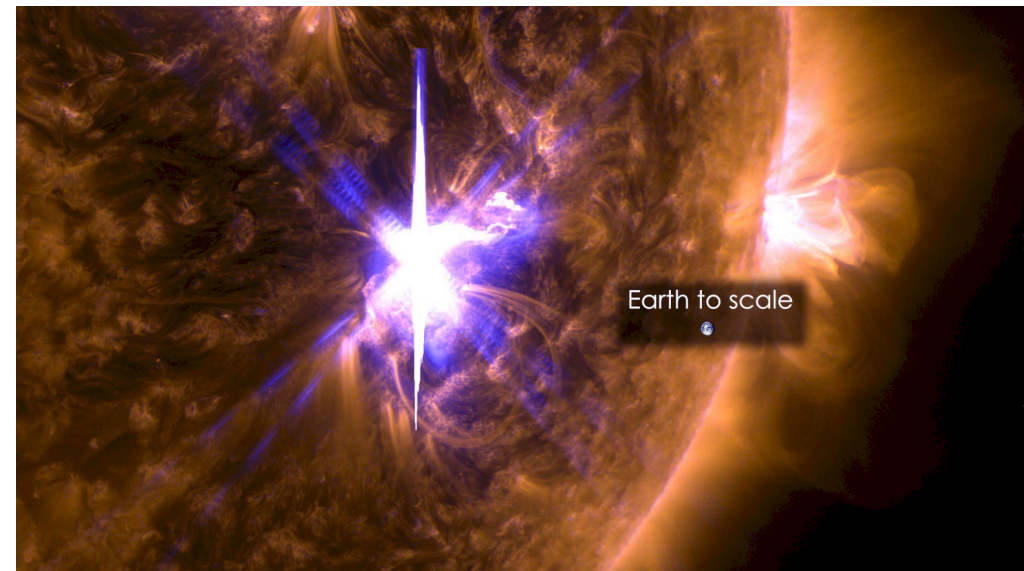


# Efficient Nonthermal Ion and Electron Acceleration in 3D Magnetic Reconnection

Qile Zhang+ 2021 PRL

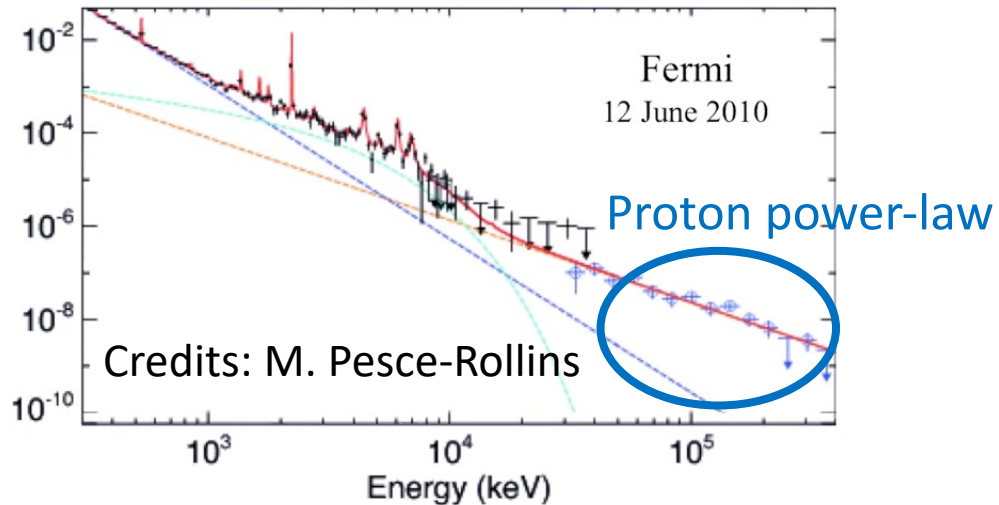
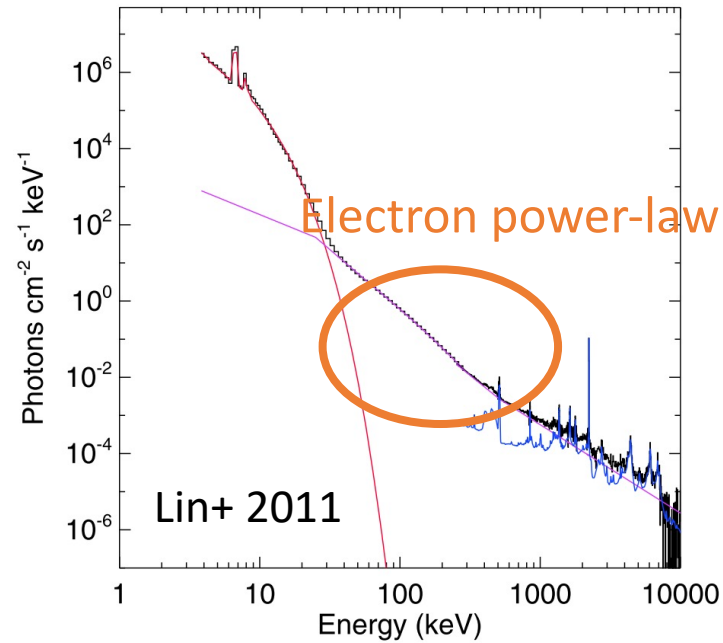
Qile Zhang, Fan Guo, Bill Daughton, Xiaocan Li, Hui Li

