.8	
00_2P	-12.8	345.3	-38.7	7.03E-04	1.3	-17.6	30.9	
01_0M	-13.6	369.6	-6.3	8.70E-04	1.4	-36.8	4.5	Matuyama
01_2M	-13.8	371.8	-21.3	5.04E-04	3.8	-28.7	3.3	

01_8M	-14.4	315.8	-49.2	1.55E-03	1.2	-1.3	53.7
02_7M	-15.3	320.5	-22.8	8.30E-04	4.2	-19.6	57.9
03_5M	-16.1	341.7	-33.3	5.45E-04	0.6	-20.5	35.1
04_2M	-16.8	305.2	-40.3	7.78E-04	1.1	-2.8	65.4
05_6M	-18.2	305.9	-32.6	5.18E-04	1.2	-7.6	67.7
05_8M	-18.4	362.3	-61	1.06E-03	1	-1.5	-165.2
06_6M	-19.2	282.8	-65.3	3.19E-04	3.5	-27.5	-165.4
08_0M	-20.6	320.1	-40.3	1.25E-04	5.4	-9.4	53.3
08_2M	-20.8	335.3	-31.6	1.37E-03	0.3	-20	41.7
09_1M	-21.7	357.1	-58.1	1.40E-04	4.2	-1.8	18.8
10_0M	-22.6	352.5	-66.2	1.27E-03	1.5	-8.2	-158.4
10_5M	-23.1	393.7	-88.7	1.72E-03	1.4	-47.2	-165.6
12_7M	-25.3	102.8	-83.7	8.05E-04	3	-50.6	177.2
12_8M	-25.4	230.5	-72.3	8.44E-04	1.6	-59.6	-108.3
13_0M	-25.6	299.2	-56.6	1.04E-03	0.5	-11.9	-118.2
14.5M	-27.1	238	-68.7	9.90E-04	0.9	-54.2	-100.5
15_M	-27.6	312	-40.3	7.71E-04	2.5	-6	60
15_4M	-28	180.1	-89.3	8.56E-04	0.4	-50.8	-163.4
17_0M	-29.6	234.5	-79.1	9.03E-04	0.5	-57.6	-130.3
17_2M	-29.8	294.7	-60.7	2.11E-03	0.5	-16.9	-117.6
17_9M	-30.5	316.8	-82.8	1.56E-03	0.7	-38.3	-151.1
19_2M	-31.8	291.6	-87.8	4.43E-04	1.3	-47.5	-157.1
20_1M	-32.7	256.8	-72	1.24E-03	1.5	-45.9	-113.8
21_5M	-34.1	323.3	-23.4	3.91E-04	3.2	-20.5	55.1

22_0M	-34.6	250.8	-61.9	2.40E-03	0.3	-42.5	-94.3
22_5M	-35.1	143.3	-78.8	1.58E-03	0.4	-63.9	166.6

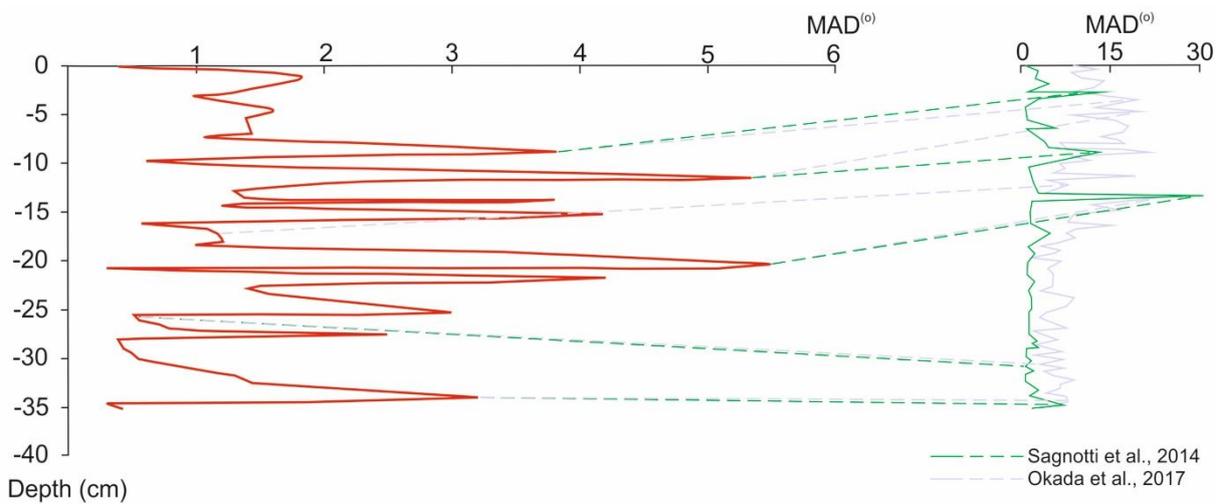
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159

