

# The Impact of Plate Motions on Long-Wavelength InSAR-Derived Velocity Fields

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## Key Points:

- Interferometric synthetic aperture radar (InSAR) time series are sensitive to uniform horizontal and vertical plate motions
- Such motions create long-wavelength spatial gradients that are visible in InSAR-derived velocity maps after removing other signals
- Plate motion effects can be easily accounted for using plate motion models, allowing long-wavelength deformation to be more clearly seen

**Abstract**

Interferometric Synthetic Aperture Radar (InSAR) measurements are increasingly being used to measure small amplitude tectonic deformations over large spatial scales. Residual signals are often present at these scales, and are interpreted to be noise of indeterminate origin, limiting studies of long-wavelength deformation. Here, we demonstrate the impact of bulk motion by the Earth's tectonic plates on InSAR-derived velocity fields. The range-dependent incidence angle of the InSAR observations, coupled with plate velocities of centimeters per year, can induce long-wavelength spatial gradients of millimeters per year over hundreds of kilometers in InSAR-derived velocity fields. We show that, after applying corrections, including for the ionosphere and troposphere, plate motion represents the dominant source of long-wavelength secular velocity gradients in multi-year time series for several study areas. This signal can be accounted for using plate motion models, allowing improved detection of regional tectonic strain at continental scales.

**Plain Language Summary**

Interferometric Synthetic Aperture Radar (InSAR) relies on repeat radar imaging to measure small motions of the Earth's surface. These motions can be used to understand a range of processes happening below the surface, from hydrology to tectonics. We show how the slow motion of Earth's tectonic plates can create a signal in InSAR data that may be confused with local tectonic signals. We also demonstrate a simple method for removing this signal, making InSAR more useful for studying very small motions over large areas of the Earth, especially in regions where we do not have good observations from other sources.

**1 Introduction**

Interferometric Synthetic Aperture Radar (InSAR) is an active imaging technique for measuring ground displacements that occur between repeat passes of an imaging platform, such as a satellite (e.g. Massonnet and Feigl (1998); Hanssen (2001); Simons and Rosen (2015)). InSAR deformation measurements are generally expressed relative to a single point, or ensemble of points, within the imaged area, usually assumed to be stable through time (e.g. Mahapatra et al. (2018)). While InSAR has been used extensively

45 for measuring large amplitude ( $> 1$  cm), deformation over short-wavelengths ( $< 100$  km)  
46 (e.g. Massonnet et al. (1993); Merryman Boncori (2019)), other signals present in the  
47 data challenge our ability to measure deformation at the scale of millimeters per year  
48 over hundreds of kilometers.

49 InSAR observations at long wavelengths are the combination of motion of the Earth's  
50 surface, changes in the atmosphere, and measurement and processing errors. The Earth  
51 motion signals comprise the surface deformation of interest, e.g. from tectonic strain,  
52 volcanic activity or subsidence (Massonnet et al. (1993); Massonnet and Feigl (1998);  
53 Amelung et al. (1999)), along with solid Earth tides (SET) (X. Xu & Sandwell, 2020),  
54 and ocean tidal loading (Dicaprio et al., 2008). Atmospheric signals come from propa-  
55 gation delay through the ionosphere (Z.-W. Xu et al., 2004) and troposphere (Tarayre  
56 & Massonnet, 1996). Errors sources include the satellite orbits (Massonnet & Feigl, 1998),  
57 local oscillator drift (Marinkovic & Larsen, 2015), phase unwrapping (Biggs et al., 2007)  
58 and topography (Berardino et al., 2002).

59 These effects can obscure small amplitude, long-wavelength signals in InSAR due  
60 to local tectonic processes, such as surface deformation from interseismic loading (e.g.  
61 Fournier et al. (2011); Parizzi et al. (2021)). Thus, it is common to not interpret long-  
62 wavelength signals from InSAR alone, instead removing them by empirically fitting 2D  
63 polynomial functions, known as “ramps”, to the data (e.g. Fialko (2006); Jolivet et al.  
64 (2015)), or combining InSAR velocities with Global Navigation Satellite System (GNSS)  
65 measurements in order to constrain the long-wavelength deformation (e.g. Weiss et al.  
66 (2020); X. Xu et al. (2021); Neely et al. (2020); Parizzi et al. (2020)). Such approaches  
67 are limiting when we wish to measure large-scale deformation in regions of sparse GNSS  
68 coverage (Chaussard et al., 2016; Neely et al., 2020).

69 The quality of InSAR data and correction methods have substantially increased  
70 over the last several years. The European Space Agency's (ESA) Sentinel-1 satellites have  
71 been regularly acquiring data for significant portions of the planet since late 2014. Sentinel-  
72 1 offers the advantages of improved orbital controls and uncertainties, reducing the noise  
73 contribution from satellite orbits (Fattahi & Amelung, 2014), as well as unrestricted data  
74 access. Split-band processing now allows for the estimation of the ionospheric signal di-  
75 rectly from the InSAR data (Gomba et al., 2016; Fattahi, Simons, & Agram, 2017; Liang  
76 et al., 2019), and higher quality weather models have improved the correction of tropo-

spheric phase (Li et al., 2005; Doin et al., 2009; Jolivet et al., 2011). Techniques for removing the SET (X. Xu & Sandwell, 2020) and ocean tidal loading signals (Dicaprio et al., 2008; Yu et al., 2020) have also been developed, among other correction methods. After corrections, there may still be long-wavelength residuals in multi-year Sentinel-1 time series, including from the troposphere, which can contribute up to 5 mm/yr over 150 km (Parizzi et al., 2021), and orbital errors, contributing around 0.5 mm/yr over 100 km for Sentinel-1 (Fattahi & Amelung, 2014).

In this work, we focus on the contribution of coherent uniform motion of Earth’s tectonic plates to the long-wavelength component of InSAR-derived velocity fields. The satellite line-of-sight (LOS) vector varies systematically in the satellite range direction (i.e. across the satellite track), causing a changing sensitivity to ground deformation with range. Bulk motion of tectonic plates in the satellite frame of reference, coupled with this LOS variation, can create quasi-linear gradients in InSAR-derived velocity fields, resulting in ramps, predominantly in the satellite range direction. This effect has been noted before, e.g., by Bähr et al. (2012), Bähr (2013) and Parizzi et al. (2020). Here, we demonstrate that plate motion creates ramps of several millimeters per year, across the 250 km track width, in six multi-year Sentinel-1 InSAR time series. After other corrections have been applied, plate motion is the dominant long-wavelength signal in our data, and we show that this signal can be straightforwardly compensated for using plate motion models. This adjustment is not currently part of several open-source InSAR time series analysis packages (e.g Hooper et al. (2012); Agram et al. (2013); Yunjun et al. (2019); Morishita et al. (2020)), and we provide an implementation of the method in the MintPy package (Yunjun et al., 2019).

## 2 The Reference Frame of InSAR Measurements

Quantifying ground deformation using InSAR requires a precise measurement of the satellite orbit ephemerides (Fattahi & Amelung, 2014; Peter, 2021). For Sentinel-1 the orbit is measured with respect to the International Terrestrial Reference Frame (ITRF) (Peter, 2021), an Earth-centered, Earth-fixed reference frame in which there is no net rotation of the Earth’s surface (Altamimi et al., 2016). Observations of absolute ground motion relative to the satellite are therefore also in ITRF (Bähr et al., 2012; Lazecky & Hooper, 2022).

108 However, it is not possible for InSAR to record absolute motions due to the  $2\pi$  am-  
 109 biguity in the interferometric phase (e.g. Massonnet and Feigl (1998)). Instead, displace-  
 110 ment measurements are generally expressed relative to a reference point within the im-  
 111 aged region, assumed to be stationary. Velocities can then be obtained from functional  
 112 fits to displacement time series, with inferred velocities also expressed relative to this point.

Selecting the reference point is not equivalent to expressing the InSAR velocities  
 in a reference frame moving with that point (Bähr et al., 2012; Bähr, 2013). We must  
 therefore consider how velocities in ITRF appear in the InSAR deformation field. We  
 represent the 3D ITRF secular velocity field of the Earth’s surface as:

$$\mathbf{v}(\mathbf{x}) = \mathbf{v}_p(\mathbf{x}) + \mathbf{v}_d(\mathbf{x}). \quad (1)$$

113  $\mathbf{v}_p(\mathbf{x})$  is the velocity field due to the strain-free motion of the relevant rigid plate in ITRF,  
 114 and  $\mathbf{v}_d(\mathbf{x})$  is the velocity due to internal deformation of the plate, for example due to  
 115 tectonic, volcanic, or hydrological processes.