AGU Fall meeting 2021

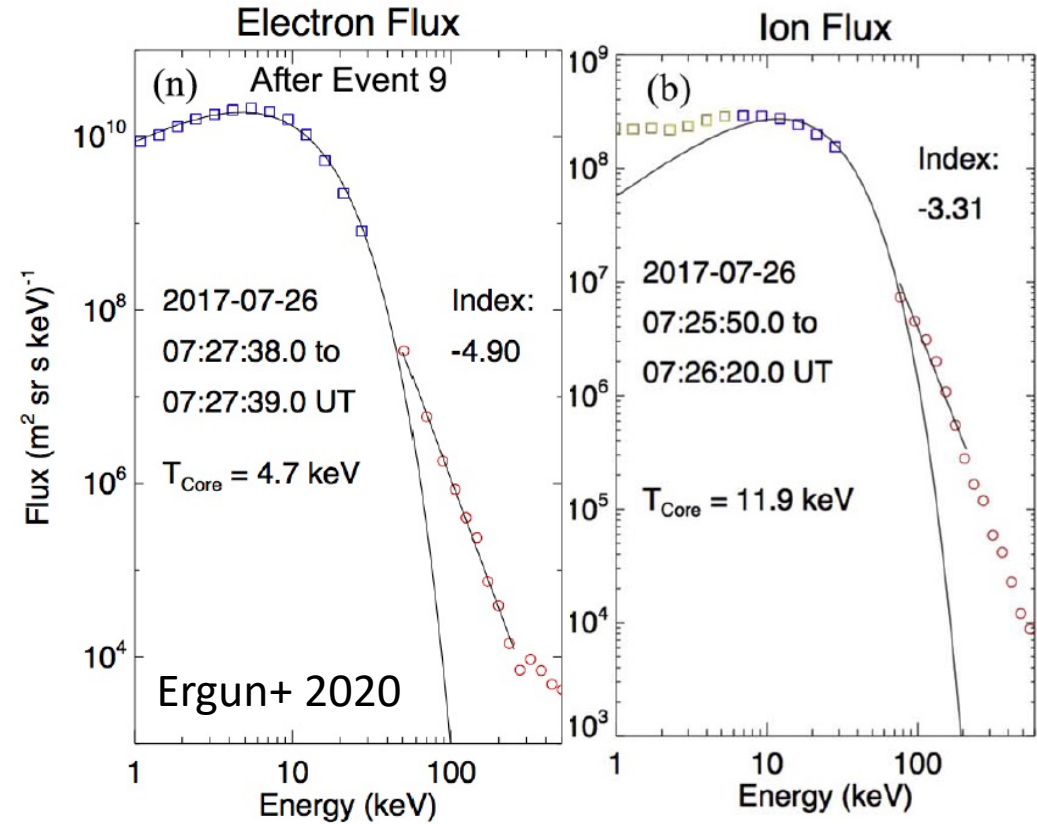


# Ions and electrons are efficiently accelerated into power-laws

## Solar flare remote sensing

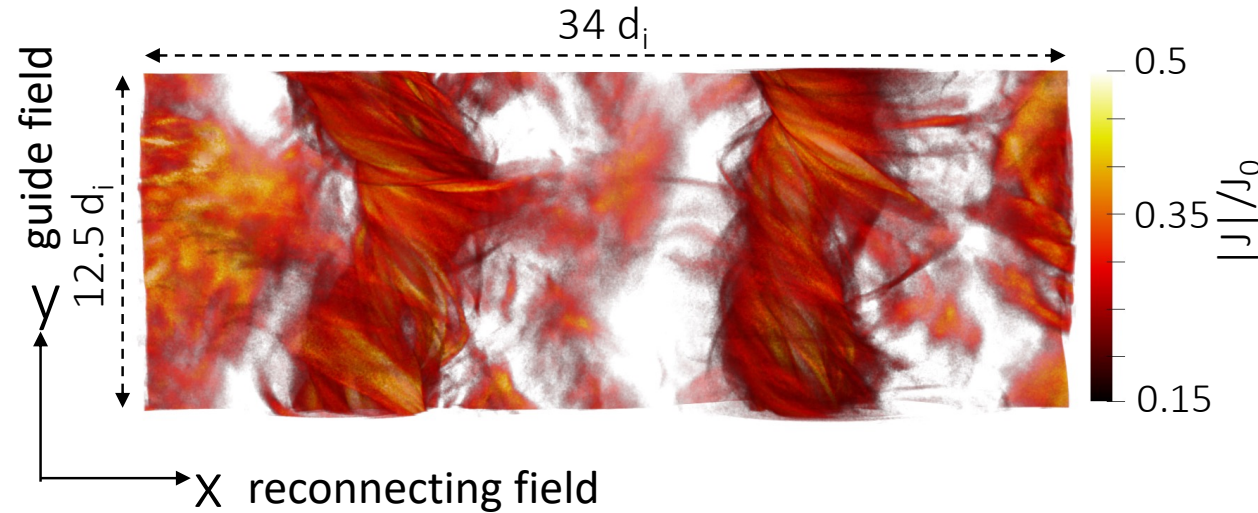


## Magnetotail in-situ



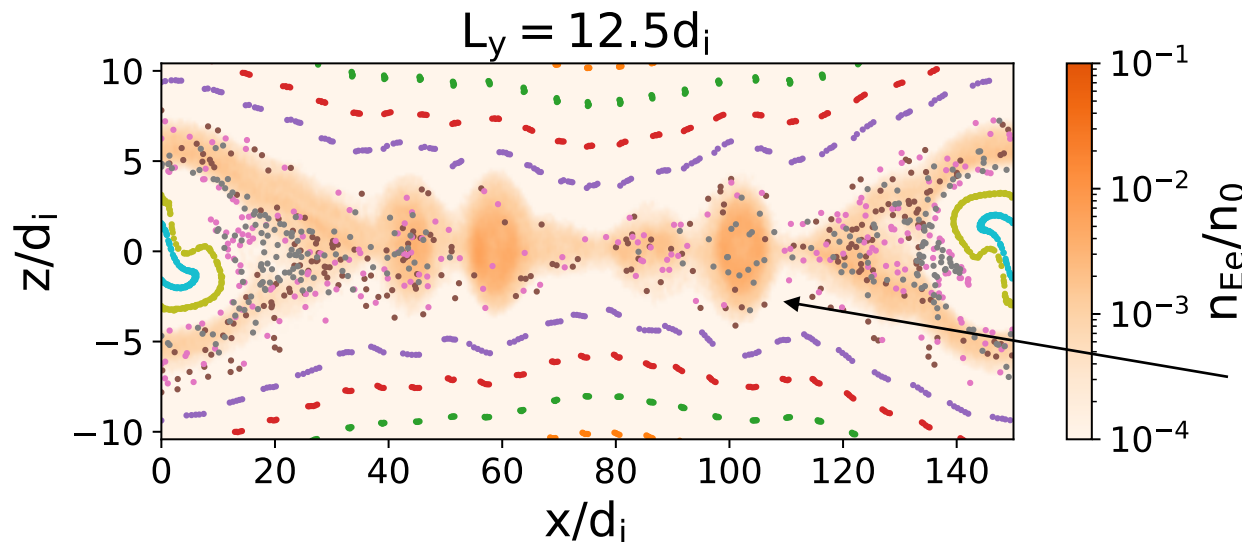
