

1 **Title: Evidence of Matuyama-Brunhes transition in the cave sediment in Central Europe**

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6 Abbreviated title: Matuyama-Brunhes record from the European cave sediment

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9 **Summary**

10 In this study, we identified the Matuyama-Brunhes magnetic reversal recorded in cave
11 sediments in Central Europe, Czech Republic. We collected discrete samples from the
12 homogeneous sedimentary profile in the Za Hajovnou cave located in the eastern part of the
13 Czech Republic. Novel use of characteristic remanent magnetization (ChRM) directions and
14 VGP (Virtual Geomagnetic Pole) path of the data revealed the Matuyama-Brunhes transition
15 boundary within 5.7 cm of the Za Hajovnou cave sediment. This result revealed a new more
16 detailed behavior of the polarity transition from the central European location. The migration
17 of the paleopole between east of Africa and west of North America is a significant marker for
18 the central European paleomagnetic record in terms of global magnetic data. The precursor of
19 the reversal occurs 4.6 kyr before the transition. As a result of the rock magnetism
20 measurements, the magnetic carrier of the samples is maghemite. Also, we estimated the
21 sedimentation rate of the cave for the first time.

22 **Keywords:** Paleomagnetism; Matuyama-Brunhes, magnetic reversal; cave; sediments

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26 **1. Introduction**

27 Matuyama-Brunhes magnetic reversal occurred approximately 781 kyr ago (Gradstein et al.
28 2004). Published studies (Channel et al., 2010; Sagnotti et al., 2010, 2014; Sugauma et al.,
29 2010; Jin and Liu, 2011; Giaccio et al., 2013; Kitaba et al., 2013; Pares et al., 2013; Valet et
30 al., 2014; Liu et al., 2016; Okada et al., 2017; Bella et al., 2019) showed that this event is well
31 recorded by respective sediments that had sufficient sedimentation rate and could be analyzed,
32 in detail, by paleomagnetism. Also, studies in the recent years have shown a younger age for
33 the reversal around 773 kyr (Channell et al., 2010 (773 ± 0.4 kyr); Sugauma et al., 2015
34 (770.2 ± 7.3 kyr); Simon et al., 2019 (773.9 kyr); Singer et al., 2019 (773 ± 2 kyr); Valet et
35 al., 2019 (772.4 ± 6.6 kyr); Haneda et al., 2020 (772.9 ± 5.4 kyr)).

36 Sediments acquire remanent magnetization during their deposition. The alignment of
37 magnetic moments of the grains occurs in the direction of the earth's magnetic field and
38 acquisition of primary magnetization due to this sedimentation process is called depositional
39 or detrital remanent magnetization (DRM) (Gubbins and Herrero, 2017). Remanent
40 magnetization protected by potential energy barriers can last over geologic time scales.
41 Nevertheless, due to thermal and/or chemical processes such as reheating, oxidation and
42 formation of iron hydroxides during the time, secondary magnetizations can be acquired by
43 crossing potential energy barriers or generation of chemical remanences. The new secondary
44 magnetization has an orientation in the direction of the Earth's field. Then rocks can acquire a
45 viscous remanent magnetization (VRM) a long time after their formation due to exposure to
46 the geomagnetic field. VRM contributes to noise in paleomagnetic data (Butler, 1992; Lanza
47 and Meloni, 2006).

48 Lock-in-depth affects the nature of the paleomagnetic recording process in sediments. It is
49 defined as the depth at which the remanent magnetization is stabilized. Lithology, grain-size
50 distribution of the sediment matrix, sedimentation rate and bioturbation, all have an influence

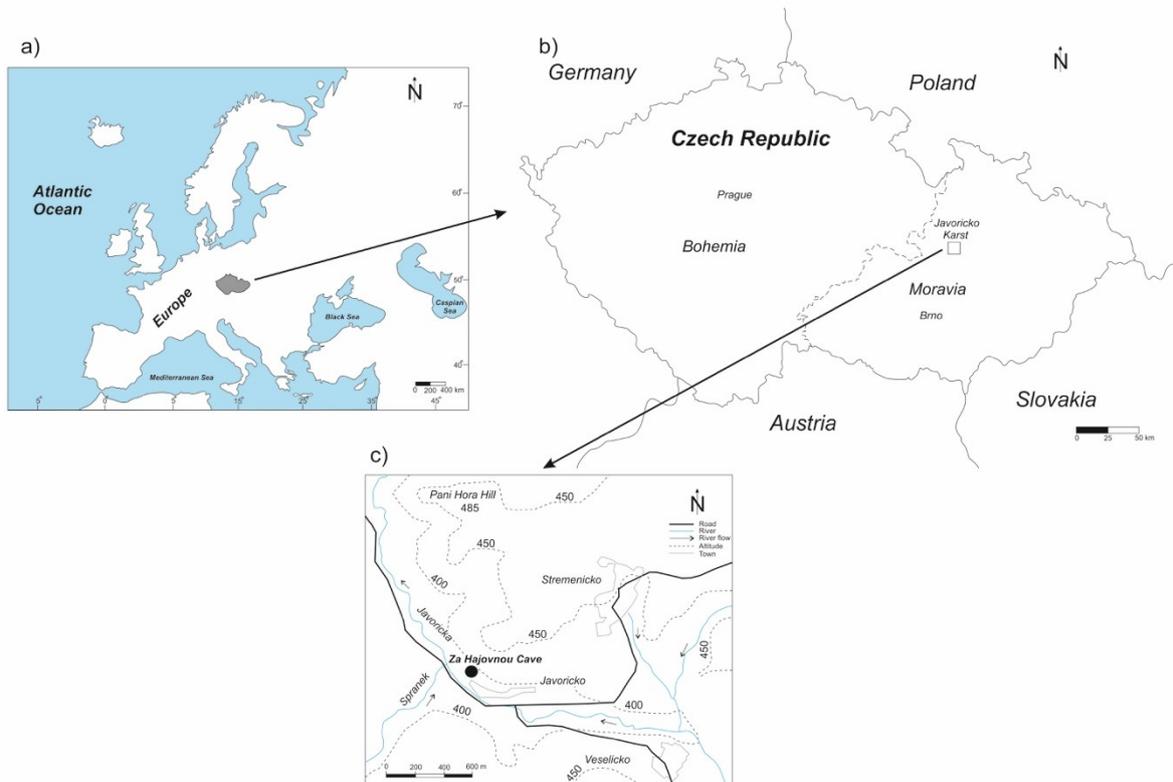
51 on the position of the lock-in-depth in the sediments (Bleil and von Dobeneck, 1999; Sagnotti
52 et al., 2005). When assuming the steady sedimentation rate, the result of lock-in-depth
53 stabilization is younger magnetization than the sediment itself by an amount of time required
54 to accumulate a sediment layer of thickness that equal to the lock-in-depth. For example, if
55 the sediment has an accumulation speed of 1 mm per 1000 years, and lock-in-depth is 10 mm,
56 the magnetization age is 10 000 years younger than the sediment itself (Sagnotti et al., 2005).
57 Kadlec et al. (2005, 2014) reported that the Central European cave (local name “Za
58 Hajovnou”) in the Moravia region of the Czech Republic records the Matuyama-Brunhes
59 transition. The aim of our study is to analyze the reversal in detail paleomagnetic method and
60 to identify the magnetic carrier of the cave sediment. Here, we obtained a new paleomagnetic
61 dataset from three vertical sediment profiles found in this cave. A central European
62 paleomagnetic record of the B/M boundary will be valuable for investigation of the
63 characteristic behavior of the Earth’s magnetic field during Matuyama-Brunhes magnetic
64 reversal. The sedimentation rate of the cave sediment is not known yet in terms of
65 understanding the transition. It makes our estimation even more crucial in this study.