Defining the LOS unit vector pointing from the ground to the satellite as  $\hat{\mathbf{l}}(\mathbf{x})$ , the  
 LOS projection of the 3D velocity field, minus the InSAR reference velocity, can be writ-  
 ten as:

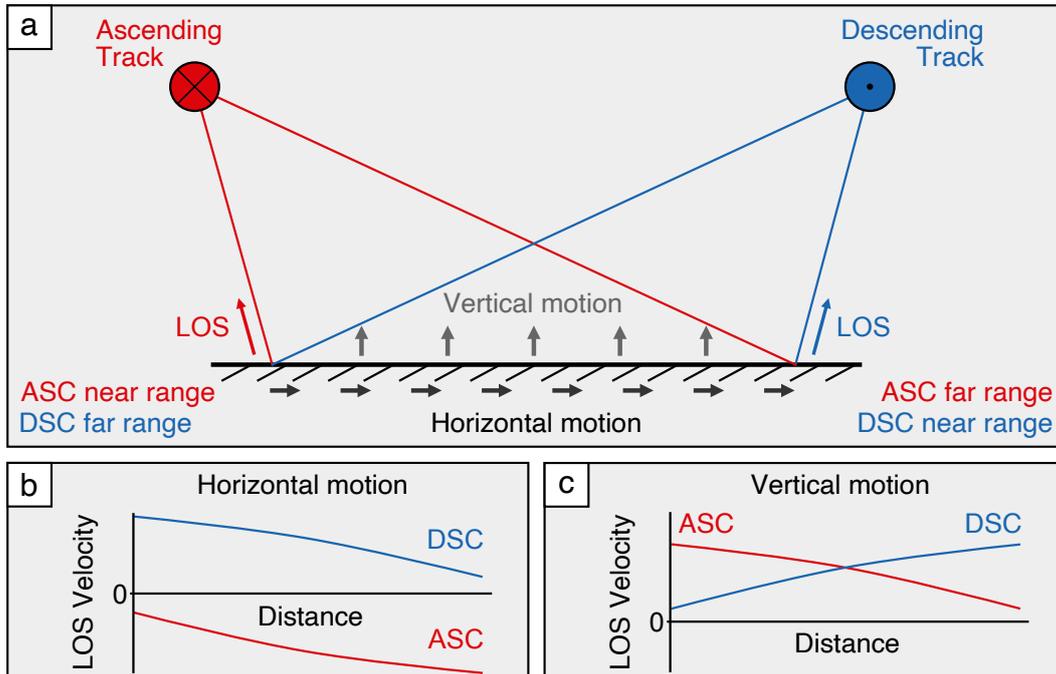
$$\mathbf{v}_l(\mathbf{x}) = \mathbf{v}(\mathbf{x}) \cdot \hat{\mathbf{l}}(\mathbf{x}) - \mathbf{v}(\mathbf{x}') \cdot \hat{\mathbf{l}}(\mathbf{x}'), \quad (2)$$

116 where the reference is at point  $\mathbf{x}'$ .  $\mathbf{v}_l(\mathbf{x})$  is the secular velocity that will be measured by  
 117 the satellite, assuming all other signals and noise can be neglected.

In ITRF,  $\mathbf{v}_l(\mathbf{x})$  has a contribution from the plate motion, which we can write as:

$$\mathbf{v}_{l,p}(\mathbf{x}) = \mathbf{v}_p(\mathbf{x}) \cdot \hat{\mathbf{l}}(\mathbf{x}) - \mathbf{v}_p(\mathbf{x}') \cdot \hat{\mathbf{l}}(\mathbf{x}'). \quad (3)$$

118 The second term, from the reference, is constant, while the first term depends on the spa-  
 119 tial variation of  $\mathbf{v}_p(\mathbf{x})$  and  $\hat{\mathbf{l}}(\mathbf{x})$ . The LOS vector  $\hat{\mathbf{l}}(\mathbf{x})$  can vary substantially over an im-  
 120 age swath. For Sentinel-1, the incidence angle (the angle between the LOS and the ver-  
 121 tical) varies approximately from  $29^\circ$  in the near range to  $46^\circ$  in the far range over the  
 122 250-km-width of the imaging swath (for data acquired in Interferometric Wideswath mode).  
 123 The range-dependent variation in  $\hat{\mathbf{l}}(\mathbf{x})$  implies a changing sensitivity to components of  
 124 the 3D deformation field across the track, with sensitivity to horizontal motion increas-  
 125 ing and vertical motion decreasing as we move from near range to far range. This range-  
 126 dependent sensitivity causes uniform plate motions to appear as velocity ramps in the  
 127 range direction when projected into the satellite LOS (Figure 1).



**Figure 1.** Illustration of how uniform horizontal and vertical motions result in ramps in InSAR-derived velocity measurements. **(a)** Satellite images acquired from ascending (ASC) and descending (DSC) orbital tracks, which have a varying LOS incidence angle across the track. The term “range” refers to the distance from the ground target to the satellite, with near range and far range the closest and furthest points from the satellite respectively. Red and blue arrows represent the ground-to-satellite LOS vector in the near range for ASC and DSC tracks, respectively. Grey and black arrows represent plate motion in the reference frame of the satellite. For illustration purposes we assume the ASC and DSC tracks are parallel to each other but in opposite directions, ignore Earth curvature. Figure not to scale. **(b)** Profile of the horizontal plate velocity projected into the LOS of the ASC and DSC tracks, against geographic distance along the ground. **(c)** Same as (b), except for vertical plate motion, resulting in opposite gradients in the LOS profiles. The observing geometry creates a small curvature in all profiles, which is exaggerated in the figure. For InSAR measurements, the LOS velocity is expressed relative to a point within the image, so each of these profiles would be vertically shifted to intersect with the  $x$  axis at the chosen reference point.

128 The plate velocity,  $\mathbf{v}_p(\mathbf{x})$ , also varies over an image swath. The motion of a rigid  
 129 plate on Earth’s surface can be represented by a rotation rate about an axis, known as  
 130 an Euler pole (McKenzie & Parker, 1967). Given the angular velocity of a chosen plate,  
 131  $\mathbf{\Omega}$ , we can write the velocity of any point,  $\mathbf{x}$ , on that plate as  $\mathbf{v}_p(\mathbf{x}) = \mathbf{\Omega} \times \mathbf{x}$ , where  
 132  $\times$  is the cross product. Thus, the velocity field due to rigid plate motion varies with dis-  
 133 tance from the plate’s Euler pole. This variation in  $\mathbf{v}_p(\mathbf{x})$  also contributes to the long-  
 134 wavelength LOS velocity field.

135 Because of the effect of plate motions, InSAR velocity measurements should not  
 136 generally be considered to be in a local reference frame, despite the use of a local ref-  
 137 erence point. Choosing a reference point within an InSAR image offsets InSAR veloc-  
 138 ity measurements from the LOS projection of ITRF velocities by an unknown constant  
 139 (Equation 1), but does not remove the long-wavelength gradients that can be induced  
 140 by plate motion (Equation 3). If ITRF plate motion is negligible when projected to the  
 141 LOS, and does not vary substantially over the InSAR track, or the satellite LOS vari-  
 142 ation across the track is small, then  $\mathbf{v}_{l,p}(\mathbf{x}) \approx 0$ . Choosing a reference point that is sta-  
 143 ble with respect to the plate is then approximately equivalent to putting the InSAR ve-  
 144 locities into the reference frame of that plate, however this should not be generally as-  
 145 sumed.

146 Several authors have investigated the reference frame of InSAR observations, gen-  
 147 erally in the context of using GNSS to put InSAR measurements into a terrestrial ref-  
 148 erence frame (e.g. Mahapatra et al. (2018); Johnston et al. (2021)). The influence of plate  
 149 motion on InSAR velocities has been noted by Bähr et al. (2012) and Bähr (2013), who  
 150 term it the *reference frame effect*. Bähr et al. (2012) present this phenomenon in terms  
 151 of a temporally increasing correction to the interferometric baseline, while Bähr (2013)  
 152 notes that this can also be framed in terms of the varying satellite LOS causing differ-  
 153 ing sensitivity to plate motion (the approach taken here). Parizzi et al. (2020) used plate  
 154 motion models to adjust their LOS velocity fields after merging InSAR with GNSS. Au-  
 155 thors have also noted the impact of plate motions on SAR geolocation accuracy (Cong  
 156 et al. (2012)). Our focus here is to demonstrate that plate motions can explain a signif-  
 157 icant fraction of observed residual long-wavelength surface velocities, after other correc-  
 158 tions have been applied, and without combining InSAR data with GNSS.