160 **Figure 4:** Changes of magnetization directions on Zijderveld diagram and Wulf stereonet  
 161 during alternative field (AF) demagnetization method and demagnetization curve for typical  
 162 samples (See Supplementary Figures (Table S2) for all other samples); a) normal polarity  
 163 from Brunhes section (12\_0P, 13\_0P) and b) reversed polarity from Matuyama section  
 164 (08\_0M, 21\_5M).

165  
166



167

168 **Figure 5:** Comparison of Maximum angular deviation (MAD) changes from this data with  
169 published studies. during Matuyama-Brunhes magnetic reversal. Maximum angular deviation  
170 (MAD) values show error and confidence limit for this data during magnetic reversal, b:  
171 Published MAD values for datasets from Okada et al. 2017; Sagnotti et al. 2014. Colourful  
172 dashed lines connect sections published studies and this study where we see presentation of  
173 similar variation.

174

### 175 3. Results

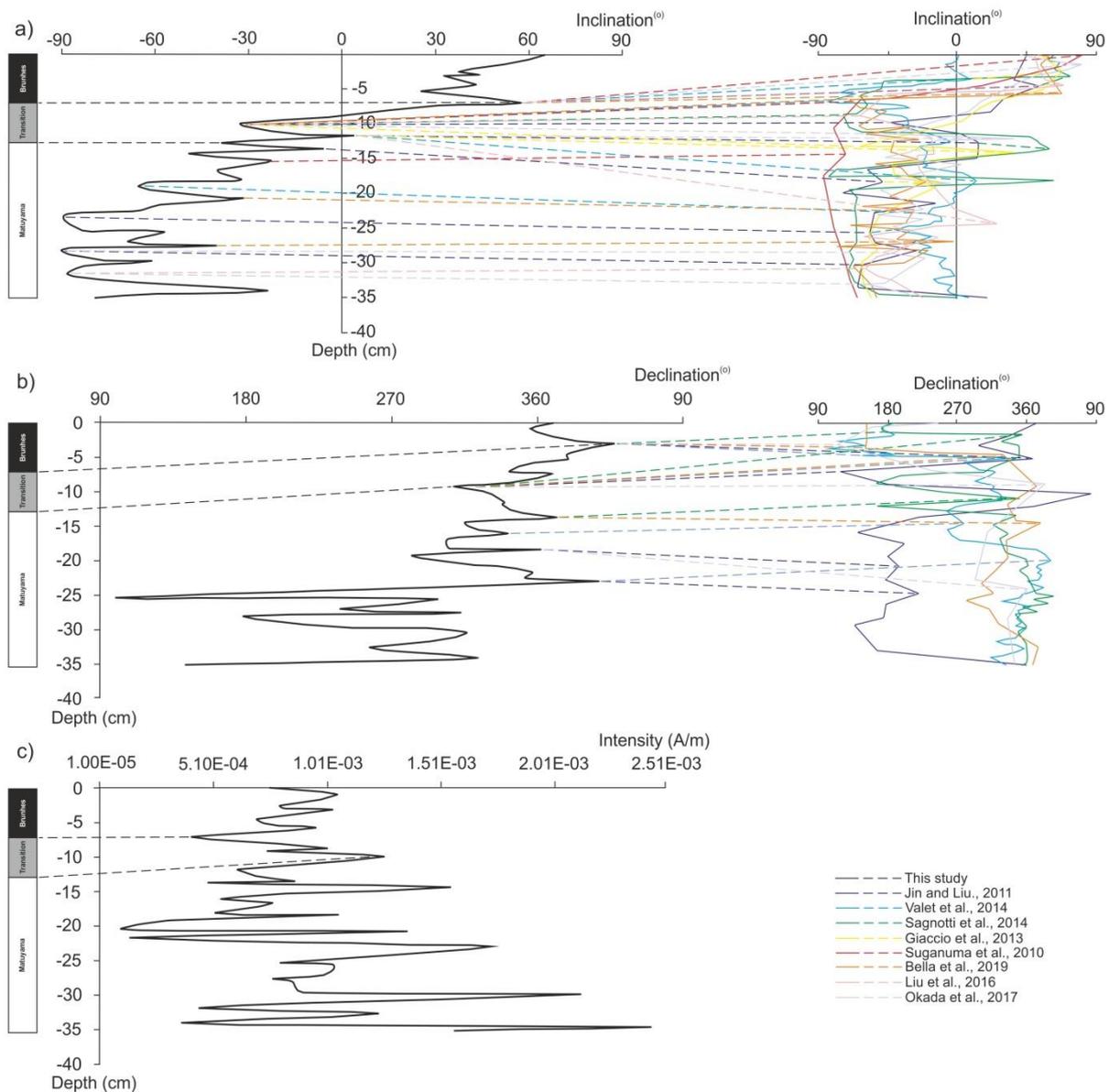
#### 176 3.1. Paleomagnetic Results

177 Sedimentary samples were demagnetized generally up to 20 mT (for details see Table S2)  
178 which removed the viscous remanent magnetization component causing change in the  
179 direction of remanent magnetization during such demagnetization for most of the samples.  
180 Intensity of the natural remanent magnetization (NRM) of the samples varies between 8.5-  
181 34.1e-3 A/m. Median destructive field (MDF) values where samples lost half of its  
182 magnetization range between 5-8 mT for the samples. NRM intensity and MDF values of the  
183 samples are shown in supplementary figures (Table S3). Maximum angular deviation (MAD)

184 values for Matuyama and Brunhes sections are between  $0.3^{\circ}$ - $5.4^{\circ}$  (Figure 5). These values for  
185 transition section are between  $0.7^{\circ}$ - $5.3^{\circ}$  which is relatively reliable for detection of the  