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67 *1.1 Geology of the Cave*

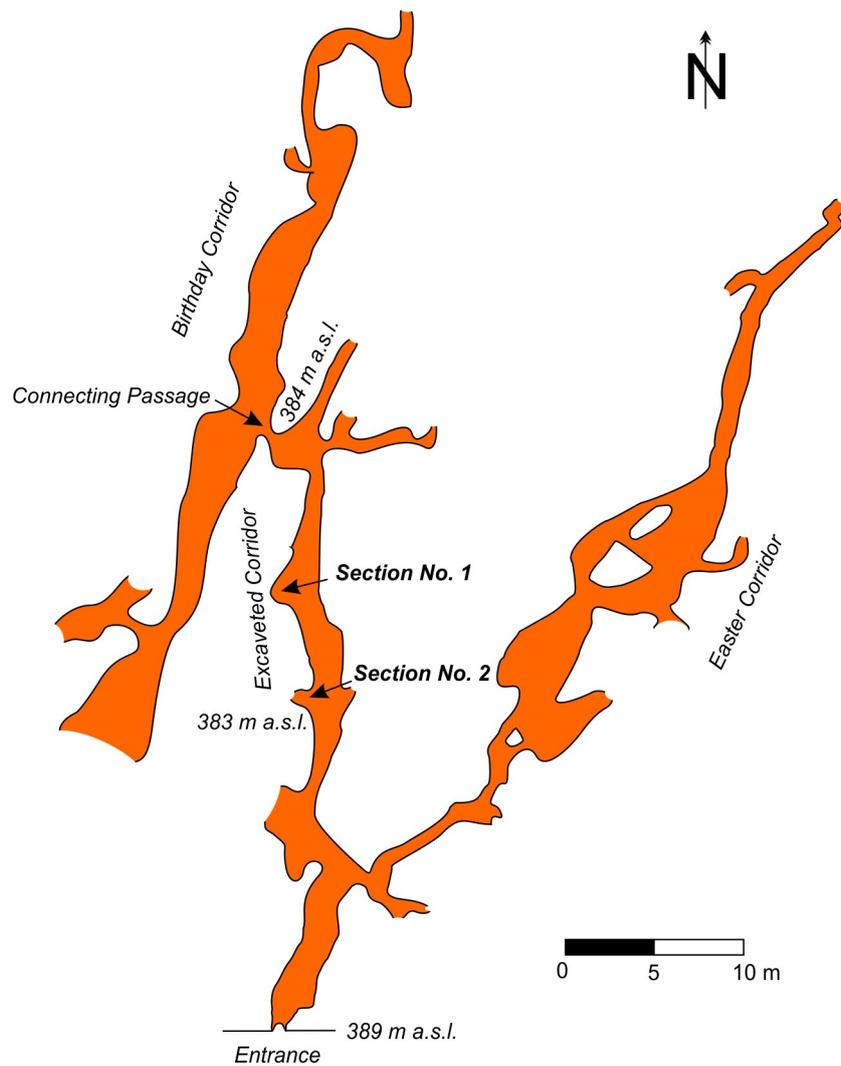
68 The Za Hájojnou Cave (49° 40' N, 16° 55' E) is a former sinkhole located in Javoricko Karst,
69 Moravia Region of the Czech Republic (Lundberg et al., 2014; Musil, 2014) (Fig. 1). The
70 Javoricko Karst is formed by light-grey-colored massive Devonian limestone that overlies
71 Pre-Cambrian phyllite (Lundberg et al., 2014; Musil, 2014). Spranek and Javoricka are two
72 rivers that flow through the Jarovicko karst. While Za Hájojnou Cave is situated on the north-
73 western bank of the Javoříčka river on the southern slope of a Pani Hora hill (Lundberg et al.,
74 2014; Žák et al., 2018), both Spranek and Javoricka watershed may have contributed to the
75 sediment development in this cave (Fig. 1).

76 The Za Hajovnou cave is approximately 500 m long system (Musil, 2005; Bábek et al., 2015).
 77 The cave's corridors were explored previously in a total length of ~200 m (Musil, 2014) (Fig.
 78 2). The cave currently consists of two main parallel corridors with a slightly different
 79 sedimentological record (Musil et al., 2014); the first corridor (local name is “Excavated
 80 Corridor” which is used to be sinkhole entrance) and the other corridor (local name is
 81 “Birthday Corridor”) has a separate entrance. These two corridors are connected by the
 82 Connecting Passage Corridor (Fig. 2). Sediments from the Excavated Corridor continue to
 83 Birthday Corridor and partially filled the Connecting Passage Corridor (Musil et al., 2014)
 84 (Fig. 2).



85
 86 **Fig. 1.** Location of the study area in (a) Central Europe. Dashed lines in (b) show more
 87 detailed placement with different regions of the Czech Republic in relation to the study area.
 88 Map in (c) shows regional detail of the Za Hajovnou cave placement (modified after
 89 Lundberg et al., 2014; Musil, 2014).

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92 **Fig. 2.** Map of the Za Hajovnou cave (modified after Kadlec et al., 2014; Lundberg et al.,
 93 2014; Musil, 2014). Locations marked by “Section No. 1” and “Section No. 2” are discussed
 94 in the text (Kadlec et al., 2005, 2014). The map shows the relation between Connecting
 95 Passage Corridor, Birthday Corridor, and Excavated Corridor (m a.s.l.: meter above sea
 96 level).

97

98 Upper sediments of the cave were dated by U/Th dating of flowstones from 118 ± 1 to $267 \pm$
 99 3 ka and the sediment sequence below was inferred to contain the Matuyama-Brunhes
 100 boundary in Section No. 1 (Fig. 3) (Musil, 2005; Kadlec et al., 2005, 2014; Lundberg et al.,
 101 2014; Bábek et al., 2015). The sedimentation in the cave corridors was active from the Early

102 Pleistocene until the beginning of the Middle Pleistocene. The sediment then consists of
103 Pleistocene glacial time in north-western Europe called the Cromerian Interglacials complex
104 (Muller, 1992), MIS19 (marine isotope stage) which is interglacial period ~780 ka (Pol *et al.*,
105 2019), and Matuyama-Brunhes reversal (Kadlec *et al.*, 2005, 2014; Lundberg *et al.*, 2014;
106 Musil, 2014; Musil *et al.*, 2014; Žák *et al.*, 2018).