But the underlying physics is not well understood.

With low guide fields, flux-rope kink instability drives field-line chaos

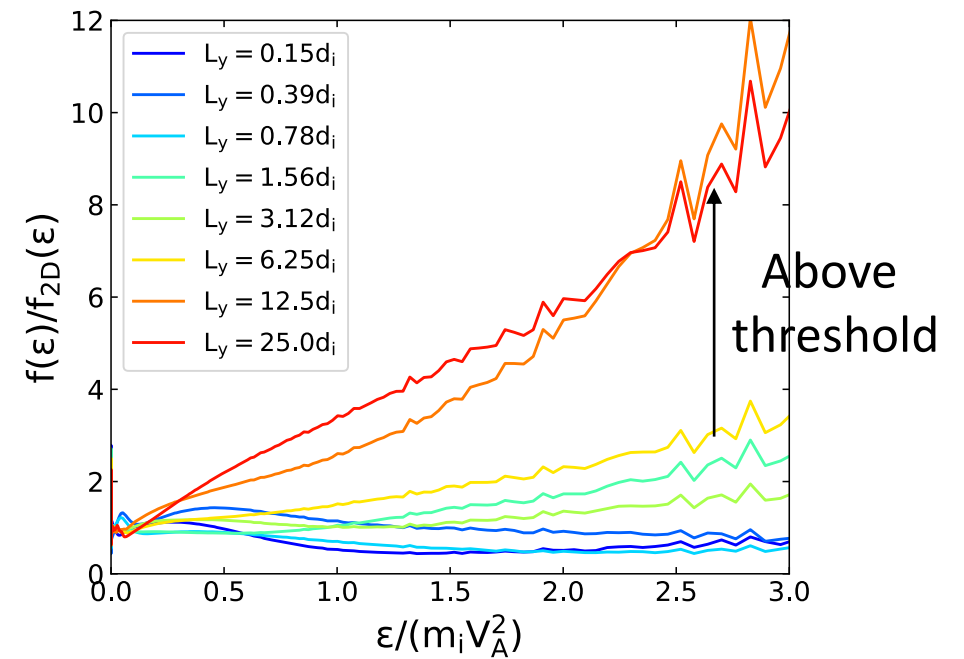
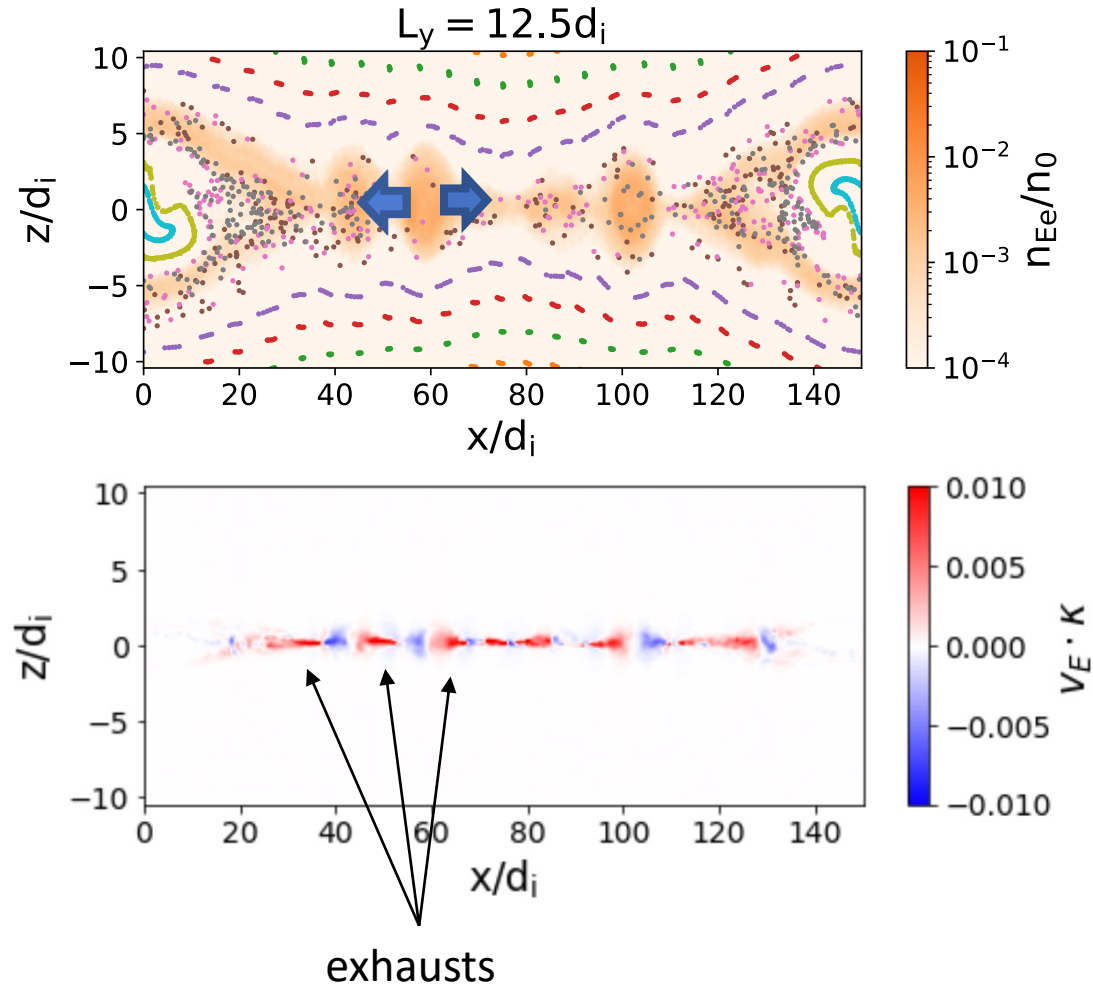