186 migration of the paleomagnetic vector from a reversed to normal polarity (Figure 5).  
187 Comparisons of MAD values with previous studies (Okada et al., 2017; Sagnotti et al., 2014)  
188 shows the relative magnitude of fluctuations.

189 In this study, paleomagnetic data showed inclination values changing by approximately  $90^{\circ}$   
190 when measuring the sediment from 12.8 to 7.1 cm depth (Figure 6). This revealed the  
191 transition nature of the Matuyama-Brunhes magnetic reversal in Za Hajovnou cave. Below  
192 this depth, there is a Matuyama section which has inclination fluctuations between  $-6.3^{\circ}$ - $88.7^{\circ}$   
193 (Figure 6). Inclination angle changes between  $33^{\circ}$ - $65.9^{\circ}$  for Brunhes section above transition  
194 (Figure 6). Also, transition from reversed to normal polarity can be seen in declination data  
195 with similar depth (Figure 6). Despite the fluctuations, magnetic intensity values which can  
196 depend on the concentration variation of magnetic carriers of every individual sample, were  
197 decreasing for the Matuyama section from the bottom to the transition (Figure 6). After the  
198 transition from reversed to normal polarity, these values kept increasing which can be seen in  
199 Brunhes section between 7.1-0 cm depth (Figure 6). Figure 6 shows the data in comparison  
200 with other studies that consisted of various sediment types and locations around the world.  
201 Depth of the data sets was normalized considering the transition zone and differences of  
202 sedimentation rate for each study and is not given in Figure 5, 6. Even though, there are some  
203 differences in absolute values, comparisons of this data set with other studies showed that  
204 fluctuations and frequency of fluctuations in our data is consistent with other data sets and  
205 serves as a supporting argument for Matuyama-Brunhes magnetic reversal in Za Hajovnou  
206 cave (Figure 6).