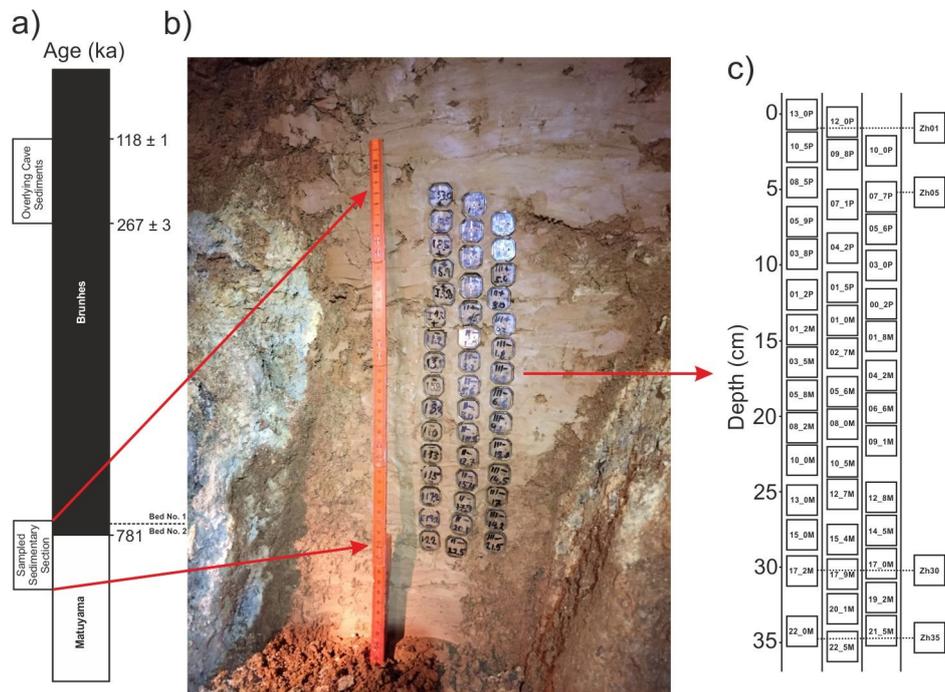
107 The Matuyama-Brunhes boundary (781 ka) was suggested in the upper part of the backwater
108 fine sediments, deposited from suspension in the flooded cave. These sediments underlay
109 mostly non-fluvial deposits which entered the cave through a steep passage and filled the
110 Connecting Passage Corridor (Kadlec *et al.*, 2014; Lundberg *et al.*, 2014; Musil *et al.*, 2014).

111 The total thickness of the sediments, deposited from suspension, reaches up to 4.3 m.

112 Sedimentary sections retrieved in the Za Hajovnou cave by Kadlec *et al.* (2005, 2014) were
113 composed of two parts. The first part (Section No. 1, in Fig. 2) was situated in the Excavated
114 Corridor about 28 m from the cave entrance (Kadlec *et al.*, 2005) (Fig. 2). It was interpreted to
115 contain the magnetic transition from a reversed to normal polarity and inferred from the age
116 dates of the overlying non-fluvial sediments, that our data indicate that the reversal is the
117 actual Matuyama-Brunhes reversal. Sediment thickness in the Section No. 1 is 0.8 m. The
118 second sedimentary section (Section No. 2) partially overlapped the Section No. 1 and was
119 located in the Excavated Corridor (Kadlec *et al.*, 2014) (Fig. 2). Kadlec *et al.* (2014) indicated
120 that this section had sediment with just reversed polarity except for the upper part of the
121 sediment where the magnetization was difficult to interpret because the sediments had weak
122 magnetization for which the sensitivity of the Agico JR-5A spinner magnetometer was
123 insufficient. Section No. 2 underlies the Section No. 1 and contains older backwater sediment
124 with reversed magnetic polarity (age > 781 ka) (Kadlec *et al.*, 2014).

125 The difficulties in the interpretation of the primary study by Kadlec *et al.* (2014) was the
126 motivation for the presented research. Here we collected 44 oriented discrete sedimentary

127 samples from the Excavated Corridor near the upper backwater sedimentary Section No. 1
 128 (Fig. 2, 3).



129

130 **Fig. 3.** Za Hajovnou cave sediments. (a) Age diagram of Za Hajovnou cave (dashed lines
 131 show the boundary between Bed No. 1 and 2), (b) sampled sedimentary Section No. 1, and
 132 (c) discrete samples for the paleomagnetism measurements and the rock magnetism samples
 133 (numbers show the sample name).

134

135 2. Materials and Methods

136 2.1. Preparation of the Samples

137 The sediment outcrop in the cave was made into a smooth plane and the samples for the
 138 paleomagnetism measurements were taken by pushing the plastic boxes (2x2x2 cm; 8cc) into
 139 the sediment (Fig. 3). We used a common geological compass (produced by the Brunton) for
 140 measuring the azimuth and tilt of the sample boxes. After the paleomagnetism sampling,
 141 another 4 samples were placed in the plastic boxes for the rock magnetism measurements
 142 (Fig. 3c). These samples were taken from the same depth as the paleomagnetic samples

143 (13_0P, 7_7P, 17_2M, 22_0M). The thickness of the sampled part of the sedimentary section
144 was 35.1 cm. The upper backwater (water accumulation) fine sediment part was brown clayey
145 silt with white angular clasts of weathered limestone and bone fragments (Bed No. 1) (Kadlec
146 et al., 2014). The lower part of the section consisted of the brown silty clay without white
147 clasts (Bed No. 2) (Kadlec et al., 2014) which is presented in Fig. 3.

148

149 *2.2. Demagnetization Measurements*

150 To clean the secondary magnetizations from the sedimentary samples, we applied a stepwise
151 alternative field (AF) demagnetization method in Pruhonice Paleomagnetism Laboratory of
152 Czech Academy of Sciences. This method was carried out using a 2G Enterprises Cryogenic
153 Magnetometer on 44 samples divided into 3 different sequences. The first sequence was 17
154 samples (13_0P, 10_5P, 09_8P, 08_5P, 05_9P, 03_8P, 01_2P, 01_2M, 03_5M, 05_8M,
155 08_2M, 10_0M, 13_0M, 15_0M, 17_2M, 19_2M, 22_0M) where we demagnetized at 1 mT
156 intervals between 0-49 mT and 10 mT intervals between 50-100 mT. The second sequence
157 was 14 samples (12_0P, 07_1P, 04_2P, 01_5P, 01_0M, 02_7M, 05_6M, 08_0M, 10_5M,
158 12_7M, 15_4M, 17_9M, 20_1M, 22_5M) where we demagnetized at 2 mT intervals between
159 0-48 mT and 10 mT intervals between 50-100 mT. The third sequence was 13 samples
160 (10_0P, 07_7M, 05_6P, 03_0P, 00_2P, 01_8M, 04_2M, 06_2M, 09_1M, 12_8M, 14_5M,
161 17_0M, 21_5M) and we demagnetized at 0.5 mT intervals between 0-39.5 mT and 10 mT
162 intervals between 40-100 mT. Demagnetization data was interpreted by using Remasoft
163 software which was written by Agico Company (Martin Chadima and Frantisek Hrouda).

164 Characteristic remanent magnetization (ChRM) directions and maximum angular deviation
165 (MAD) values were determined from principal component analysis (PCA) (Kirschvink, 1980)
166 on the Zijderveld diagram (Zijderveld, 1967). Virtual geomagnetic pole's (VGP's) latitudes
167 and longitudes were calculated using PMGSC software by Randy Enkin. Table 1 shows the

168 data of alternative field demagnetization for each sample. Examples of alternative field
 169 demagnetization results for the Matuyama and Brunhes intervals (two examples each) are
 170 shown in Fig. 4. The rest of the samples are in the Supporting Information Tables S1 and
 171 Supporting Information Figs S2. Maximum angular deviation (MAD) values of this study and
 172 other studies (Sagnotti et al., 2014; Okada et al., 2017) are shown in Fig. 5.