### 3 Data and Methods

#### 3.1 Data Processing

We present several examples using InSAR data from the ESA’s Sentinel-1 satellites, taken from ascending (ASC) and descending (DSC) tracks covering the Makran subduction zone (Iran), the Gulf of Aqaba (at the northern end of the Red Sea), and western Australia. For each track, we process at least 5 years of data using the InSAR Scientific Computing Environment (ISCE) (Rosen et al., 2012; Fattahi, Agram, & Simons, 2017). After forming the interferogram networks, we create deformation time series using MintPy (Yunjun et al., 2019).

Before examining residual signals due to plate motion, we apply corrections for the ionosphere, troposphere, SET, and digital elevation model (DEM) error. We use split-band processing to correct for the ionosphere (Liang et al., 2019), PyAPS and the ERA5 weather model to mitigate the tropospheric delay (Jolivet et al., 2014; Hersbach et al., 2020), the method of Fattahi and Amelung (2013) for DEM error correction and PySolid to correct for SET (Milbert, 2018; Yunjun et al., 2022). Further details of our data and processing are presented in the supporting information (Text S1, S2 and Table S1).

#### 3.2 Adjusting InSAR Measurements for Plate Motion

After all other corrections have been applied, we can then observe and account for the signal of plate motion. InSAR observations of ground motion are generally used to study regional deformation, rather than plate translations or rotations. For such purposes, a useful reference frame is one that moves with the plate in which we are trying to measure strain. Translating into this reference frame requires us to remove the signal of plate motion in the satellite’s frame of reference, i.e. ITRF (Bähr, 2013; Parizzi et al., 2020).

GNSS networks can be used to connect InSAR measurements to ITRF (e.g. Mahapatra et al. (2018); Johnston et al. (2021)), which can then be transformed into a reference frame moving with the chosen plate. In the absence of sufficient GNSS coverage, we can estimate the transformation into the plate’s frame of reference using the following steps:

1. Choose an InSAR reference point,  $\boldsymbol{x}'$ , that is stable with respect to the plate
2. Find the velocity field of the plate within ITRF, i.e.  $\boldsymbol{v}_p(\boldsymbol{x})$

- 189        3. Project that velocity field into the satellite LOS direction
- 190        4. Subtract the LOS velocity of the reference point,  $\mathbf{v}_{l,p}(\mathbf{x}')$ , from the projected plate
- 191        velocity to compute  $\mathbf{v}_{l,p}(\mathbf{x})$ , which is then removed from the InSAR velocity map.

192        Note that, after these steps, InSAR-derived velocities are still expressed relative to a ref-

193        erence point, meaning that deformation and other signals seen at the reference point will

194        still affect the entire scene.

195        We use the geodetically constrained ITRF plate motion model of Altamimi et al.

196        (2017) to estimate the plate velocity field. For each study region, we identify our refer-

197        ence plate (Table S1), then use the modeled angular velocity of the plate to calculate hor-

198        izontal velocities for our observation region. We then project these velocities into the LOS

199        direction and remove them from the velocity map.

## 200        4 Results

### 201        4.1 The Importance of Removing Other Signals for Revealing Plate Mo-

202        tion

203        We expect plate motion to contribute below 8 mm/yr across the 250 km width of

204        the Sentinel-1 tracks for our chosen regions (Figure S6), making it important to remove

205        other signals to show what fraction of the residual velocity can be explained by plate mo-

206        tion. For ASC tracks in the Makran and Gulf of Aqaba, ionosphere corrections have a

207        particularly large effect on the long-wavelength velocity signal (e.g. contributing a 25

208        mm/yr ramp along track 86 for the Makran, Figure 2), with DSC tracks showing sub-

209        stantially less ionospheric signal. ASC tracks are acquired at dusk—a period of greater

210        ionosphere activity than dawn, when DSC tracks are acquired. This impact is still no-

211        table in C-band Sentinel-1 data, even though it suffers much less from ionospheric ef-

212        fects than L-band (Fattahi, Simons, & Agram, 2017; Liang et al., 2019).

213        We find that troposphere corrections have a less significant impact on the long-wavelength

214        velocity signal than the ionosphere for ASC tracks, and a comparable effect for DSC tracks.

215        Corrections for the SET have a small effect on the long-wavelength secular velocity, con-

216        tributing below 0.5 mm/yr over several hundred kilometers. The range of DEM error

217        corrections is less than  $\pm 0.5$  mm/yr in our results and has a minimal contribution to the

218 long-wavelength velocity field. We show the impact of the above corrections for all tracks  
219 in Figures 2 and S1-S5, and present more details in Text S2.

## 220 **4.2 The Impact of Accounting for Plate Motion**

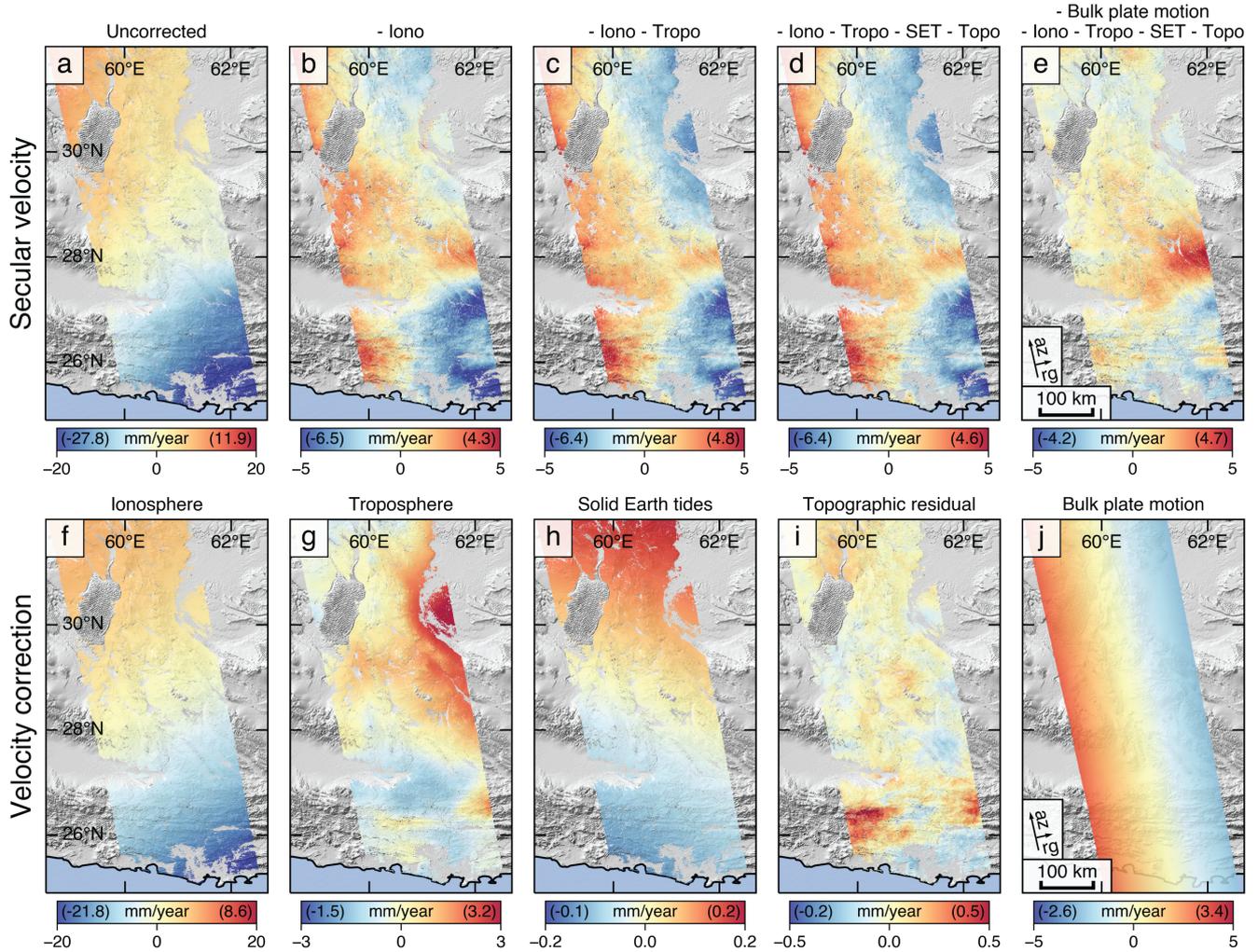
221 After applying the suite of corrections we are left with residual velocity ramps in  
222 all of our tracks, predominantly in the range direction. We present the results of plate  
223 motion adjustments for several tracks in Figures 3 and 4. Our results show that account-  
224 ing for plate motion removes a significant fraction of the residual velocity ramp in ev-  
225 ery case, reducing the across-track ramps from 4-7 mm/yr/track to below 1.5 mm/yr/track.  
226 For our data the plate motion signal is comparable to the troposphere in its effect on the  
227 long-wavelength velocity field.