Kink instability criteria: safety factor  $q < 1$

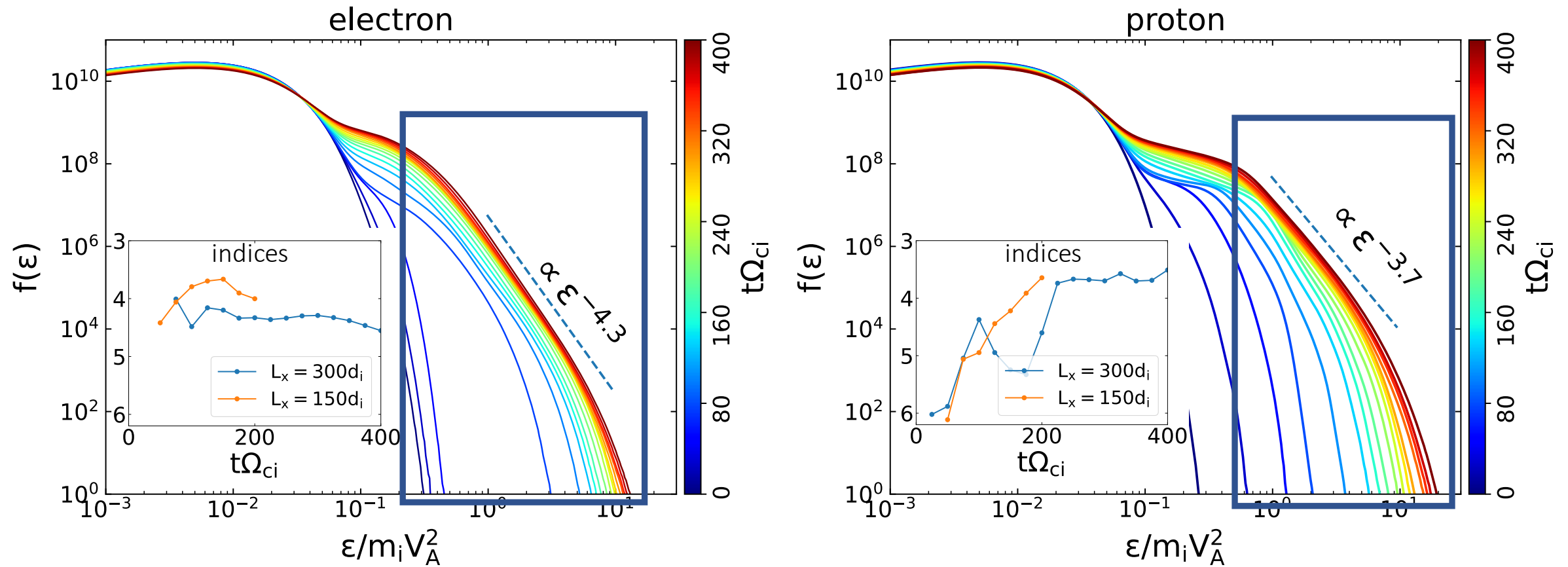
$$L_{th} \sim \pi b_g D \sim 0.1 \pi b_g L_x$$