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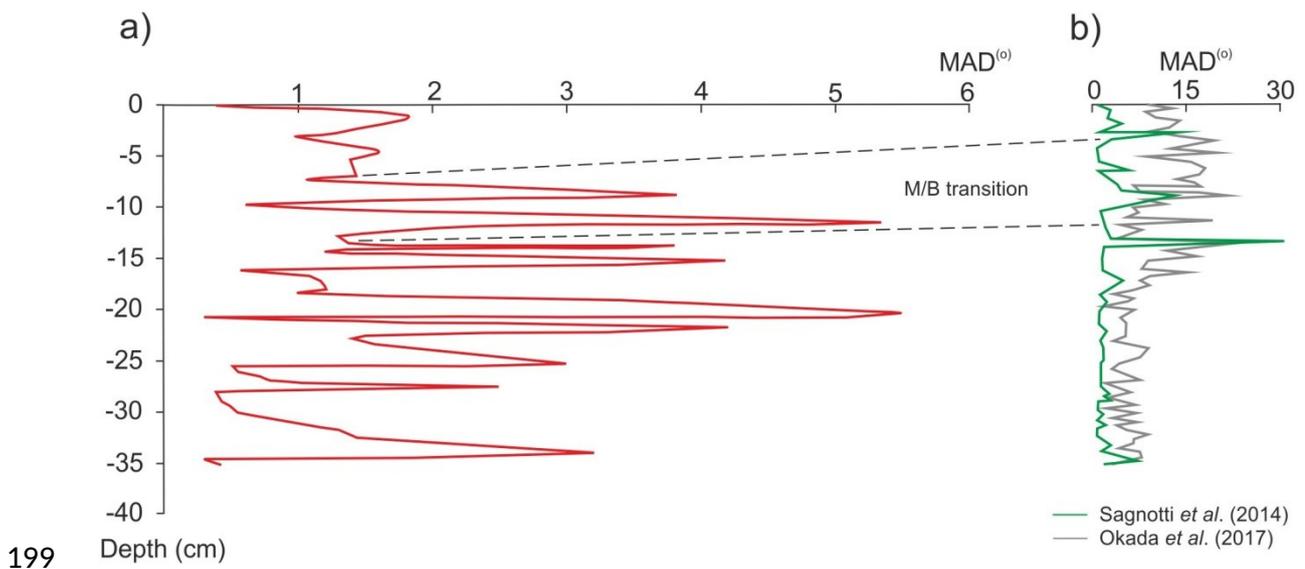
186 **Table 1.** Alternative field demagnetization, virtual geomagnetic pole (VGP) data, and
 187 Matuyama-Brunhes magnetic reversal scale for this study. Minus (-) values for virtual
 188 geomagnetic pole (VGP) latitudes and longitudes indicate the southern and western
 189 hemispheres. MAD: maximum angular deviation, ϕ_p : VGP latitude, λ_p : VGP longitude.

Sample Name	Depth (cm)	Dec ^(o)	Inc ^(o)	Intensity (A/ m)	MAD ^(o)	ϕ_p (N°/S°)	λ_p (E°/W°)	
13_0P	0	369.2	65.9	7.56E-04	0.4	83.8	114.5	Brun

12_0P	-1	356.1	58.5	1.06E-03	1.8	79.4	-146.7	hes
10.5P	-2.5	379.7	37.8	8.03E-04	1.4	57.8	160.5	
10_0P	-3	407.1	44.3	7.96E-04	1.2	47	121.8	
09_8P	-3.2	401.1	33	1.04E-03	1	44.5	135.4	
08_5P	-4.5	378.5	43.2	7.02E-04	1.6	61.8	159.2	
07_7P	-5.3	378.7	25.2	7.47E-04	1.4	50.7	167.1	
07_1P	-5.9	356.1	37.3	9.54E-04	1.4	61.3	-155.8	
05_9P	-7.1	342.5	57.7	4.15E-04	1.4	73.3	-108.3	Transition
05_6P	-7.4	369.1	30.4	4.48E-04	1.1	56.1	-179.3	
04_2P	-8.8	338.7	2.1	1.01E-03	3.8	38.3	-135.9	
03_8P	-9.2	308.2	-8.4	7.45E-04	2.3	-20.2	73.2	
03_0P	-10	327.2	-32.7	1.26E-03	0.7	-16.8	49.1	
01_5P	-11.5	339.3	-15.8	6.87E-04	5.3	-29.8	40.3	
01_2P	-11.8	338.1	3.7	6.12E-04	2.6	38.9	-134.8	
00_2P	-12.8	345.3	-38.7	7.03E-04	1.3	-17.6	30.9	
01_0M	-13.6	369.6	-6.3	8.70E-04	1.4	-36.8	4.5	Matuyama
01_2M	-13.8	371.8	-21.3	5.04E-04	3.8	-28.7	3.3	
01_8M	-14.4	315.8	-49.2	1.55E-03	1.2	-1.3	53.7	
02_7M	-15.3	320.5	-22.8	8.30E-04	4.2	-19.6	57.9	
03_5M	-16.1	341.7	-33.3	5.45E-04	0.6	-20.5	35.1	
04_2M	-16.8	305.2	-40.3	7.78E-04	1.1	-2.8	65.4	
05_6M	-18.2	305.9	-32.6	5.18E-04	1.2	-7.6	67.7	
05_8M	-18.4	362.3	-61	1.06E-03	1	-1.5	-165.2	
06_6M	-19.2	282.8	-65.3	3.19E-04	3.5	-27.5	-165.4	
08_0M	-20.6	320.1	-40.3	1.25E-04	5.4	-9.4	53.3	

08_2M	-20.8	335.3	-31.6	1.37E-03	0.3	-20	41.7
09_1M	-21.7	357.1	-58.1	1.40E-04	4.2	-1.8	18.8
10_0M	-22.6	352.5	-66.2	1.27E-03	1.5	-8.2	-158.4
10_5M	-23.1	393.7	-88.7	1.72E-03	1.4	-47.2	-165.6
12_7M	-25.3	102.8	-83.7	8.05E-04	3	-50.6	177.2
12_8M	-25.4	230.5	-72.3	8.44E-04	1.6	-59.6	-108.3
13_0M	-25.6	299.2	-56.6	1.04E-03	0.5	-11.9	-118.2
14.5M	-27.1	238	-68.7	9.90E-04	0.9	-54.2	-100.5
15_M	-27.6	312	-40.3	7.71E-04	2.5	-6	60
15_4M	-28	180.1	-89.3	8.56E-04	0.4	-50.8	-163.4
17_0M	-29.6	234.5	-79.1	9.03E-04	0.5	-57.6	-130.3
17_2M	-29.8	294.7	-60.7	2.11E-03	0.5	-16.9	-117.6
17_9M	-30.5	316.8	-82.8	1.56E-03	0.7	-38.3	-151.1
19_2M	-31.8	291.6	-87.8	4.43E-04	1.3	-47.5	-157.1
20_1M	-32.7	256.8	-72	1.24E-03	1.5	-45.9	-113.8
21_5M	-34.1	323.3	-23.4	3.91E-04	3.2	-20.5	55.1
22_0M	-34.6	250.8	-61.9	2.40E-03	0.3	-42.5	-94.3
22_5M	-35.1	143.3	-78.8	1.58E-03	0.4	-63.9	166.6

193 **Fig. 4.** Changes of magnetization directions on the Zijderveld diagram and Wulf stereonet
 194 during alternative field (AF) demagnetization method and demagnetization curve for typical
 195 samples (see Supporting Information Figs S2 for all other samples); (a) normal polarity from
 196 Brunhes section (12_0P, 13_0P) and (b) reversed polarity from Matuyama section (08_0M,
 197 21_5M).
 198



200 **Fig. 5.** Maximum angular deviation (MAD) changes during the Matuyama-Brunhes magnetic
 201 reversal a) from this data, b) from other published studies. Dashed lines show the M/B
 202 transition.