228 The proximity of the Arabian plate Euler pole to the Gulf of Aqaba study area re-  
229 sults in the plate velocity field varying appreciably within the tracks (Altamimi et al.,  
230 2017). This variation causes an additional LOS velocity ramp along the track, with an  
231 opposite direction for the ascending and descending tracks. Figures 4 (a) and (b) show  
232 how this along track gradient can be clearly seen in the data, and is well corrected for  
233 by the plate motion model. We do not see similar along-track ramps for Australia and  
234 Makran, which is consistent with the plate motion velocity field.

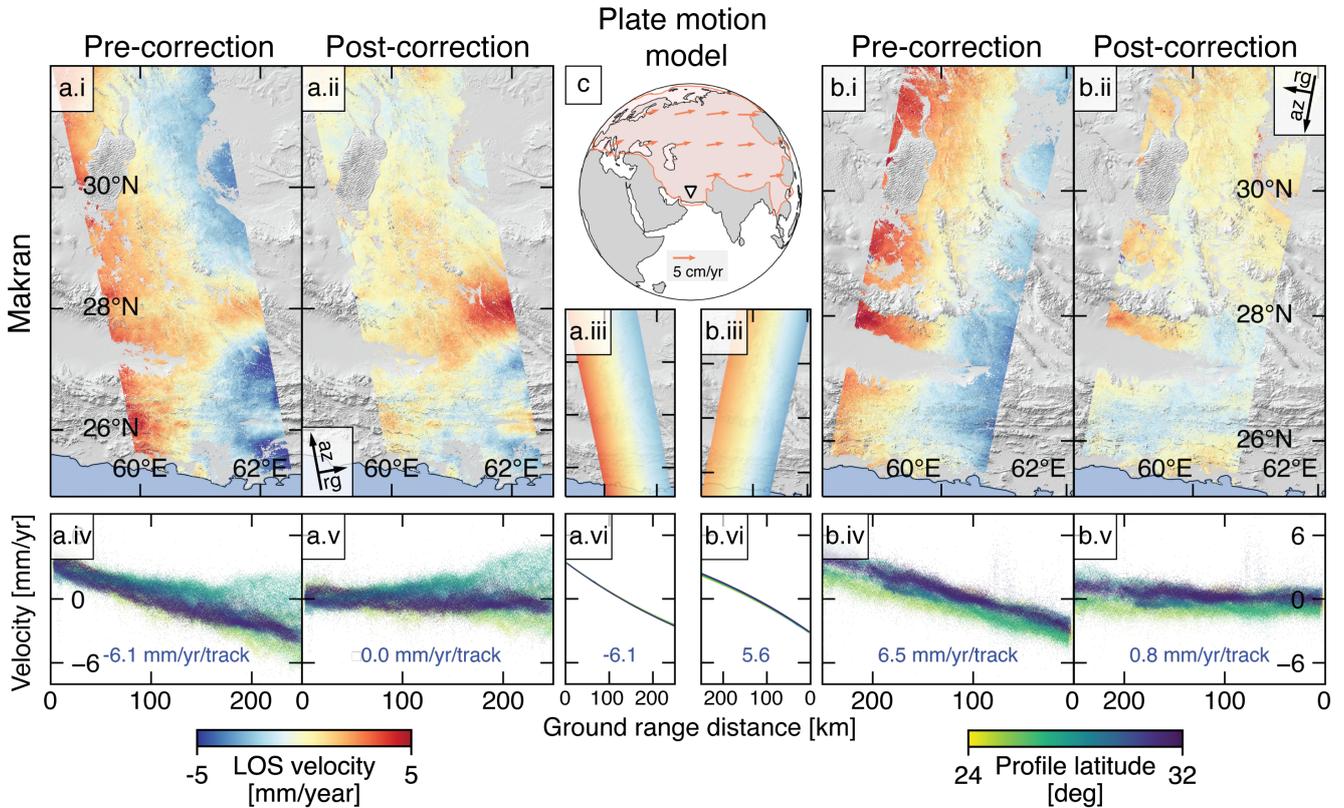
## 235 **5 Discussion**

236 Other authors have previously noted that plate motion will affect InSAR velocity  
237 measurements (e.g. Bähr et al. (2012)), however the narrower variation of the satellite  
238 LOS angle for earlier satellites, more limited data, and the presence of other significant  
239 long-wavelength signals, has made the signal difficult to isolate. The quality of recently  
240 available data and correction methods, and the wide swath of Sentinel-1, allow us to show  
241 the plate motion signal and the clear impact of accounting for it. Our results from the  
242 Gulf of Aqaba illustrate that plate rotation is an important part of the correction.

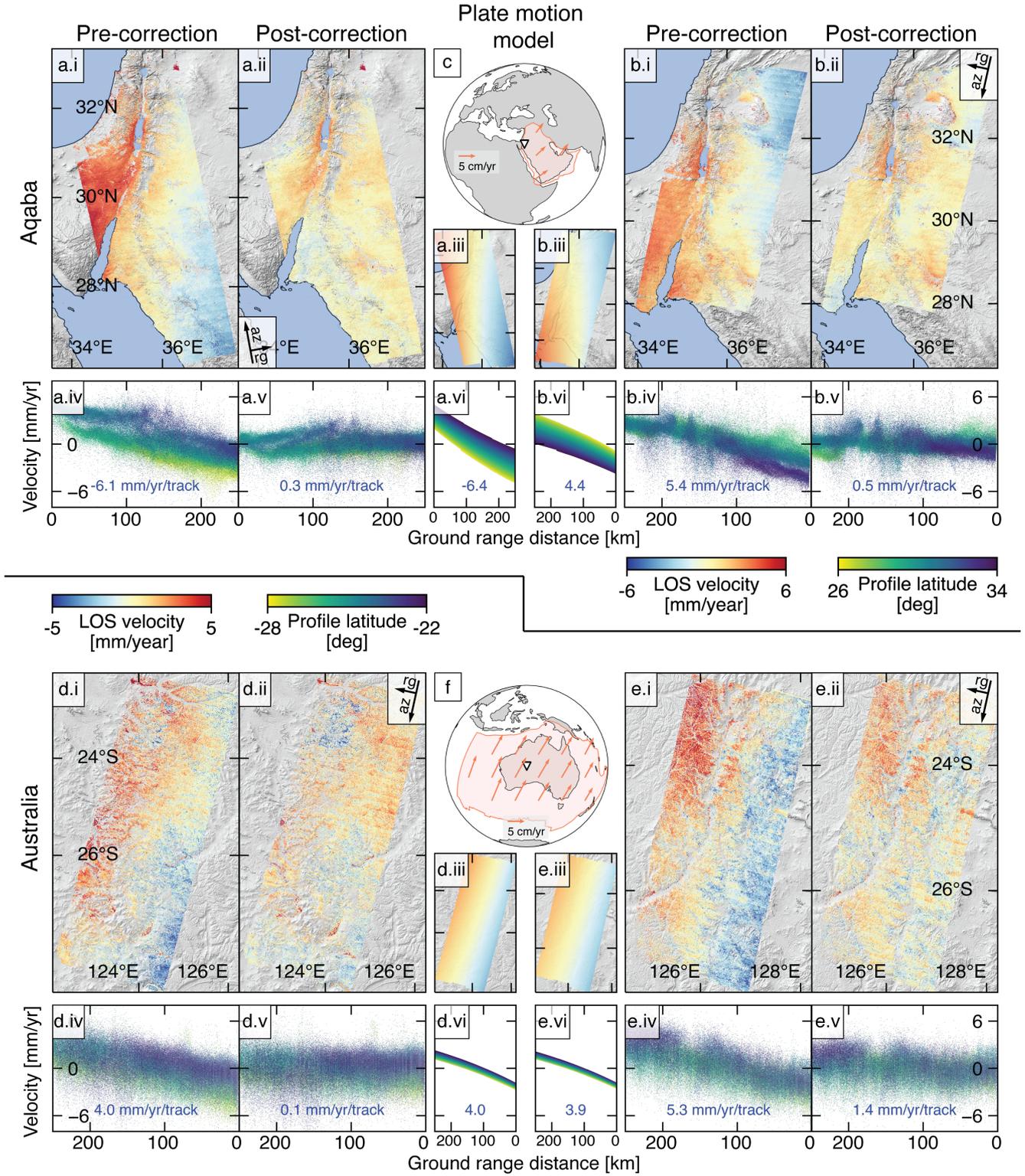
243 After adjusting for plate motion, remaining long-wavelength signals could be due  
244 to the incomplete removal of some signals (predominantly the troposphere (Fattahi &  
245 Amelung, 2015; Parizzi et al., 2021)), sources that we have not corrected for (e.g. ocean  
246 tidal loading (Dicaprio et al., 2008) and orbital errors (Fattahi & Amelung, 2014)), or  
247 actual strain accumulation in the lithosphere—the signal that InSAR measurements of-