# Field-line chaos leads to particle transport out of flux ropes, allowing more efficient acceleration

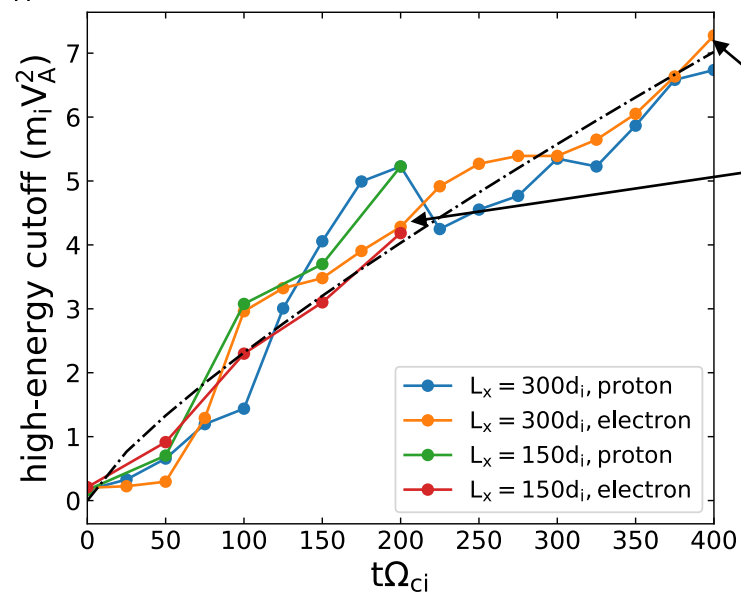
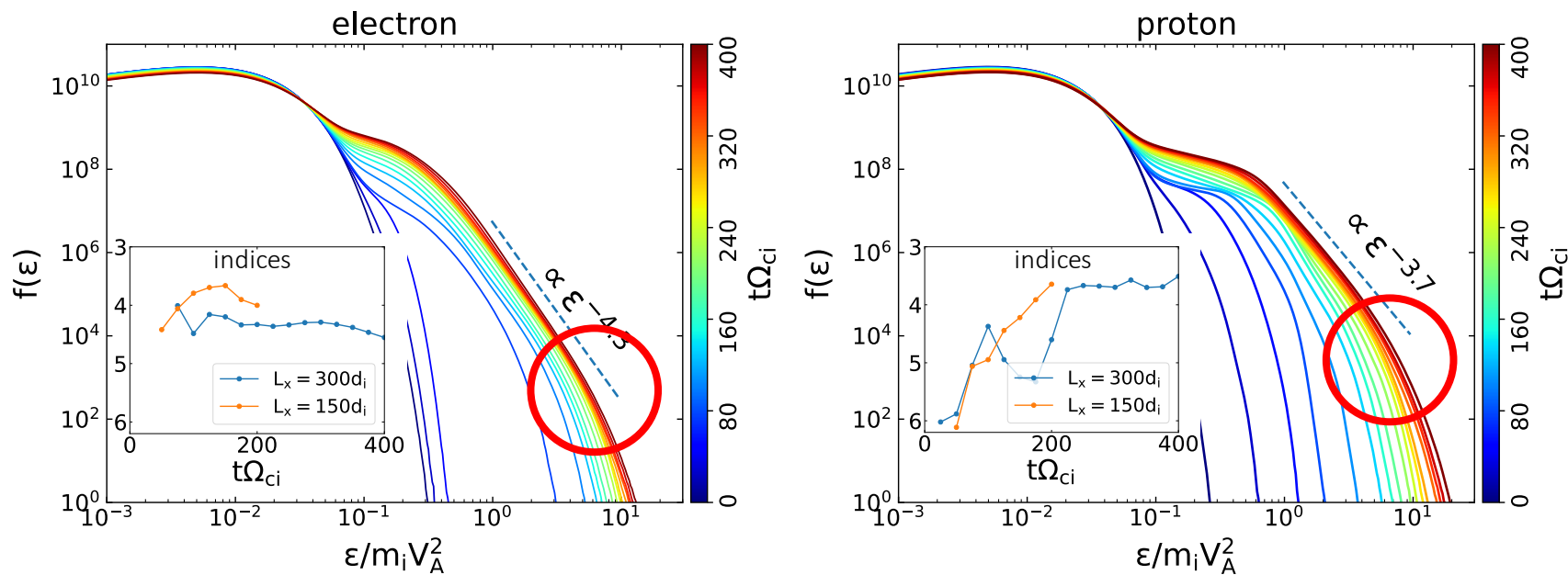


# Electrons and protons form power-laws containing a large fraction of energy



Nonthermal components:  $\sim 20\%$  number,  $50\%$  energy

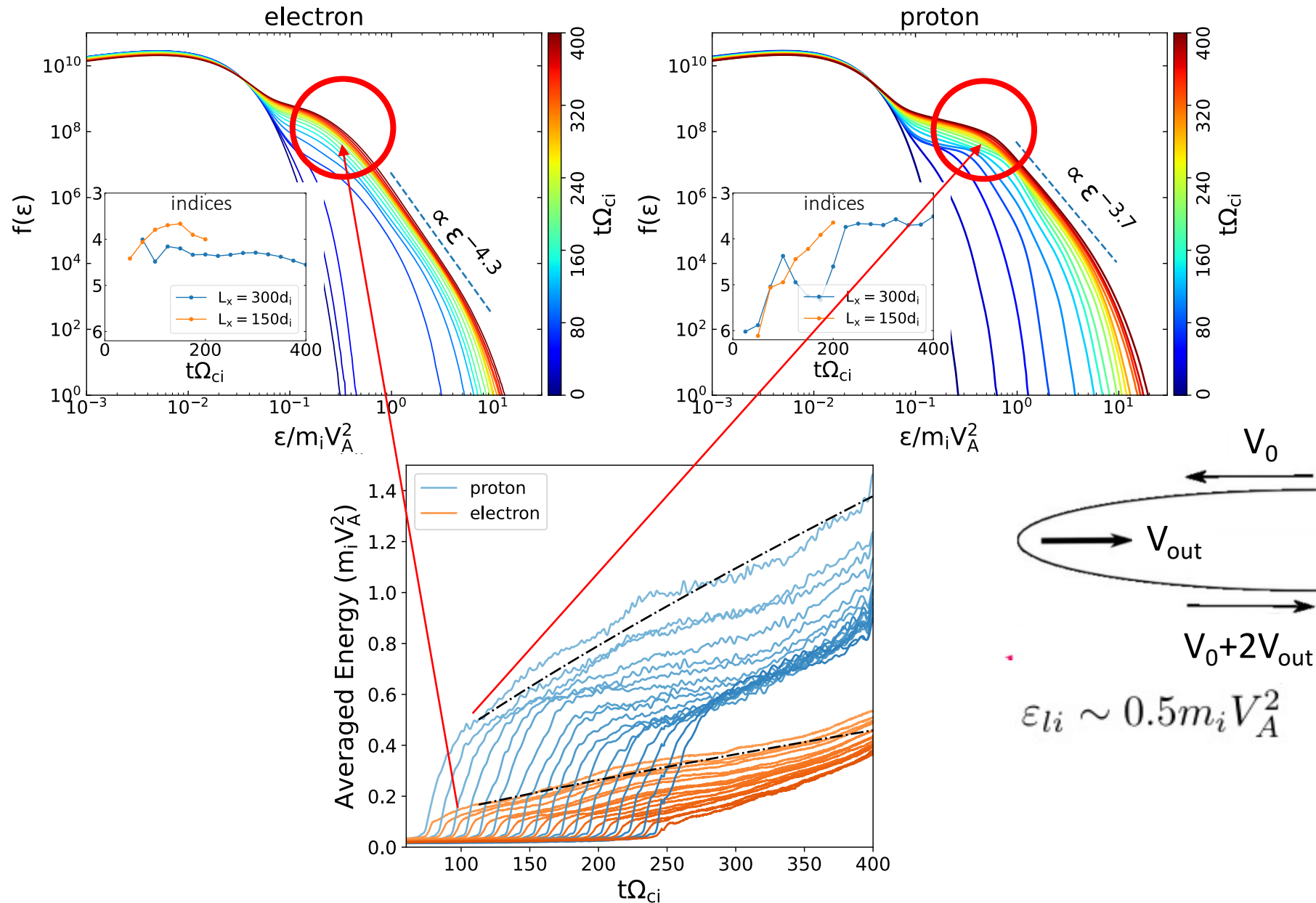
# High-energy cutoff is limited by system size