203

204 2.3. Rock Magnetism Measurements

205 To determine the magnetic minerals in the samples, High Temperature Magnetic
 206 Susceptibility measurements were done up to 635 °C by using an Agico Kappabridge MFK1-
 207 FA susceptibility meter. Samples used for the measurements are ZH01, ZH05, ZH30, ZH35
 208 which correspond to the paleomagnetic samples 13_0P, 7_7P, 17_2M, 22_0M respectively.
 209 The samples for the susceptibility measurements (χ mass normalized) were chosen as a
 210 result of the paleomagnetic data.

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212 **3. Results**

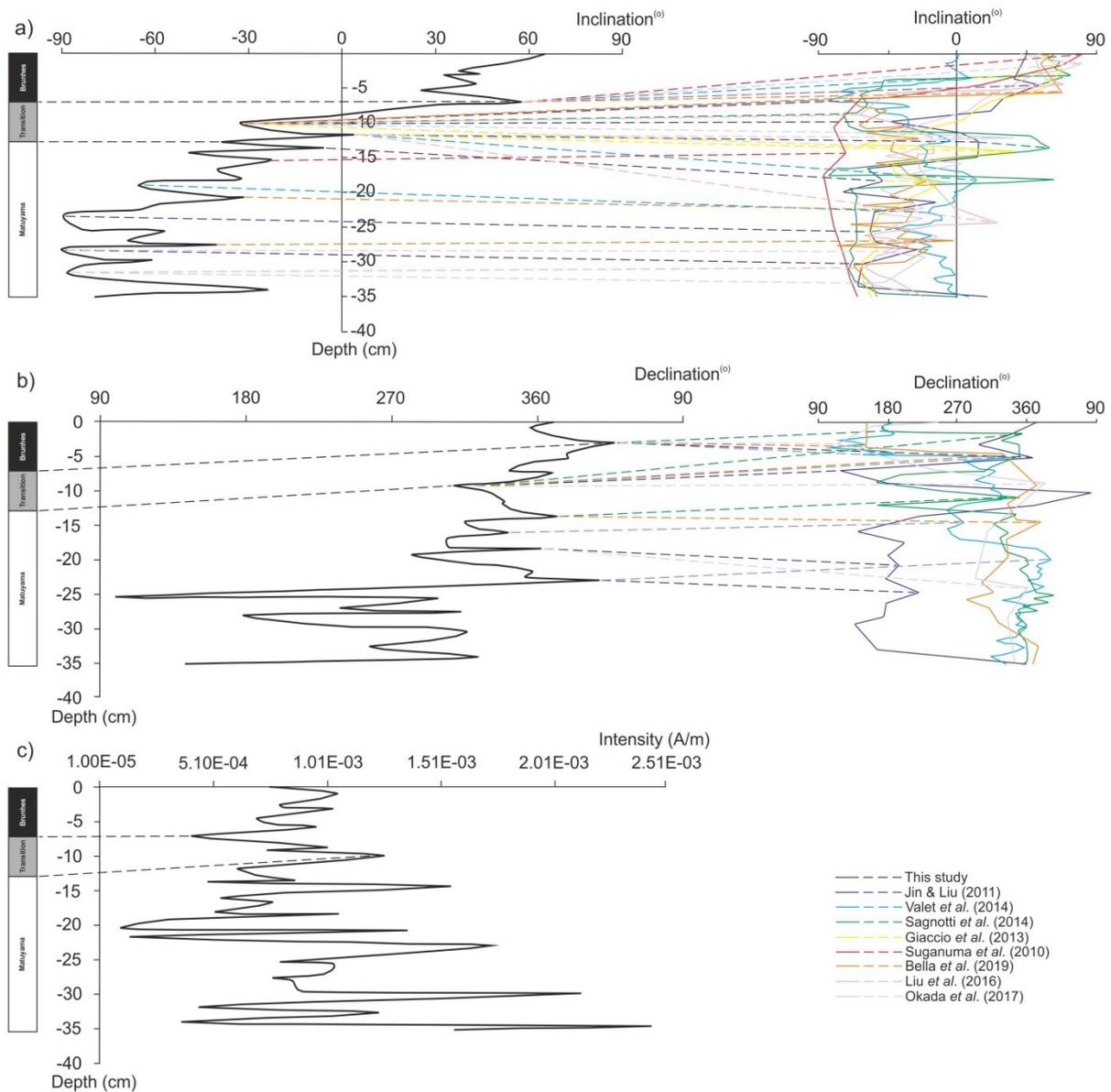
213 *3.1. Paleomagnetic Results*

214 The samples were demagnetized generally up to 20 mT (for details see Supporting
215 Information Figs S2) which removed the viscous remanent magnetization (VRM) component
216 causing a change in the direction of remanent magnetization during such demagnetization for
217 most of the samples. This soft VRM component in the samples has a mean D: 13.8° and I:
218 56.8° values which is close to the present day's Earth's magnetic field direction values for the
219 Czech Republic (D: 4.4° and I: 66.8°) (see Supporting Information Figs S3).

220 The intensity of the natural remanent magnetization (NRM) of the samples varies between
221 8.5-34.1e-3 A/m. Median destructive field (MDF) values where samples lost half of its
222 magnetization range between 5-8 mT for the samples. NRM intensity and MDF values of the
223 samples are shown in Supporting Information Figs S3. The maximum angular deviation
224 (MAD) values for Matuyama and Brunhes sections are between 0.3°-5.4° (Fig. 5a). These
225 values for the transition section are between 0.7°-5.3° which is relatively reliable for the
226 detection of the migration of the paleomagnetic vector from reversed to normal polarity (Fig.
227 5a). The trend of the MAD values increases across the transition which can also be seen in the
228 other studies (Sagnotti et al., 2014; Okada et al., 2017) (Fig. 5b).