**Figure 2.** Cumulative impact of corrections on the InSAR-derived velocity field for Sentinel-1, track 86 (ASC) over the Makran subduction zone. For plotting purposes, we remove the median value from each velocity field. Positive values represent apparent motion towards the satellite. Color bars are re-scaled between plots. Numbers in parentheses within the color bars refer to the 2nd and 98th percentiles of the velocity. “az” is the azimuth direction (satellite direction of motion), and “rg” the range direction (perpendicular to the satellite direction of motion). (a) No corrections applied. (b) Estimated ionosphere removed. (c) Tropospheric model removed. (d) SET model removed and DEM error (Topo) correction applied. (e) Plate motion correction applied. The positive signal around (28 °N, 62 °E) is post-seismic deformation from the 2013 Khash earthquake (Barnhart et al., 2014). (f) Applied ionospheric correction. (g) Applied tropospheric correction. (h) Applied SET correction. (i) Applied DEM error correction. Larger signals in the south may be bias from tropospheric residuals (Fattahi & Amelung, 2014). (j) Applied plate motion correction.



**Figure 3.** The impact of plate motion adjustments for InSAR tracks over the Makran subduction zone. For plotting purposes, we remove the median from each velocity field. **(a)** track 86 (ASC). After plate motion correction, the post-seismic signal from the 2013 Khash earthquake (Barnhart et al., 2014) can be more clearly seen in the south-east of the figure. **(b)** track 20 (DSC). **(c)** Location of tracks (a) and (b) and the velocity field of the Eurasian plate, used to correct the tracks. **(i)** Velocity before plate motion correction, but after other corrections have been applied. **(ii)** Velocity after plate motion correction. **(iii)** Applied plate motion correction. **(iv).** Across track profile of the velocity before plate motion correction. The number below each profile is the gradient of the linear least squares fit to the profile. Note that profiles are plotted as a function of ground range, which increases with distance from the satellite. **(v)** Across track profile of the velocity after plate motion correction. **(vi)** Across track profile of the applied plate motion correction.



**Figure 4.** Same as Figure 3, but showing the tracks for the Gulf of Aqaba and Australia. (a) Gulf of Aqaba, track 87 (ASC). (b) Gulf of Aqaba, track 21 (DSC). (c) Location of tracks (a) and (b) and the velocity field of the Arabian plate, used to correct the tracks. (d) Australia, track 119 (DSC) (e) Australia, track 46 (DSC). (f) Location of tracks (d) and (e) and the velocity field of the Australian plate, used to correct the tracks. (i)-(vi) are as described in Figure 3.

248 ten target. See Text S2 for more details on the contributors to long-wavelength resid-  
249 uals.

250 Deficiencies in the plate motion model, or motion of the InSAR track reference point  
251 relative to the assumed plate, could also create long-wavelength residuals. Motion rel-  
252 ative to the plate will be of particular importance in areas of diffuse plate boundary de-  
253 formation, where it is not possible to choose a reference point that is stable with respect  
254 to the rigid plate. This could be the case for tracks covering the Makran subduction zone  
255 and the Gulf of Aqaba, both of which span plate boundary zones. In these situations,  
256 plate motion models may not fully account for the impacts of bulk motion, and using  
257 local GNSS measurements to put InSAR measurements into a local terrestrial reference  
258 frame could be necessary (Bähr, 2013).

259 These results emphasise the importance of accounting for the reference frame be-  
260 fore interpreting long-wavelength InSAR-derived velocity fields. When using InSAR for  
261 studies of tectonic deformation, the most natural reference frame is one that is fixed to  
262 a stable region within the scene, so that we can interpret velocity gradients in terms of  
263 tectonic strain rather than strain-free translation and rotation. There are several situ-  
264 ations in which failing to account for the reference frame could bias the results:

- 265 1. Combining multiple tracks to estimate 3D deformation (Fialko et al., 2001; Wright  
266 et al., 2004). In Text S4 and Figure S7 we show how plate motion can bias esti-  
267 mates of the 3D velocity field when we use an overlapping ascending and descend-  
268 ing track to calculate horizontal and vertical velocities.
- 269 2. Modeling InSAR signals. If the long-wavelength signals in an InSAR velocity field  
270 are being modeled, and the model is assumed to not be rotating or translating,  
271 then a velocity ramp from plate motion may be modeled as strain accumulation  
272 and bias the results (e.g. changing the locking depth in a subduction zone model).
- 273 3. Comparisons between GNSS and InSAR. Both data sets must be in the same ref-  
274 erence frame (Parizzi et al., 2020). If the GNSS are in a local reference frame, the  
275 InSAR and GNSS velocities will diverge at long wavelengths due to the signal of  
276 plate motion in the InSAR.

277 Studies which removed ramps from InSAR-derived velocities to account for orbital  
278 errors may have inadvertently removed the impact of plate motion in their observations  
279 as well, reducing the biases we outlined above.

280 In this work, our primary focus is on the impact of horizontal plate motions, and  
281 we have not considered the contribution of long-wavelength vertical velocities. Horizontal  
282 plate motions in ITRF are generally at the scale of centimeters per year (Altamimi  
283 et al., 2017), with long-wavelength vertical motions, for example due to post-glacial re-  
284 bound, significantly smaller at millimeters per year (e.g. Riddell et al. (2020); Lau et al.  
285 (2020)). If an InSAR track is taken within a region that is experiencing constant ver-  
286 tical motion, this motion will also create a velocity ramp in the satellite LOS velocity  
287 field, but with ASC and DSC tracks having opposite gradients (Figure 1(c)). However,  
288 the amplitude of vertical velocities will result in smaller velocity gradients across the satel-  
289 lite track than those caused by horizontal motion (Text S3, Figure S6).

## 290 **6 Conclusion**

291 We have illustrated how InSAR velocity measurements are sensitive to tectonic plate  
292 motion in the satellite reference frame. This motion will induce ramps in the InSAR ve-  
293 locity fields, predominantly in the satellite range direction, of up to several millimeters  
294 per year. In all of our multi-year time series, plate motion was the dominant long-wavelength  
295 signal after ionospheric and tropospheric corrections were applied. We have presented  
296 a simple adjustment method, which uses plate motion models to remove the plate mo-  
297 tion signal from the InSAR velocity field. This adjustment substantially reduces long-  
298 wavelength ramps in multiple InSAR tracks from three different regions of the Earth.  
299 Routinely accounting for plate motion in InSAR could reduce biases when constraining  
300 long-wavelength tectonic strain induced by local geophysical phenomena. This adjust-  
301 ment is likely to be particularly useful where GNSS is not available to constrain the long-  
302 wavelength deformation. The signal of plate motion in InSAR data could also be used  
303 to improve plate motion models, which may be helpful where GNSS observations are sparse  
304 but high-quality InSAR data are available.

## 305 **7 Open Research**

306 The Sentinel-1 data were provided by the European Space Agency and downloaded  
307 from the Alaska Satellite Facility. InSAR data were processed using the InSAR Scien-

308 tific Computing Environment (ISCE) (Rosen et al., 2012), available at: [https://github](https://github.com/isce-framework/isce2)  
 309 [.com/isce-framework/isce2](https://github.com/isce-framework/isce2). Time series analysis was performed using the MintPy soft-  
 310 ware (Yunjun et al., 2019), available at: <https://github.com/insarlab/MintPy>. Our  
 311 plate motion correction method is implemented in MintPy as `bulk_plate_motion.py`  
 312 (for review purposes this is available at: [https://github.com/yuankailiu/MintPy/blob/](https://github.com/yuankailiu/MintPy/blob/GRL/mintpy/bulk_plate_motion.py)  
 313 [GRL/mintpy/bulk\\_plate\\_motion.py](https://github.com/yuankailiu/MintPy/blob/GRL/mintpy/bulk_plate_motion.py)). Other data processing was performed using Python.  
 314 Plots were produced using Matplotlib and Cartopy in Jupyter Notebooks, available at  
 315 <https://zenodo.org/record/6606282>.