Increase with system sizes



# Low-energy bound is determined by injection



# The Fermi acceleration process can be elucidated by a scaling analysis

Fermi acceleration at exhausts + particle acceleration theory gives:

$$p \sim 1 + \frac{B_x}{B_z} \frac{\Delta_z}{L} \left(1 + \frac{B_g^2}{B_x^2}\right).$$

$$p \sim 4$$

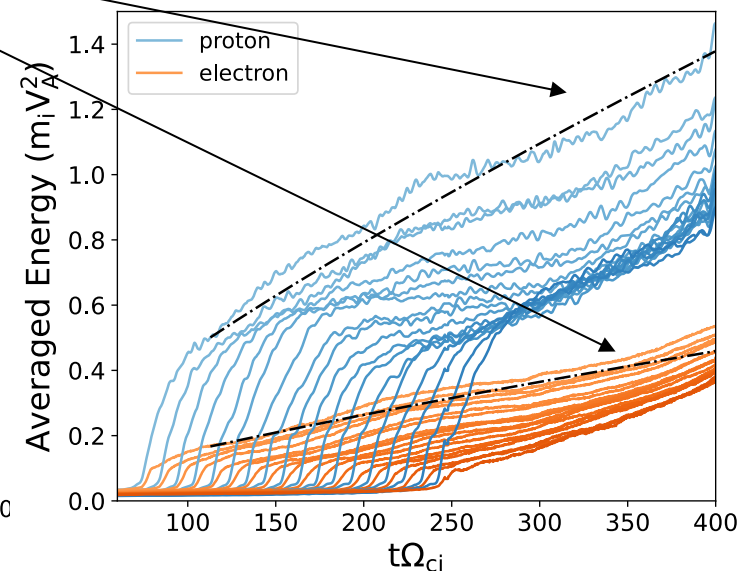
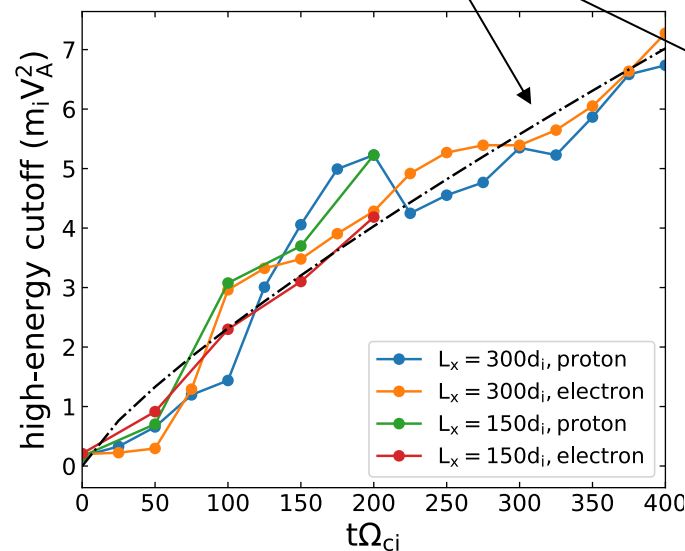
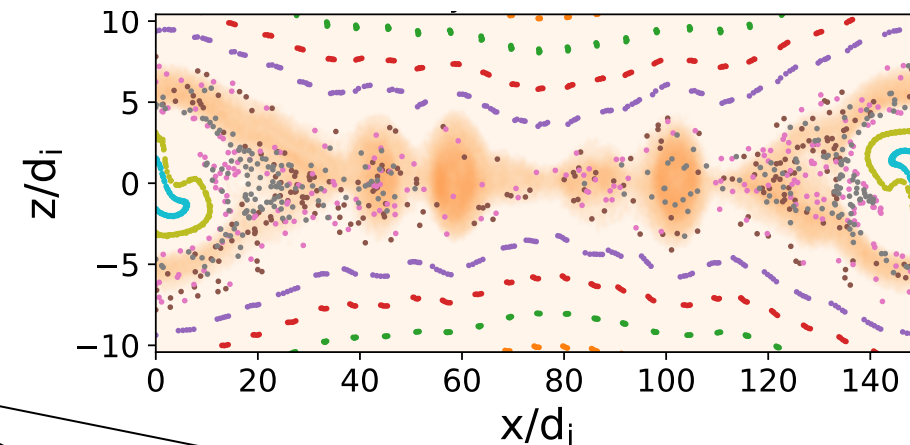
$$\alpha \equiv \dot{E}/E \sim \frac{B_z V_{Ax} B_x^2}{B_x (B_x^2 + B_g^2) \Delta_z}.$$