229 In this study, paleomagnetic data showed inclination values changing by approximately 90°
230 when measuring the sediment from 12.8 to 7.1 cm depth (Fig. 6). This revealed the transition
231 nature of the Matuyama-Brunhes magnetic reversal in Za Hajovnou cave. Below this depth,
232 there is a Matuyama section which has inclination fluctuations between -6.3°-88.7° (Fig. 6).
233 Inclination angle changes between 33°-65.9° for Brunhes section above transition (Fig. 6).
234 Also, the transition from reversed to normal polarity can be seen in declination data with
235 similar depth. It has more frequent oscillations which show ~180° change from reversed to

236 normal polarity between 2.5-9.2 cm depth (Fig. 6b). Despite the fluctuations, the intensity
237 values of ChRM which can depend on the concentration variation of magnetic carriers of
238 every individual sample were decreasing for the Matuyama section from the bottom to the
239 transition between 35.1-15 cm depth (Fig. 6). After the transition from reversed to normal
240 polarity, these values kept increasing which can be seen in the Brunhes section between 7.1-0
241 cm depth (Fig. 6). Fig. 6 shows the data in comparison with other studies that consisted of
242 various sediment types and locations around the world. The depth of the data sets was
243 normalized considering the transition zone and differences of sedimentation rate for each
244 study and is not given in Fig. 5, 6. Even though there are some differences in absolute values,
245 comparisons of this data set with other studies showed that fluctuations and frequency of
246 fluctuations in our data are consistent with other data sets and serves as a supporting argument
247 for Matuyama-Brunhes magnetic reversal in Za Hajovnou cave (Fig. 6).



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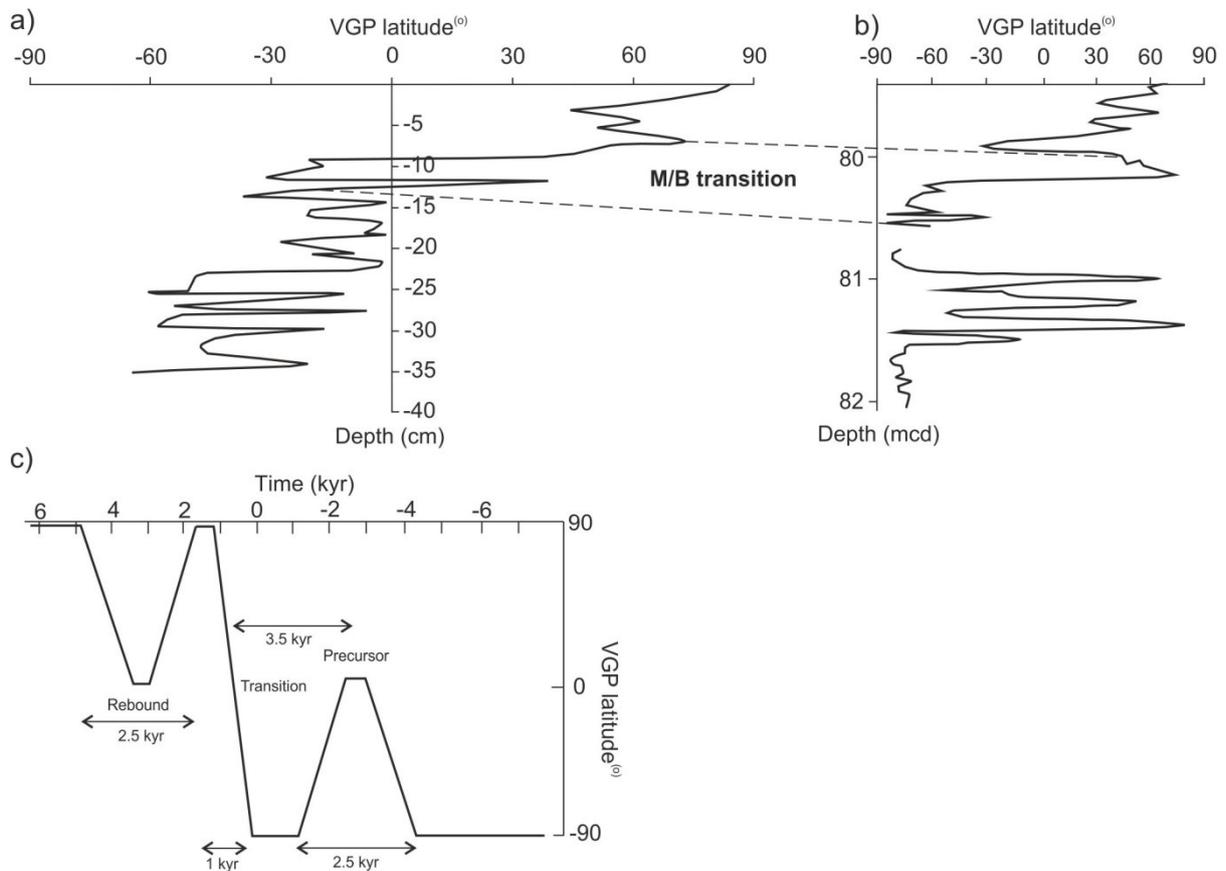
249 **Fig. 6.** Comparisons of inclination and declination data with previous studies. As a result of
 250 the alternative field demagnetization method, data shows (a) inclination, (b) declination, and
 251 (c) intensity of ChRM of the samples. Dashed lines in color connect sections of published
 252 studies and this study where we see the presentation of similar variation.

253

254 3.2. VGP's and Pole Migration

255 The virtual geomagnetic pole (VGP) shows the position of the geomagnetic paleopole (Lanza
 256 and Meloni, 2006). VGP's latitudes from this data set show fluctuations in the Matuyama

257 section which are similar with the data from Yamazaki and Oda (2001) (Fig. 7a, b). These
 258 values indicate 90° change from Matuyama to Brunhes transition as a result of pole migration
 259 (Fig. 7a). 75° change in VGP values at 11.8 cm depth (Fig. 7a) shows the precursor of the
 260 reversal according to Valet et al. (2012) (Fig. 7c). In addition, we plotted the VGP path for
 261 Matuyama, Brunhes, and transition sections using VGP latitudes and longitudes based on
 262 characteristic remanent magnetization (ChRM) directions detected in our samples (Fig. 8).
 263 VGP locations for the Matuyama section are in the southern hemisphere (Fig. 8). During the
 264 transition from reversed to normal polarity, the magnetic pole migrates from southern to
 265 northern hemisphere (Fig. 8). After the geomagnetic transition, paleopoles fluctuate around
 266 the geographic north pole (Fig. 8). This VGP path of pole migration during the transition from
 267 the southern to northern hemisphere compares well with the Matuyama-Brunhes transition
 268 found by Okada et al. (2017) recorded in marine sediments near Japan (Fig. 8).



269

270 **Fig. 7.** Virtual geomagnetic pole (VGP) latitudes of a) this study, b) Yamazaki and Oda
271 (2001), and c) shows the precursor model of Valet et al. (2012). Dashed lines show the main
272 M/B transition for a) and b).