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### 320 References

- 321 Agram, P. S., Jolivet, R., Riel, B., Lin, Y. N., Simons, M., Hetland, E., . . . Lasserre,  
 322 C. (2013). New Radar Interferometric Time Series Analysis Toolbox Re-  
 323 leased. *Eos, Transactions American Geophysical Union*, *94*(7), 69–70.  
 324 Retrieved from <http://doi.wiley.com/10.1002/2013E0070001> doi:  
 325 [10.1002/2013E0070001](https://doi.org/10.1002/2013E0070001)
- 326 Altamimi, Z., Métivier, L., & Collilieux, X. (2012). ITRF2008 plate motion  
 327 model. *Journal of Geophysical Research: Solid Earth*, *117*(7), 1–14. doi:  
 328 [10.1029/2011JB008930](https://doi.org/10.1029/2011JB008930)
- 329 Altamimi, Z., Métivier, L., Rebischung, P., Rouby, H., & Collilieux, X. (2017).  
 330 ITRF2014 plate motion model. *Geophysical Journal International*, *209*(3),  
 331 1906–1912. doi: [10.1093/gji/ggx136](https://doi.org/10.1093/gji/ggx136)
- 332 Altamimi, Z., Rebischung, P., Métivier, L., & Collilieux, X. (2016). ITRF2014: A  
 333 new release of the International Terrestrial Reference Frame modeling nonlin-  
 334 ear station motions. *Journal of Geophysical Research: Solid Earth*, *121*(8),  
 335 6109–6131. Retrieved from <http://doi.wiley.com/10.1002/2016JB013098>  
 336 doi: [10.1002/2016JB013098](https://doi.org/10.1002/2016JB013098)
- 337 Amelung, F., Galloway, D. L., Bell, J. W., Zebker, H. A., & Laczniak, R. J. (1999).  
 338 Sensing the ups and downs of Las Vegas: InSAR reveals structural control of

- 339 land subsidence and aquifer-system deformation. *Geology*, 27(6), 483. Re-  
340 trieved from [https://pubs.geoscienceworld.org/geology/article/27/6/](https://pubs.geoscienceworld.org/geology/article/27/6/483-486/207094)  
341 [483-486/207094](https://pubs.geoscienceworld.org/geology/article/27/6/483-486/207094) doi: 10.1130/0091-7613(1999)027<0483:STUADO>2.3.CO;2
- 342 Ansari, H., De Zan, F., & Parizzi, A. (2021). Study of Systematic Bias in Mea-  
343 suring Surface Deformation with SAR Interferometry. *IEEE Transactions on*  
344 *Geoscience and Remote Sensing*, 59(2), 1285–1301. doi: 10.1109/TGRS.2020  
345 .3003421
- 346 Bähr, H. (2013). *Orbital effects in spaceborne synthetic aperture radar interferometry*  
347 (Doctoral dissertation). doi: 10.5445/KSP/1000037166
- 348 Bähr, H., Samiei-Esfahany, S., & Hanssen, R. F. (2012). On The Effect Of Reference  
349 Frame Motion On InSAR Deformation Estimates. *Proceedings of Fringe 2011*.
- 350 Barnhart, W. D., Hayes, G. P., Samsonov, S. V., Fielding, E. J., & Seidman, L. E.  
351 (2014). Breaking the oceanic lithosphere of a subducting slab: The 2013  
352 Khash, Iran earthquake. *Geophysical Research Letters*, 41(1), 32–36. doi:  
353 10.1002/2013GL058096
- 354 Bekaert, D., Walters, R., Wright, T., Hooper, A., & Parker, D. (2015). Statisti-  
355 cal comparison of InSAR tropospheric correction techniques. *Remote Sensing*  
356 *of Environment*, 170, 40–47. Retrieved from [https://linkinghub.elsevier](https://linkinghub.elsevier.com/retrieve/pii/S0034425715301231)  
357 [.com/retrieve/pii/S0034425715301231](https://linkinghub.elsevier.com/retrieve/pii/S0034425715301231) doi: 10.1016/j.rse.2015.08.035
- 358 Berardino, P., Fornaro, G., Lanari, R., & Sansosti, E. (2002). A new algorithm  
359 for surface deformation monitoring based on small baseline differential SAR  
360 interferograms. *IEEE Transactions on Geoscience and Remote Sensing*,  
361 40(11), 2375–2383. Retrieved from [http://ieeexplore.ieee.org/document/](http://ieeexplore.ieee.org/document/1166596/)  
362 [1166596/](http://ieeexplore.ieee.org/document/1166596/) doi: 10.1109/TGRS.2002.803792
- 363 Biggs, J., Wright, T., Lu, Z., & Parsons, B. (2007). Multi-interferogram method for  
364 measuring interseismic deformation: Denali Fault, Alaska. *Geophysical Journal*  
365 *International*, 170(3), 1165–1179. doi: 10.1111/j.1365-246X.2007.03415.x
- 366 Chaussard, E., Johnson, C. W., Fattahi, H., & Bürgmann, R. (2016). Poten-  
367 tial and limits of InSAR to characterize interseismic deformation indepen-  
368 dently of GPS data: Application to the southern San Andreas Fault system.  
369 *Geochemistry, Geophysics, Geosystems*, 17(3), 1214–1229. Retrieved from  
370 <https://onlinelibrary.wiley.com/doi/10.1002/2015GC006246> doi:  
371 10.1002/2015GC006246

- 372 Chen, C. W., & Zebker, H. A. (2002). Phase unwrapping for large SAR interfero-  
373 grams: Statistical segmentation and generalized network models. *IEEE Trans-*  
374 *actions on Geoscience and Remote Sensing*, 40(8), 1709–1719. doi: 10.1109/  
375 TGRS.2002.802453
- 376 Cong, X., Balss, U., Eineder, M., & Fritz, T. (2012). Imaging Geodesy—Centimeter-  
377 Level Ranging Accuracy With TerraSAR-X: An Update. *IEEE Geoscience and*  
378 *Remote Sensing Letters*, 9(5), 948–952. Retrieved from [http://ieeexplore](http://ieeexplore.ieee.org/document/6170871/)  
379 [.ieeexplore.org/document/6170871/](http://ieeexplore.ieee.org/document/6170871/) doi: 10.1109/LGRS.2012.2187042
- 380 De Zan, F., Parizzi, A., Prats-Iraola, P., & López-Dekker, P. (2014). A SAR interfer-  
381 ometric model for soil moisture. *IEEE Transactions on Geoscience and Remote*  
382 *Sensing*, 52(1), 418–425. doi: 10.1109/TGRS.2013.2241069
- 383 Dicaprio, C. J., Simons, M., Kenner, S. J., & Williams, C. A. (2008). Post-seismic  
384 reloading and temporal clustering on a single fault. *Geophysical Journal Inter-*  
385 *national*, 172(2), 581–592. doi: 10.1111/j.1365-246X.2007.03622.x
- 386 Doin, M. P., Lasserre, C., Peltzer, G., Cavalié, O., & Doubre, C. (2009). Correc-  
387 tions of stratified tropospheric delays in SAR interferometry: Validation with  
388 global atmospheric models. *Journal of Applied Geophysics*, 69(1), 35–50. Re-  
389 trieved from <http://dx.doi.org/10.1016/j.jappgeo.2009.03.010> doi:  
390 10.1016/j.jappgeo.2009.03.010
- 391 Emardson, T. R., Simons, M., & Webb, F. H. (2003). Neutral atmospheric delay  
392 in interferometric synthetic aperture radar applications: Statistical descrip-  
393 tion and mitigation. *Journal of Geophysical Research: Solid Earth*, 108(B5),  
394 1–8. Retrieved from <http://doi.wiley.com/10.1029/2002JB001781> doi:  
395 10.1029/2002JB001781
- 396 Fattahi, H., Agram, P., & Simons, M. (2017). A Network-Based Enhanced  
397 Spectral Diversity Approach for TOPS Time-Series Analysis. *IEEE*  
398 *Transactions on Geoscience and Remote Sensing*, 55(2), 777–786. Re-  
399 trieved from <http://ieeexplore.ieee.org/document/7637021/> doi:  
400 10.1109/TGRS.2016.2614925
- 401 Fattahi, H., & Amelung, F. (2013). DEM Error Correction in InSAR Time Series.  
402 *Geoscience and Remote Sensing, IEEE Transactions on*, 51(7), 4249–4259.  
403 Retrieved from <http://ieeexplore.ieee.org/document/6423275/> doi:  
404 10.1109/TGRS.2012.2227761