$$\alpha \sim C/t$$

$$C \sim 0.8$$

$$E \propto t^C \sim t^{0.8}$$

$$t^{0.8}$$





# Results are roughly consistent with observations

## Solar flares

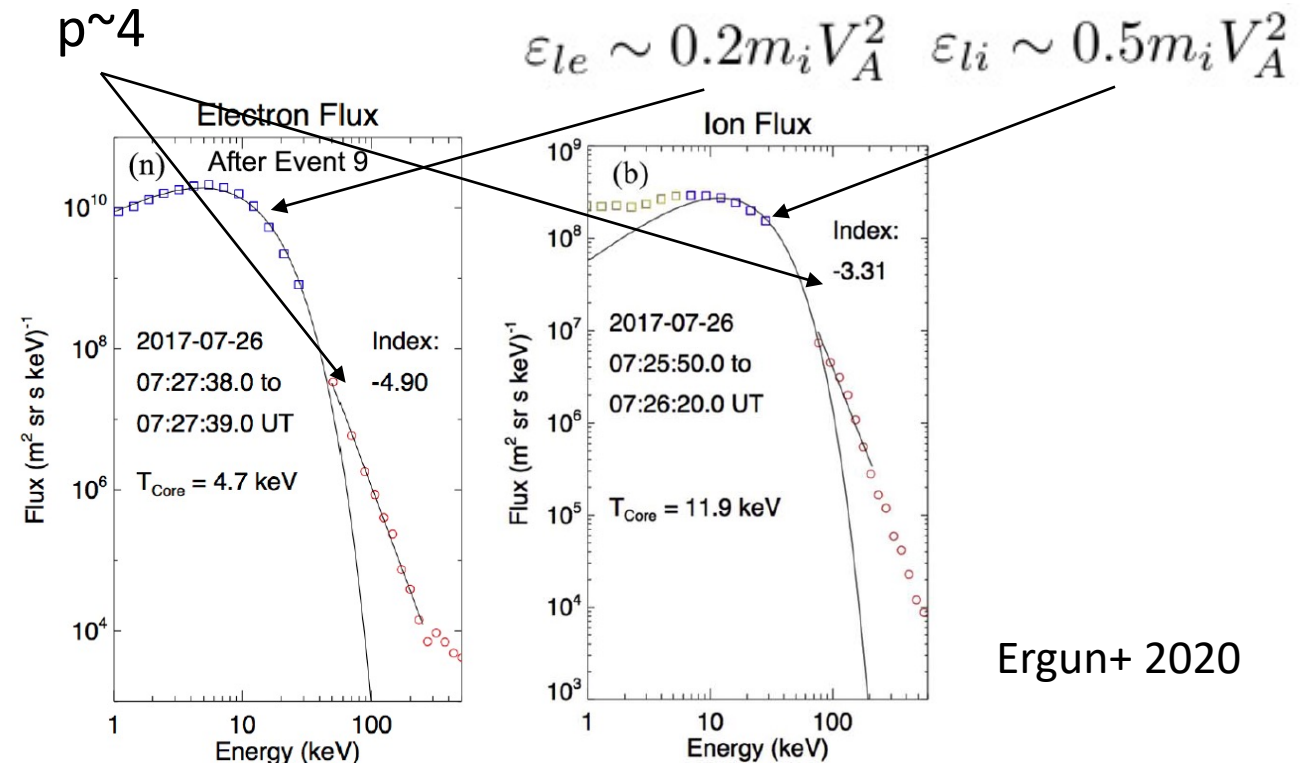
- Power-law indices  $p \sim 4$ : Electrons: x-ray (Oka+ 2018), microwave (Gary+ 2018)  
Protons: gamma-ray (Omodei+ 2018), solar energetic particles (Cohen+ 2020)
- Significant nonthermal fraction (Oka+ 2015, Aschwanden+ 2016)