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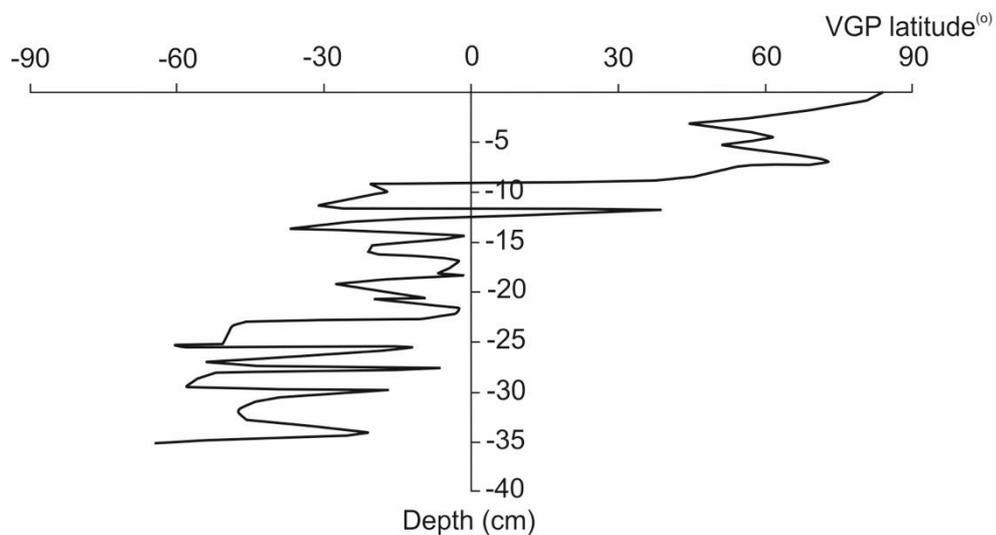


208  
 209 **Figure 6:** Comparisons of inclination and declination data with previous studies. As a result  
 210 of alternative field demagnetization method, data shows a) Inclination, b) Declination and c)  
 211 Magnetic intensity of the samples. Dashed lines in color connect sections published studies  
 212 and this study where we see presentation of similar variation.