- 405 Fattahi, H., & Amelung, F. (2014). InSAR uncertainty due to orbital errors. *Geo-*  
406 *physical Journal International*, *199*(1), 549–560. doi: 10.1093/gji/ggu276
- 407 Fattahi, H., & Amelung, F. (2015). InSAR bias and uncertainty due to the system-  
408 atic and stochastic tropospheric delay. *Journal of Geophysical Research: Solid*  
409 *Earth*, *120*(12), 8758–8773. Retrieved from [https://onlinelibrary.wiley](https://onlinelibrary.wiley.com/doi/abs/10.1002/2015JB012419)  
410 [.com/doi/abs/10.1002/2015JB012419](https://onlinelibrary.wiley.com/doi/abs/10.1002/2015JB012419) doi: 10.1002/2015JB012419
- 411 Fattahi, H., Simons, M., & Agram, P. (2017). InSAR Time-Series Estimation  
412 of the Ionospheric Phase Delay: An Extension of the Split Range-Spectrum  
413 Technique. *IEEE Transactions on Geoscience and Remote Sensing*, *55*(10),  
414 5984–5996. doi: 10.1109/TGRS.2017.2718566
- 415 Fialko, Y. (2006). Interseismic strain accumulation and the earthquake potential on  
416 the southern San Andreas fault system. *Nature*, *441*(7096), 968–971. doi: 10  
417 [.1038/nature04797](https://doi.org/10.1038/nature04797)
- 418 Fialko, Y., Simons, M., & Agnew, D. (2001). The complete (3-D) surface displace-  
419 ment field in the epicentral area of the 1999 Mw 7.1 Hector Mine earthquake,  
420 California, from space geodetic observations. *Geophysical Research Letters*,  
421 *28*(16), 3063–3066. doi: 10.1029/2001GL013174
- 422 Fournier, T., Pritchard, M. E., & Finnegan, N. (2011). Accounting for Atmo-  
423 spheric Delays in InSAR Data in a Search for Long-Wavelength Deformation  
424 in South America. *IEEE Transactions on Geoscience and Remote Sensing*,  
425 *49*(10), 3856–3867. Retrieved from [http://ieeexplore.ieee.org/document/](http://ieeexplore.ieee.org/document/5771553/)  
426 [5771553/](http://ieeexplore.ieee.org/document/5771553/) doi: 10.1109/TGRS.2011.2139217
- 427 Gomba, G., Parizzi, A., De Zan, F., Eineder, M., & Bamler, R. (2016). Toward  
428 operational compensation of ionospheric effects in SAR interferograms: The  
429 split-spectrum method. *IEEE Transactions on Geoscience and Remote Sens-*  
430 *ing*, *54*(3), 1446–1461. doi: 10.1109/TGRS.2015.2481079
- 431 Hanssen, R. F. (2001). *Radar Interferometry: Data Interpretation and Error Anal-*  
432 *ysis* (Vol. 2). Dordrecht: Springer Netherlands. Retrieved from [http://link](http://link.springer.com/10.1007/0-306-47633-9)  
433 [.springer.com/10.1007/0-306-47633-9](http://link.springer.com/10.1007/0-306-47633-9) doi: 10.1007/0-306-47633-9
- 434 Hersbach, H., Bell, B., Berrisford, P., Hirahara, S., Horányi, A., Muñoz-Sabater,  
435 J., ... Thépaut, J. (2020). The ERA5 global reanalysis. *Quarterly Jour-*  
436 *nal of the Royal Meteorological Society*, *146*(730), 1999–2049. Retrieved  
437 from <https://onlinelibrary.wiley.com/doi/10.1002/qj.3803> doi:

438 10.1002/qj.3803

439 Hooper, A., Bekaert, D., Spaans, K., & Arikan, M. (2012). Recent advances in SAR  
440 interferometry time series analysis for measuring crustal deformation. *Tectono-*  
441 *physics*, *514-517*, 1–13. doi: 10.1016/j.tecto.2011.10.013

442 Johnston, P. J., Filmer, M. S., & Fuhrmann, T. (2021). Evaluation of methods  
443 for connecting InSAR to a terrestrial reference frame in the Latrobe Valley,  
444 Australia. *Journal of Geodesy*, *95*(10). Retrieved from [https://doi.org/](https://doi.org/10.1007/s00190-021-01560-2)  
445 [10.1007/s00190-021-01560-2](https://doi.org/10.1007/s00190-021-01560-2) doi: 10.1007/s00190-021-01560-2

446 Jolivet, R., Agram, P. S., Lin, N. Y., Simons, M., Doin, M.-p., Peltzer, G., & Li, Z.  
447 (2014). Improving InSAR geodesy using Global Atmospheric Models. *Jour-*  
448 *nal of Geophysical Research: Solid Earth*, *119*(3), 2324–2341. Retrieved from  
449 <https://onlinelibrary.wiley.com/doi/abs/10.1002/2013JB010588> doi:  
450 [10.1002/2013JB010588](https://doi.org/10.1002/2013JB010588)

451 Jolivet, R., Grandin, R., Lasserre, C., Doin, M. P., & Peltzer, G. (2011). Sys-  
452 tematic InSAR tropospheric phase delay corrections from global meteorolo-  
453 gical reanalysis data. *Geophysical Research Letters*, *38*(17), 1–6. doi:  
454 [10.1029/2011GL048757](https://doi.org/10.1029/2011GL048757)

455 Jolivet, R., Simons, M., Agram, P. S., Duputel, Z., & Shen, Z.-K. (2015). Aseismic  
456 slip and seismogenic coupling along the central San Andreas Fault. *Geophysical*  
457 *Research Letters*, *42*(2), 297–306. Retrieved from [http://doi.wiley.com/10](http://doi.wiley.com/10.1002/2014GL062222)  
458 [.1002/2014GL062222](http://doi.wiley.com/10.1002/2014GL062222) doi: 10.1002/2014GL062222

459 Lau, N., Borsa, A. A., & Becker, T. W. (2020). Present-Day Crustal Vertical Veloc-  
460 ity Field for the Contiguous United States. *Journal of Geophysical Research:*  
461 *Solid Earth*, *125*(10). Retrieved from [https://onlinelibrary.wiley.com/](https://onlinelibrary.wiley.com/doi/10.1029/2020JB020066)  
462 [doi/10.1029/2020JB020066](https://doi.org/10.1029/2020JB020066) doi: 10.1029/2020JB020066

463 Lazecky, M., & Hooper, A. (2022). InSAR-derived horizontal velocities in  
464 a global reference frame. *ESSOAr (preprint)*, 1–14. Retrieved from  
465 <https://www.essoar.org/doi/abs/10.1002/essoar.10511058.1> doi:  
466 <https://doi.org/10.1002/essoar.10511058.1>

467 Li, Z., Muller, J. P., Cross, P., & Fielding, E. J. (2005). Interferometric synthetic  
468 aperture radar (InSAR) atmospheric correction: GPS, Moderate Resolution  
469 Imaging Spectroradiometer (MODIS), and InSAR integration. *Journal of*  
470 *Geophysical Research: Solid Earth*, *110*(3), 1–10. doi: 10.1029/2004JB003446

- 471 Li, Z., Wright, T., Hooper, A., Crippa, P., Gonzalez, P., Walters, R., . . . Par-  
 472 sons, B. (2016). Towards InSAR everywhere, all the time, with Sentinel-  
 473 1. *International Archives of the Photogrammetry, Remote Sensing and*  
 474 *Spatial Information Sciences - ISPRS Archives*, 41(July), 763–766. doi:  
 475 10.5194/isprsarchives-XLI-B4-763-2016
- 476 Liang, C., Agram, P., Simons, M., & Fielding, E. J. (2019). Ionospheric Correction  
 477 of InSAR Time Series Analysis of C-band Sentinel-1 TOPS Data. *IEEE Trans-*  
 478 *actions on Geoscience and Remote Sensing*, 57(9), 6755–6773. doi: 10.1109/  
 479 tgrs.2019.2908494
- 480 Mahapatra, P., der Marel, H. v., van Leijen, F., Samiei-Esfahany, S., Klees,  
 481 R., & Hanssen, R. (2018). InSAR datum connection using GNSS-  
 482 augmented radar transponders. *Journal of Geodesy*, 92(1), 21–32. doi:  
 483 10.1007/s00190-017-1041-y
- 484 Marinkovic, P., & Larsen, Y. (2015). On Resolving the Local Oscillator Drift In-  
 485 duced Phase Ramps in ASAR and ERS1/2 Interferometric Data—The Final  
 486 Solution. *Fringe 2015 Workshop (ESA SP - 731)(1)*, 20.
- 487 Massonnet, D., & Feigl, K. L. (1998). Radar interferometry and its application to  
 488 changes in the Earth’s surface. *Reviews of Geophysics*, 36(4), 441. Retrieved  
 489 from <http://doi.wiley.com/10.1029/97RG03139> doi: 10.1029/97RG03139
- 490 Massonnet, D., Rossi, M., Carmona, C., Adragna, F., Peltzer, G., Feigl, K., &  
 491 Rabaute, T. (1993). The displacement field of the Landers earthquake  
 492 mapped by radar interferometry. *Nature*, 364(6433), 138–142. Retrieved  
 493 from <http://www.nature.com/articles/364138a0> doi: 10.1038/364138a0
- 494 McKenzie, D. P., & Parker, R. L. (1967). The North Pacific: an Example of Tecton-  
 495 ics on a Sphere. *Nature*, 216(5122), 1276–1280. Retrieved from [https://www](https://www.nature.com/articles/2161276a0)  
 496 [.nature.com/articles/2161276a0](https://www.nature.com/articles/2161276a0) doi: 10.1038/2161276a0
- 497 Merryman Boncori, J. P. (2019). Measuring Coseismic Deformation With Space-  
 498 borne Synthetic Aperture Radar: A Review. *Frontiers in Earth Science*,  
 499 7(February), 1–20. doi: 10.3389/feart.2019.00016
- 500 Milbert, D. (2018). *solid: Solid Earth Tide*. Retrieved from [https://geodesyworld](https://geodesyworld.github.io/SOFTS/solid.htm)  
 501 [.github.io/SOFTS/solid.htm](https://geodesyworld.github.io/SOFTS/solid.htm)
- 502 Morishita, Y., Lazecky, M., Wright, T., Weiss, J., Elliott, J., & Hooper, A. (2020).  
 503 LiCSBAS: An Open-Source InSAR Time Series Analysis Package Integrated

