

INVESTIGATING
GEOCORONAL
ABSORPTION
 FOR
WAVELENGTH
CALIBRATION

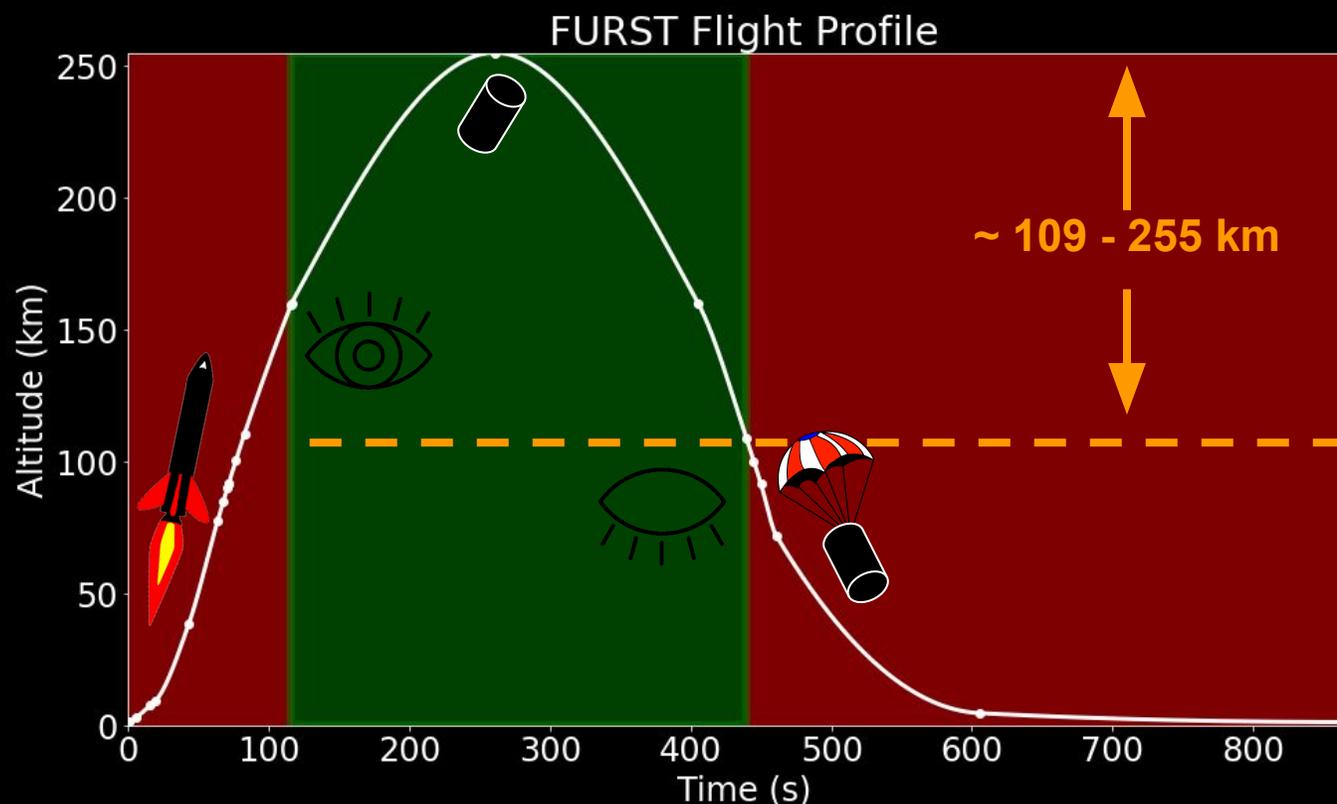
OF
SOUNDING
ROCKETS

NICOLAS DONDEES,
 AMY WINEBARGER,
 CHARLES KANKELBORG,
 GENEVIEVE VIGIL,
 LARRY PAXTON,
 & GARY ZANK

Sounding Rockets are spectroscopic and imaging instruments on-board sub-orbital flights

FURST will image the first **full-sun integrated high-resolution UV** spectra (1200-1810 Å)

- Current UV spectral measurement sources have a limited FOV (such as HRTS) or low resolution
- Will serve as a Hubble-analog