213  
 214 *3.2. VGP's and Pole Migration*

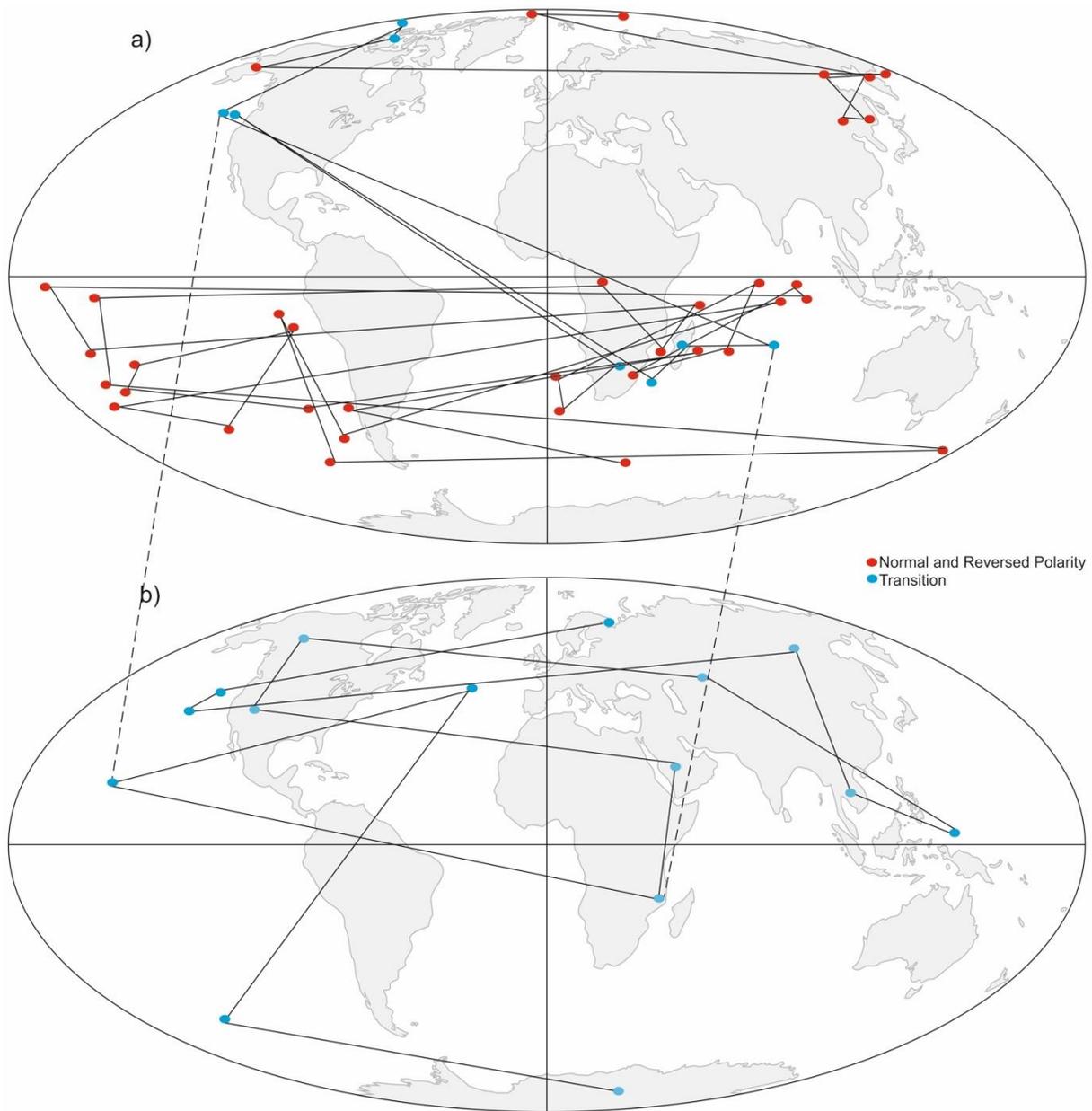
215 Virtual geomagnetic pole (VGP) shows the position of the geomagnetic paleopole (Lanza and  
 216 Meloni, 2006). Virtual geomagnetic pole's (VGP) latitudes from this data set show fluctuations

217 in Matuyama section (Figure 7). These values indicate 90° change from Matuyama to  
218 Brunhes transition as a result of pole migration (Figure 7). In addition, we plotted VGP path  
219 for Matuyama, Brunhes and transition sections using VGP latitudes and longitudes based on  
220 characteristic remanent magnetization (ChRM) directions detected in our samples (Figure 8).  
221 VGP locations for Matuyama section is in the southern hemisphere (Figure 8). During the  
222 transition from reversed to normal polarity, magnetic pole migrates from southern to northern  
223 hemisphere (Figure 8). After the geomagnetic transition, paleopoles fluctuate around  
224 geographic north pole (Figure 8). This VGP path of pole migration during the transition from  
225 the southern to northern hemisphere compares well with the Matuyama-Brunhes transition  
226 found by Okada et al. (2017) recorded in marine sediments near Japan (Figure 8).  
227



228  
229 **Figure 7:** Virtual geomagnetic pole (VGP) latitudes for this study.

230



231

232 **Figure 8:** a) Virtual geomagnetic pole (VGP) path of this study b) VGP path of transition  
 233 section from Okada et al. 2017 (dashed lines show similarities between two study).