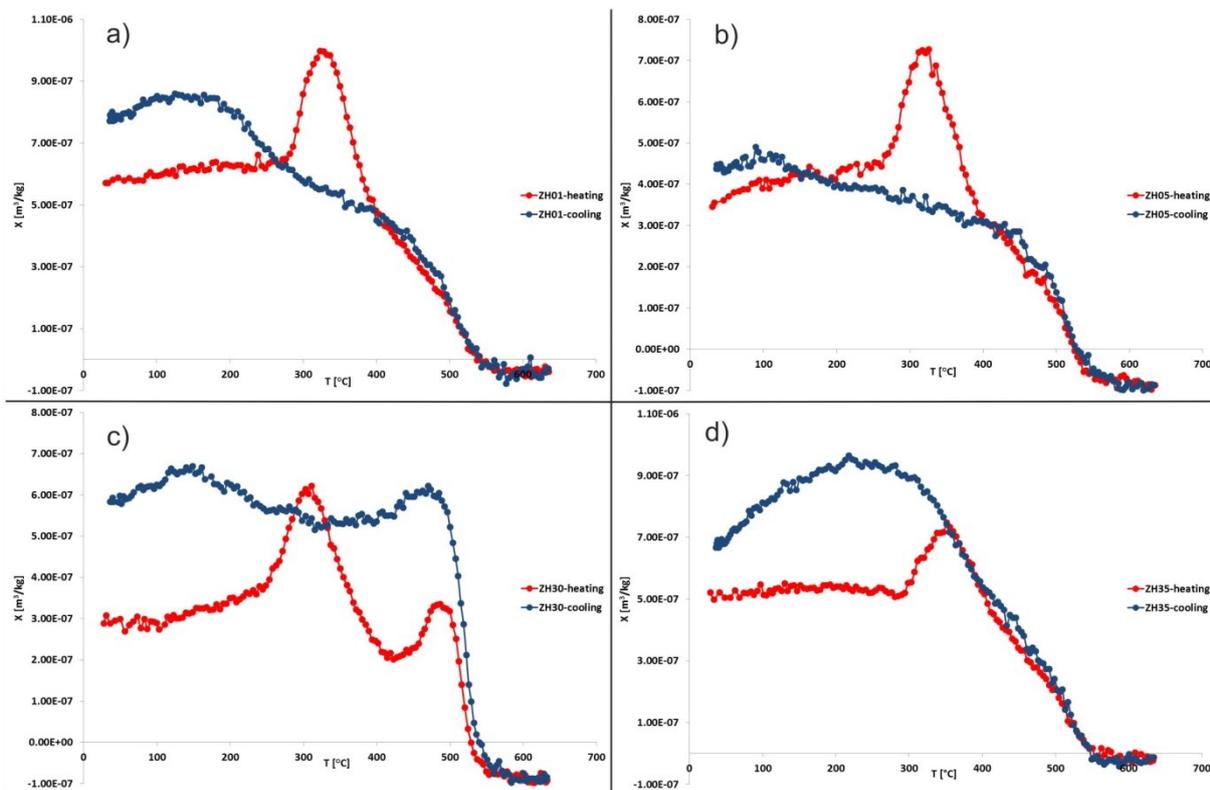


273
274 **Fig. 8.** Virtual geomagnetic pole (VGP) path of (a) this study and (b) VGP path of transition
275 section from Okada et al. (2017) (dashed lines show similarities between two study).

276

277 *3.3. Rock Magnetism Results*

278 According to the High Temperature Magnetic Susceptibility measurements, the transition of
 279 maghemite to magnetite can be seen with an increase in susceptibility values at approximately
 280 250-350 °C (Fig. 9). The samples have a Curie temperature between 530-550 °C which may
 281 be a sign of titanium in the minerals and of a new sulphite phase created from decomposing
 282 the maghemite and incorporation of the sulfur from the surrounding clay. The increase of
 283 susceptibility values in the cooling curves corresponds to the percentage of maghemite
 284 decrease after the heating in the samples (Fig. 9a, c, d). Two different drops in susceptibility
 285 values at 410 °C and 530 °C in Fig. 9c shows the existence of maghemite and magnetite
 286 together in the sample.



287
 288 **Fig. 9.** High Temperature Magnetic Susceptibility measurement results (χ : mass normalized
 289 magnetic susceptibility, T: temperature in Celcius).

290

291 *3.4. Sedimentation Rate Estimation*

292 The sedimentation rate of the sediment from this study in the Za Hajovnou cave is not known.
 293 We compared the thickness of the transition section of our study (cm) with the thickness of
 294 the transition section of other studies (cm) and estimated the sedimentation rate. Let n be the
 295 ratio between the two data sets which shows how thick or thin the transition section of the
 296 previous study compared to our study (Equation 1.1, 1.2). This, we can estimate the
 297 sedimentation rate of Za Hajovnou (Equation 1.2). In equations, tso is the transition section
 298 thickness from our study (in cm), tsp is the transition section thickness of the published study
 299 (in cm), srp is the sedimentation rate of the published study (in cm/kyr), sro is sedimentation
 300 rate of our study (in cm/kyr).

301 Equations;

302 $tso (cm) \times n = tsp (cm) \quad (1.1)$

303 $srp (cm/kyr) / n = sro (cm/kyr) \quad (1.2)$

304

305 Then, Za Hajovnou sedimentation rate ranges between 0.07-1.40 cm/kyr. Dropping the
 306 highest (1.41 cm/kyr) and lowest (0.066 cm/kyr) values gives rate between 0.2-1.0 cm/kyr.
 307 Thus, the average sedimentation rate is 0.6 cm/kyr. Sedimentation rate estimates compared
 308 with other studies are shown in Table 2.

309

310 **Table 2.** Sedimentation rate estimations for Za Hajovnou from the previous studies.

311 Reference Sedimentation Rate Sedimentation Rate* *: Sedimentation rate

Reference	Sedimentation Rate	Sedimentation Rate*	estimation for Za Hajovnou.
Jin and Liu (2011)	15 cm/kyr	0.6 cm/kyr	
Okada et al. (2017)	61 cm/kyr	0.91 cm/kyr	
Sagnotti et al. (2010)	61 cm/kyr	0.6 cm/kyr	
Giaccio et al. (2013)	26 cm/kyr	1 cm/kyr	
Sagnotti et al. (2014)	21 cm/kyr	0.235 cm/kyr	
Suganuma et al. (2010)	0.66 cm/kyr	0.23 cm/kyr	
Bella et al. (2019)	0.64 cm/kyr	1.41 cm/kyr	
Liu et al. (2016)	8.78 cm/kyr	0.066 cm/kyr	

317

318 **4. Discussion**

319 Our data indicate that the Matuyama-Brunhes transition boundary constitutes 5.7 cm, between
320 7.1-12.8 cm depth of the sampled sedimentary section, of the Za Hajovnou cave sediment.
321 The magnetic reversal is characterized and represented by frequent fluctuations of inclination
322 angle (Fig. 6a) and VGP latitude (Fig. 7a). We think that oscillations in declination data show
323 instability of the Earth's magnetic field. On the other hand, similar fluctuations which are
324 seen in the previous studies (Fig. 6) show the reliability of the data.

325 Although the data in this study and Okada et al. (2017) belong to geographically different
326 locations and sediment types, the similarity during polar migration (Fig. 8) shows that the
327 reversal was a dipole transition, and the non-dipole field component was less significant (Oda
328 et al., 2000; Mochizuki et al., 2011; Simon et al., 2019).

329 In most of the samples with demagnetization generally at 20 mT, some large fluctuations in
330 the data may be considered as instability of remanent magnetization. Although some samples
331 (01_8M, 04_2M, 17_2M, 17_9M, 22_0M) are not demagnetized up to 100 mT, it shows that
332 minerals with low coercivity are responsible for the magnetization of the cave sediments in
333 our study. It can be clearly seen in the rock magnetism measurements that the samples show
334 similar behaviour during the measurements that is maghemite in this case.