## Magnetotail

In-situ spectra for both species

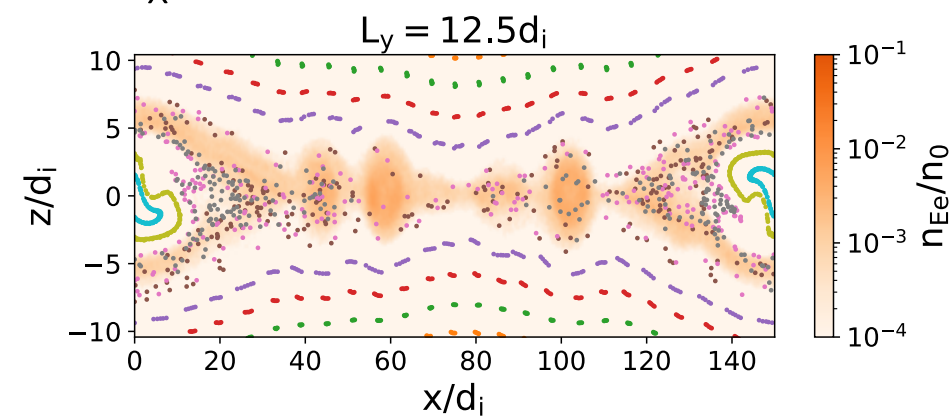
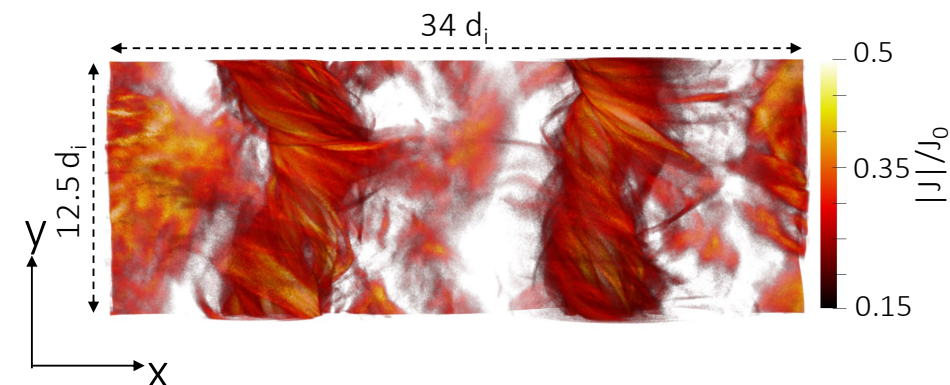
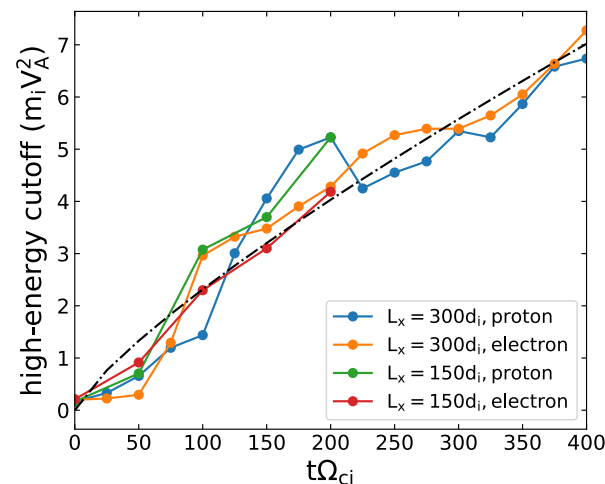
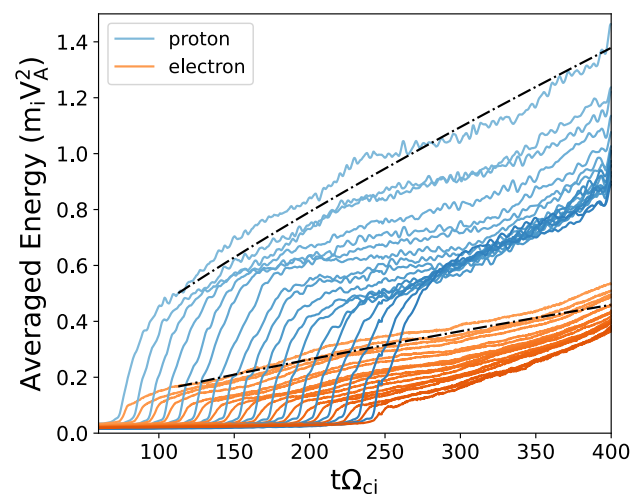
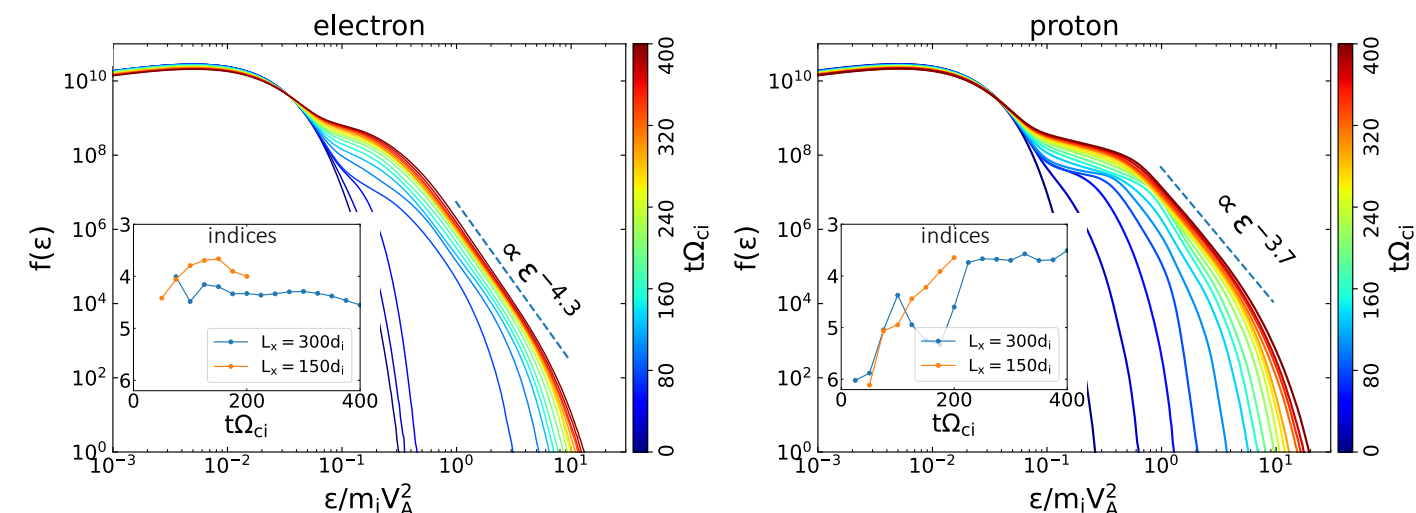
Power-law indices

Low-energy bound “shoulder”



# Take-aways

- 3D reconnection drives efficient nonthermal acceleration for both protons and electrons.
- Efficient Fermi acceleration is enabled by field-line chaos driven by flux-rope kink instability.
- Low-energy bound is determined by injection; high-energy cutoff grows with system size.



Qile Zhang+ 2021 PRL  
qlzhang@lanl.gov