234

### 235 3.3. Sedimentation Rate

236 Sedimentation rate of the sediment from this study in the Za Hajovnou cave is not known. We  
 237 compared the thickness of transition section of our study (cm) with the thickness of transition  
 238 section of other studies (cm) and estimated the sedimentation rate. Let  $n$  be the ratio between

239 the two data sets which shows how thick or thin the transition section of previous study  
 240 compared to our study (Equation 1.1, 1.2). This, we can estimate the sedimentation rate of Za  
 241 Hajovnou (Equation 1.2). In equations,  $t_{so}$  is the transition section thickness from our study  
 242 (in cm),  $t_{sp}$  is transition section thickness of published study (in cm),  $s_{rp}$  is sedimentation rate  
 243 of published study (in cm/kyr),  $s_{ro}$  is sedimentation rate of our study (in cm/kyr).

244 Equations;

$$t_{so} (cm) \times n = t_{sp} (cm) \quad (1.1)$$

$$s_{rp} (cm/kyr) / n = s_{ro} (cm/kyr) \quad (1.2)$$

245

246 Then, Za Hajovnou sedimentation rate ranges between 0.07-1.40 cm/kyr. Dropping the  
 247 highest (1.41 cm/kyr) and lowest (0.066 cm/kyr) values gives rate between (0.2 cm/kyr – 1.0  
 248 cm/kyr) Sedimentation rate estimates are shown in Table 2 compared with other studies.

249

250 **Table 2:** Sedimentation rate estimations for Za Hajovnou from previous studies.

251 \*: Sedimentation rate estimation for Za Hajovnou.

Reference	Sedimentation Rate	Sedimentation Rate*
Jin and Liu. (2011)	15 cm/kyr	0.6 cm/kyr 253
Okada et al. (2017)	61 cm/kyr	0.91 cm/kyr 254
Sagnotti et al. (2010)	61 cm/kyr	0.6 cm/kyr 255
Giaccio et al. (2013)	26 cm/kyr	1 cm/kyr 256
Sagnotti et al. (2014)	0.0224 cm/kyr	0.235 cm/kyr 257
Suganuma et al. (2010)	0.66 cm/kyr	0.23 cm/kyr 258
Bella et al. (2019)	0.64 cm/kyr	1.41 cm/kyr 259
Liu et al. (2016)	8.78 cm/kyr	0.066 cm/kyr 260

#### 261 **4. Discussion**

262 Our data indicate that Matuyama Brunhes transition boundary constitutes 5.7 cm, between  
263 7.1-12.8 cm depth of the sampled sedimentary section, of the Za Hajovnou cave sediment.  
264 Magnetic reversal is characterized by frequent fluctuations of inclination angle (Figure 6a)  
265 and VGP latitude (Figure 7). Maximum angular deviation (MAD) values of this study are  
266 within the error limit as seen in Figure 5. Similarities in comparisons of this data set with  
267 other studies indicate that Za Hajovnou cave sediment dates to Matuyama-Brunhes magnetic  
268 transition.

269 Because of low coercivity in most of the samples with demagnetization generally at 20 mT,  
270 some large fluctuations in the data may be considered as instability of remanent  
271 magnetization. This shows that minerals with low coercivity is responsible for the  
272 magnetization of the cave sediments in our study. On the other hand, similar fluctuations  
273 which are seen in previous studies (Figure 6) show reliability of the data.

274 Although the data in this study and Okada et al. 2017 belong to geographically different  
275 locations and sediment types, this similarity during polar migration shows that the reversal  
276 was a dipole transition, and non-dipole field component was less significant (Simon et al.  
277 2019; Mochizuki et al. 2011; Oda et al., 2000).

278 Our estimate of the Za Hajovnou cave's sedimentation rate seems to be significantly lower  
279 than the sediments from other studies (Jin and Liu., 2011; Okada et al., 2017; Sagnotti et al.,  
280 2010; Sagnotti et al., 2014; Giaccio et al., 2013; Suganuma et al., 2010; Bella et al., 2019; Liu  
281 et al., 2016). This is likely due to contrasting sediment types.