- 504 with the LiCSAR Automated Sentinel-1 InSAR Processor. *Remote Sensing*,  
505 *12*(3), 424. Retrieved from <https://www.mdpi.com/2072-4292/12/3/424>  
506 doi: 10.3390/rs12030424
- 507 Neely, W. R., Borsa, A. A., & Silverii, F. (2020). GInSAR: A cGPS Correction for  
508 Enhanced InSAR Time Series. *IEEE Transactions on Geoscience and Remote*  
509 *Sensing*, *58*(1), 136–146. Retrieved from [https://ieeexplore.ieee.org/](https://ieeexplore.ieee.org/document/8839742/)  
510 [document/8839742/](https://ieeexplore.ieee.org/document/8839742/) doi: 10.1109/TGRS.2019.2934118
- 511 Parizzi, A., Brcic, R., & De Zan, F. (2021). InSAR Performance for Large-Scale De-  
512 formation Measurement. *IEEE Transactions on Geoscience and Remote Sens-*  
513 *ing*, *59*(10), 8510–8520. doi: 10.1109/TGRS.2020.3039006
- 514 Parizzi, A., Gonzalez, F. R., & Brcic, R. (2020). A covariance-based approach to  
515 merging InSAR and GNSS displacement rate measurements. *Remote Sensing*,  
516 *12*(2). doi: 10.3390/rs12020300
- 517 Peter, H. (2021). *Copernicus POD Product Handbook: Copernicus Sentinel-1, -2*  
518 *and -3 Precise Orbit Determination Service (CPOD)* (Tech. Rep.). GMV.  
519 Retrieved from [https://sentinel.esa.int/documents/247904/4599719/](https://sentinel.esa.int/documents/247904/4599719/Copernicus-POD-Product-Handbook.pdf)  
520 [Copernicus-POD-Product-Handbook.pdf](https://sentinel.esa.int/documents/247904/4599719/Copernicus-POD-Product-Handbook.pdf)
- 521 Petit, G., & Luzum, B. (2010). *IERS Conventions (IERS Technical Note No. 36)*.  
522 Frankfurt, Germany. Retrieved from <http://www.iers.org/TN36/>
- 523 Riddell, A. R., King, M. A., & Watson, C. S. (2020). Present-Day Vertical Land Mo-  
524 tion of Australia From GPS Observations and Geophysical Models. *Journal of*  
525 *Geophysical Research: Solid Earth*, *125*(2). doi: 10.1029/2019JB018034
- 526 Rosen, P. A., Gurrola, E., Sacco, G. F., & Zebker, H. (2012). The InSAR scientific  
527 computing environment. In *Proceedings of the european conference on synthetic*  
528 *aperture radar, eusar* (Vol. 2012-April, pp. 730–733).
- 529 Simons, M., & Rosen, P. (2015). Interferometric Synthetic Aperture Radar Geodesy.  
530 In *Treatise on geophysics* (Vol. 3, pp. 339–385). Elsevier. Retrieved from  
531 <https://linkinghub.elsevier.com/retrieve/pii/B9780444538024000610>  
532 doi: 10.1016/B978-0-444-53802-4.00061-0
- 533 Tarayre, H., & Massonnet, D. (1996). Atmospheric Propagation heterogeneities  
534 revealed by ERS-1 interferometry. *Geophysical Research Letters*, *23*(9), 989–  
535 992. Retrieved from <http://doi.wiley.com/10.1029/96GL00622> doi: 10  
536 .1029/96GL00622

- 537 Weiss, J. R., Walters, R. J., Morishita, Y., Wright, T. J., Lazecky, M., Wang, H., ...  
538 Parsons, B. (2020). High-Resolution Surface Velocities and Strain for Anatolia  
539 From Sentinel-1 InSAR and GNSS Data. *Geophysical Research Letters*, *47*(17).  
540 doi: 10.1029/2020GL087376
- 541 Wright, T. J., Parsons, B. E., & Lu, Z. (2004). Toward mapping surface deforma-  
542 tion in three dimensions using InSAR. *Geophysical Research Letters*, *31*(1), 1–  
543 5. doi: 10.1029/2003GL018827
- 544 Xu, X., & Sandwell, D. T. (2020). Toward Absolute Phase Change Recovery  
545 with InSAR: Correcting for Earth Tides and Phase Unwrapping Ambiguities.  
546 *IEEE Transactions on Geoscience and Remote Sensing*, *58*(1), 726–733. doi:  
547 10.1109/TGRS.2019.2940207
- 548 Xu, X., Sandwell, D. T., Klein, E., & Bock, Y. (2021). Integrated Sentinel-1 InSAR  
549 and GNSS Time-Series Along the San Andreas Fault System. *Journal of Geo-  
550 physical Research: Solid Earth*, *126*(11), 1–14. doi: 10.1029/2021JB022579
- 551 Xu, Z.-W., Wu, J., & Wu, Z.-S. (2004). A survey of ionospheric effects on space-  
552 based radar. *Waves in Random Media*, *14*(2), S189-S273. Retrieved from  
553 <http://www.tandfonline.com/doi/abs/10.1088/0959-7174/14/2/008> doi:  
554 10.1088/0959-7174/14/2/008
- 555 Yu, C., Penna, N. T., & Li, Z. (2020). Ocean Tide Loading Effects on InSAR Ob-  
556 servations Over Wide Regions. *Geophysical Research Letters*, *47*(15). doi: 10  
557 .1029/2020GL088184
- 558 Yunjun, Z., Fattahi, H., & Amelung, F. (2019). Small baseline InSAR time se-  
559 ries analysis: Unwrapping error correction and noise reduction. *Computers and  
560 Geosciences*, *133*(May), 104331. Retrieved from [https://doi.org/10.1016/j  
561 .cageo.2019.104331](https://doi.org/10.1016/j.cageo.2019.104331) doi: 10.1016/j.cageo.2019.104331
- 562 Yunjun, Z., Fattahi, H., Pi, X., Rosen, P., Simons, M., Agram, P., & Aoki, Y.  
563 (2022). Range Geolocation Accuracy of C/L-band SAR and its Implications  
564 for Operational Stack Coregistration. *IEEE Transactions on Geoscience and  
565 Remote Sensing*, *2892*(c), 1–1. Retrieved from [https://ieeexplore.ieee  
566 .org/document/9759304/](https://ieeexplore.ieee.org/document/9759304/) doi: 10.1109/TGRS.2022.3168509
- 567 Zebker, H. A., & Villasenor, J. (1992). Decorrelation in interferometric radar echoes.  
568 *IEEE Transactions on Geoscience and Remote Sensing*, *30*(5), 950–959. doi:  
569 10.1109/36.175330

570 Zheng, Y., Fattahi, H., Agram, P., Simons, M., & Rosen, P. (2022). On  
571 Closure Phase and Systematic Bias in Multilooked SAR Interferometry.  
572 *IEEE Transactions on Geoscience and Remote Sensing*, 60, 1–11. Re-  
573 trieved from <https://ieeexplore.ieee.org/document/9758802/> doi:  
574 10.1109/TGRS.2022.3167648