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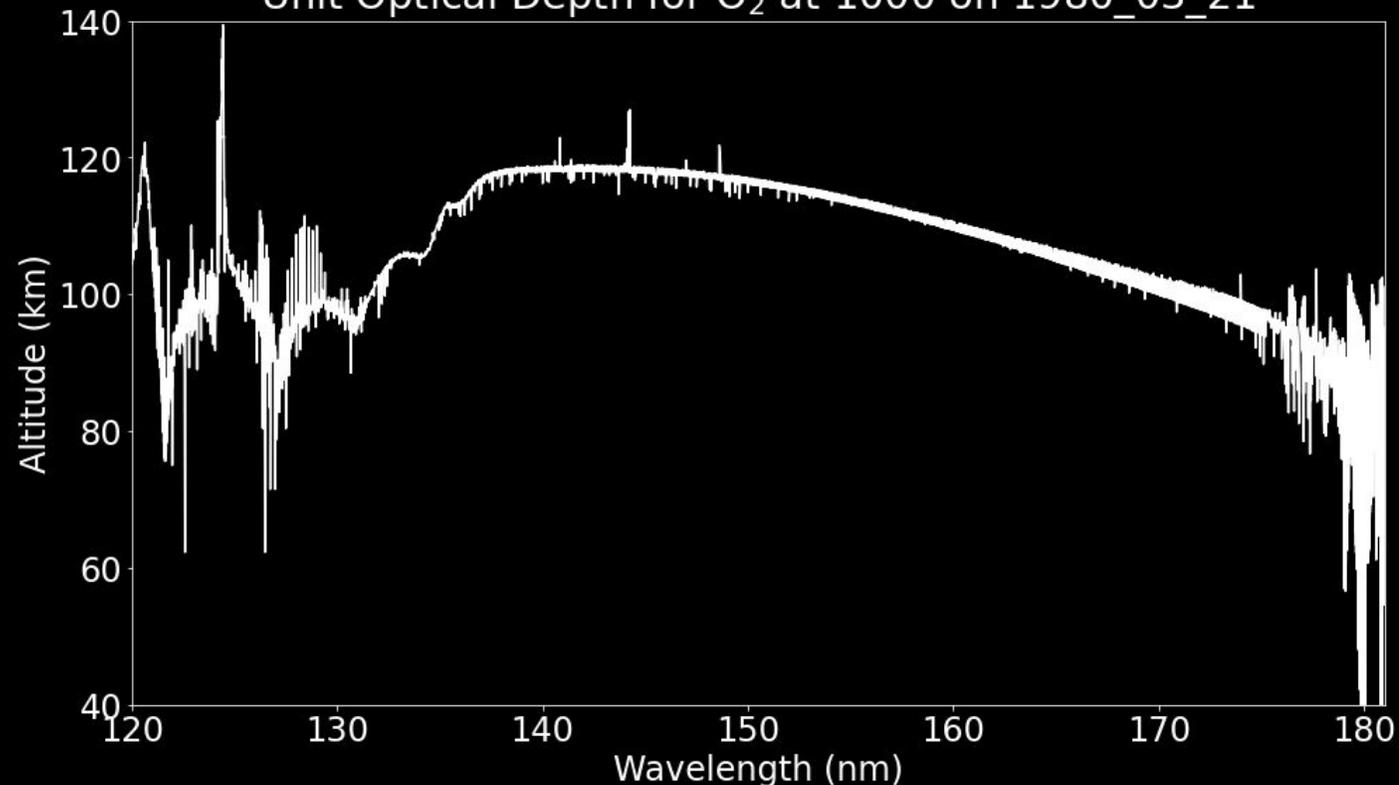
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Geocoronal absorption is caused by molecules in the upper atmosphere

Optical depth is the product of the **absorption cross section and number density** integrated vertically with altitude

$$\tau(\lambda, z) = \sigma(\lambda) \int_{z'}^{\infty} \eta(z) dz' = 1$$

Unit Optical Depth for O₂ at 1000 on 1980_03_21



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Absorption lowers the spectral
signal at known locations

$$I = I_0 e^{-\tau / \cos \theta}$$

These patterns can be useful
for **wavelength calibration!**

Additionally, we may also be
able to **validate atomic and
atmospheric properties**

HRTS Spectra with Geocoronal Absorption