282 Analyzes of cave sediments by paleomagnetism carried out in different locations around the  
283 earth such as in Western Europe (Pares et al., 2018), South Africa (Nami et al., 2016), South  
284 America (Jaqueto et al., 2016), North America (Stock et al., 2005), Southern Europe (Pruner  
285 et al., 2010), Eastern Asia (Morinaga et al., 1992) showed that cave sediments recorded

286 magnetic reversals. Morinaga et al. (1992) suggested low sedimentation rate for the Western  
287 Japan cave sediments 1.6 cm/kyr which shows similar rate with our Central European cave  
288 sediment estimation. King and Channel (1991) suggested that large "lock-in" depths are  
289 associated with interparticle rigidity and strength, characteristic of clayey low accumulation  
290 rate sediments (<1 cm/kyr) which results in delays of magnetic acquisition. This shows that  
291 magnetic polarity reversal could have a large (25 kyr) apparent age offset between sediments  
292 with high and very low accumulation rates (King and Channel, 1991).

293 Glass and Heezen (1967) claimed that a meteorite impact could result in a magnetic reversal.  
294 Large meteorite impacts may provide the sufficient moment to the exterior of the outer core to  
295 cause motion relative to the outer core. In this way, the angular difference disrupts the  
296 convection and the position of the magnetic poles. Changing in the convection pattern would  
297 affect the electric currents in the liquid core to have a new core convection with modified  
298 Coriolis forces. In addition the inner core's angular momentum change due to the new core  
299 convection dynamics, in respect to the mantle, would change the heat transfer in the inner  
300 core and have an effect on dynamo reset. This process may cause variations in the currents  
301 producing the earth's magnetic field and cause a geomagnetic reversal. In our study, the  
302 inclination and VGP change which have an anomaly about 12 cm depth just before the  
303 reversal (Figure 6, 7) could be a reason of a meteorite impact which was shown in a study in  
304 Indonesia from marine sediments (Hyodo et al., 2011) by micro-tektite level formed due to  
305 cosmic impacts. This can be supported with oscillations in VGP's latitude approximately at  
306 the same depth (1.5 m) below the reversal in the study of Yamazaki and Oda (2001) in South  
307 Atlantic. Both of the studies (Hyodo et al., 2001; Yamazaki and Oda, 2001) have very similar  
308 sedimentation rates between 8-10 cm/kyr. The same anomaly can be seen in sediments from  
309 other studies (Sagnotti et al., 2014; Giaccio et al., 2013; Valet et al., 2014; Jin and Liu, 2011;  
310 Liu et al., 2016; Okada et al., 2017) (Figure 6). Migration of the pole during Matuyama

311 polarity is towards to South America (Figure 8) which may be caused by the angular  
312 momentum and convection pattern change because of the northward directed meteorite impact  
313 in Asia (Sieh et al., 2020).

## 314 **5. Conclusions**

315 We compared the paleomagnetic data from the cave sediments with published magnetic  
316 reversal record and were able for the first time to use the detailed magnetic characteristic of  
317 cave sediment and to infer the specific magnetic reversal (Matuyama-Brunhes). This is  
318 possible due to nature of the magnetic reversal that we discovered. Note that the paleopole  
319 was residing in the east of Africa and than quickly reappeared west of North America. We  
320 consider this as an important marker signature for dating the central European paleomagnetic  
321 record from this time period.

322 Additionally, we were able to design a new method for estimation of the accumulation rate of  
323 the cave sediment in Za Hajovnou cave, which until now was not known. Also we provided  
324 an evidence that a meteorite impact could be the cause of the Matuyama-Brunhes magnetic  
325 reversal.

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## 331 **Data Availability**

332 Data is currently uploaded as Supplementary Information.

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487 **Additional Information**

488 The authors declare no competing interests.