335 Note that most of the section contains samples from the polarity transition. The data shows
336 that the magnetic field was unstable for our oldest sample already, when in reversed polarity.
337 This observation goes well with Yamazaki and Oda (2001) where they show the magnetic
338 pole was unstable long time before the actual reversal boundary and the magnetic field started
339 to fluctuate almost 150 cm deeper than the actual transition (Fig. 7b, 10). We think that our
340 data illustrate the same instability, and this is why no paleomagnetic samples have VGP
341 latitudes that deviate less than 25° from the reversed position. We provide a more detailed

342 explanation of the reversed VGP behavior in our data that show reversed polarity unrest well
343 before the actual magnetic reversal.

344 *4.1. Precursor Event*

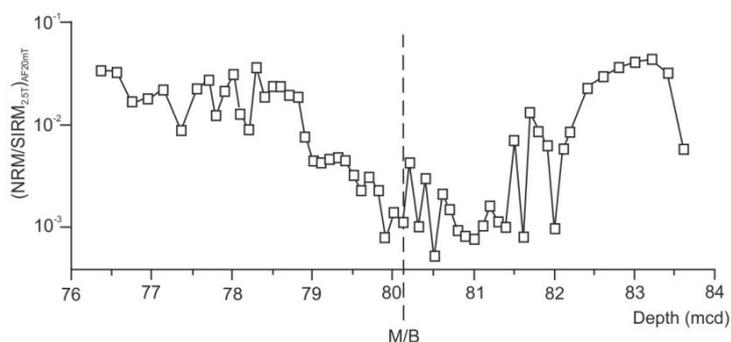
345 Valet et al. (2012) showed a 90° change in VGP during reversed polarity before the transition
346 from reversed to normal polarity (Fig. 7c). According to this model, the precursor prior to the
347 magnetic reversal has a 2.5 kyr duration and it occurs ~ 3.5 kyr before the actual transition
348 (mid-point) which has a 1 kyr duration (Fig. 7c). The model showed another 90° change as
349 the rebound with 2.5 kyr duration after the transition (Fig. 7c). Sagnotti et al. (2014) reported
350 Valet et al. (2012)'s precursor with 140° change in VGP, 0.7 kyr duration, and 5 kyr prior to
351 the actual transition. In our VGP data (Fig. 7a), 75° change between 13.6-11.8 cm depth
352 shows 3 kyr duration (according to 0.6 cm/kyr average sedimentation rate estimation) that can
353 be interpreted as the precursor of the Matuyama-Brunhes magnetic reversal. The pick point of
354 the precursor (11.8 cm depth) is 4.6 kyr before the actual transition (9.0 cm depth). The actual
355 transition duration is 0.8 kyr between 9.2-8.8 cm. These values are consistent with the Valet et
356 al. (2012) model and show the unique behaviour of the Earth's magnetic field during the
357 reversal time. The rebound after the transition in the model is not seen in our study since VGP
358 change between 7.1-3.2 cm is not enough to interpret it as the rebound.

359 *4.2. Sediment Deposition and Sedimentation Rate*

360 According to Lundberg et al. (2014) the cave was filled with water during the sedimentation
361 which was continuously active without any climatic change or hiatus. This provides the
362 continuous magnetic record of the reversal in the cave sediment and allows the sediment to
363 acquire and keep the primary magnetization without the possibility of secondary
364 mineralization. Also, there are no obvious marks in the studied sediment section that one
365 would expect if breaks in deposition (e.g. lithological boundaries, desiccation cracks).

366 Our estimate of the Za Hajovnou cave's sedimentation rate (0.2 cm/kyr – 1.0 cm/kyr) seems
367 to be significantly lower than the sediments from other studies (Sagnotti et al., 2010; Jin and
368 Liu, 2011; Giaccio et al., 2013; Liu et al., 2016; Okada et al., 2017). This is likely due to
369 contrasting sediment types. While the duration of the M/B transition was reported to last
370 between 4-13 kyr (Yamazaki and Oda, 2001; Suganuma et al., 2010; Valet et al., 2014; Okada
371 et al., 2017), the averaged sedimentation rate of 0.6 cm/kyr in this study suggests transition
372 duration of 9.66 kyr (7.1-12.8 cm sediment transition depth section) and thus strengthen our
373 confidence to the reliability of our sedimentation rate and paleomagnetic record estimates.
374 Additionally, low sedimentation rates (<1 cm/kyr) were reported in previous Matuyama-
375 Brunhes magnetic reversal studies (Clark, 1970; Love and Mazaud, 1997; Nowaczyk et al.,
376 2001).

377 Analyzes of cave sediments by paleomagnetism carried out in different locations around the
378 earth such as in Western Europe (Parés et al., 2018), South Africa (Nami et al., 2016), South
379 America (Jaqueto et al., 2016), North America (Stock et al., 2005), Southern Europe (Pruner
380 et al., 2010), Eastern Asia (Morinaga et al., 1992) showed that cave sediments recorded
381 magnetic reversals. Morinaga et al. (1992) suggested a low sedimentation rate for the Western
382 Japan cave sediments 1.6 cm/kyr which shows a similar rate with our Central European cave
383 sediment estimation. King and Channell (1991) suggested that large "lock-in" depths are
384 associated with interparticle rigidity and strength, characteristic of clayey low accumulation
385 rate sediments (<1 cm/kyr) which results in delays of magnetic acquisition. This shows that
386 magnetic polarity reversal could have a large (25 kyr) apparent age offset between sediments
387 with high and very low accumulation rates (King and Channell, 1991).



388

389 **Fig. 10.** NRM/SIRM ratio from Yamazaki and Oda (2001). The oscillations start from 81.6
 390 mcd in the ratio below the reversal. Dashed lines shows the M/B reversal.

391

392 **5. Conclusions**

393 We compared the paleomagnetic data from the cave sediments with published magnetic
 394 reversal record and were able for the first time to use the detailed magnetic characteristic of
 395 cave sediment and to infer the specific magnetic reversal (Matuyama-Brunhes). This is
 396 possible due to the nature of the magnetic reversal that we discovered. Note that the paleopole
 397 was residing in the east of Africa and then quickly reappeared west of North America. We
 398 consider this as an important marker signature for dating the central European paleomagnetic
 399 record from this time period.

400 The precursor event in our data is a significant anomaly to identify the behaviour of the
 401 Matuyama-Brunhes magnetic reversal. Additionally, we were able to design a new method for
 402 estimation of the accumulation rate of the cave sediment in Za Hajovnou cave, which until
 403 now was not known.

404

405 **Author contributions**

406 HU and GK designed research and wrote the paper, JK contributed to the fieldwork, gave
 407 detailed information about the cave sediment, and commented significantly.

408

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416

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577 **Data Availability Statement**

578 All data are incorporated into the article and its online supplementary material.

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580 **Supporting Information**

581 Supplementary data are available online.

582 Supporting Information Tables S1: Demagnetization data for the samples.

583 Supporting Information Figures S2: Zijderveld diagrams, Wulf stereonet, and

584 demagnetization curves of the rest of the samples for the alternative field (AF)
585 demagnetization method.
586 Supporting Information Figures S3: NRM intensity, MDF values, and VRM data