

How frequent are Antarctic sudden stratospheric warmings in present and future climate?

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Key Points:

- Antarctic sudden stratospheric warmings occur once every 22 years in present-day (1990) climate conditions.
- The warmings will become much rarer under future climate change, irrespective of their exact definition.
- The future decrease in frequency is linked to a strengthening of the Antarctic polar vortex.

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Abstract

Southern Hemisphere (SH) Stratospheric Sudden Warmings (SSWs) result in smaller Antarctic ozone holes and are linked to extreme midlatitude weather on subseasonal to seasonal timescales. Therefore, it is of interest how often such events occur and whether we should expect more events in the future. Here, we use a pair of novel multi-millennial simulations with a stratosphere-resolving coupled ocean-atmosphere climate model to show that the frequency of SSWs, such as observed 2002 and 2019, is about one in 22 years for 1990 conditions. In addition, we show that we should expect the frequency of SSWs—and that of more moderate vortex weakening events—to strongly decrease by the end of this century.

Plain Language Summary

The stratosphere at 10-50 km height can influence surface weather for several months. In 2002 and 2019, the stratosphere warmed over Antarctica within a few days to weeks. This caused dry and hot summers in Australia and South America. And it reduced the size of the ozone hole. Since these warming events are rare, it is difficult to say how often they occur. We therefore use long computer simulations to answer that question. We find that without climate change, warming events occur about every 22 years. But with climate change, the warming events will happen only once every 300 years. From this, we believe that the quick succession of two events in 2002 and 2019 will remain special in history.

1 Introduction

The stratospheric polar vortex forms in the winter hemisphere due to the lack of solar heating at high latitudes and the resulting strong equator-to-pole temperature gradient. In the Northern Hemisphere (NH), strong and planetary scale waves originating in the troposphere from orographic forcing and land-sea contrast periodically propagate upward into the stratosphere and perturb the polar vortex via momentum deposition when the waves break (Eliassen & Palm, 1960; Charney & Drazin, 1961; Matsuno, 1971). In extreme cases, this disruption of the polar vortex leads to a rapid warming and reversal of wind directions in the polar stratosphere, a so-called (major) Sudden Stratospheric Warming (SSW) (Butler et al., 2015). These SSWs occur around every other winter in the NH.

However, over the six decades that we have station records (and later satellite observations) of the Southern Hemisphere (SH) polar vortex, only one such wind reversal has been recorded in 2002 (Roscoe et al., 2005; Esler et al., 2006). This event substantially decreased the size of the ozone hole thanks to higher than usual stratospheric polar temperatures and transport of ozone-rich air from lower latitudes into the polar regions (Fig. S2a) (Stolarski et al., 2005). There was also a dynamical effect of the 2002 SSW at the surface, as an extreme negative polarity of the Southern Annular Mode (SAM) was recorded at the surface for the 10-90 day period following the event (Thompson et al., 2005). Even though no wind reversal at 60° S and 10 hPa was registered in 2019, the polar vortex in this more recent event weakened dramatically and also led to a smaller ozone hole (Fig. S2b) with almost 30% higher total column ozone values compared to the previous decade (Safieddine et al., 2020). The event has also been linked to the severe bushfire season in South Eastern Australia the following spring and summer (Lim et al., 2021).

Due to the impacts on stratospheric ozone and surface weather on the subseasonal to seasonal timescale, it is important to determine how rare SSWs are in the SH, and whether we should expect more or less frequent SSWs under future climate change. However, given the shortness of the observational record it is impossible to get an observa-

65 tional estimate of how often SSWs do occur on average. Recently, Wang et al. (2020)
 66 analyzed hindcasts of a seasonal forecasting system and found an average Antarctic SSW
 67 frequency of one every 25 years. However, the underlying model of this study had a strong
 68 mean westerly wind bias, raising some doubts on the validity of their results. Here, we
 69 revisit the question of how frequent Antarctic SSWs are in present climate and also ad-
 70 dress possible changes under future climate change. This is accomplished by investigat-
 71 ing two nearly 10,000-year long simulations with a well-performing stratosphere-resolving
 72 coupled ocean-atmosphere model based on present-day (1990) and future (increased CO₂)
 73 conditions and by considering integrations from the sixth Climate Model Intercompar-
 74 ison Project (CMIP6).

75 **2 Model data and SSW definitions**

76 **2.1 Multi-millennial coupled GCM simulations**

77 We use a set of two 9,900-year long simulations with the stratosphere-resolving ver-
 78 sion of the the Geophysical Fluid Dynamics Laboratory’s CM2.1 atmosphere-ocean cou-
 79 pled climate model (Delworth et al., 2006; Horan & Reichler, 2017), which has been used
 80 in particular for studies of stratosphere-troposphere coupling in the past (Horan & Re-
 81 ichler, 2017; Jucker & Reichler, 2018). The model has 48 vertical levels with approxi-
 82 mately half of the levels situated in the stratosphere and a model top at 0.002 hPa. The
 83 horizontal resolution is approximately 2° in latitude and 2.5° in longitude. The bound-
 84 ary conditions are set to perpetual 1990 conditions. More specifically, ozone in the year
 85 1990 is comparable to both 2002 and the 2010s (Newman & Nash, 2019). The two sim-
 86 ulations differ in their greenhouse gas forcing; CO₂ is set to 353 ppm in the ‘present-day’
 87 and 1120 ppm in the ‘future’ simulation, which is a quadrupling relative to pre-industrial
 88 CO₂ concentration (and 3.2 times present-day concentration). This is the only difference
 89 between the two simulations. Atmospheric variables are stored on a daily frequency to
 90 allow for detailed dynamical analysis, including Eliassen-Palm fluxes.

91 In agreement with Horan and Reichler (2017), who have shown that this model com-
 92 pares well to reanalysis in the troposphere and northern hemisphere stratosphere, both
 93 the southern hemisphere stratospheric zonal mean zonal wind and vertical component
 94 of the Eliassen-Palm flux from our present-day simulation show excellent agreement with
 95 those from ERA5 reanalysis (1979-2019) (Hersbach et al., 2020), for both mean and stan-
 96 dard deviation (Figs. 1a,c and S1). We also note that the model intercomparison work
 97 by Reichler and Kim (2008) showed that CM2.1 had the best performance index among
 98 CMIP3 models, even though that version had only half the number of vertical levels com-
 99 pared to the version used here. Besides its performance in the atmosphere, which is of
 100 particular relevance here, the oceanic component has been validated extensively and also
 101 found to have a good representation of tropical (including ENSO, Wittenberg et al., 2006)
 102 as well as extratropical southern hemisphere ocean dynamics (Gnanadesikan et al., 2006).

103 Having multi-millennial simulations with a model showing such small bias will al-
 104 low us to robustly estimate SSW frequencies. In addition, having future projections will
 105 make it possible to address the question of whether or not we should expect another SSW
 106 to occur in the future, and we will show that increased greenhouse gas concentrations
 107 have a strong impact on SSW frequency.

108 **2.2 SSW definitions**

109 We follow the most common definition of Sudden Stratospheric Warmings as the
 110 reversal of u_{1060} , the zonal mean zonal wind at 60°S and 10 hPa (‘SSW-reversal’, Charl-
 111 ton & Polvani, 2007). However in observations, only the September 2002 event is an SSW-
 112 reversal event, while the 2019 event is widely considered an SSW but did not show wind
 113 reversal at 60°S and 10 hPa. Therefore, we have performed our analysis with an addi-

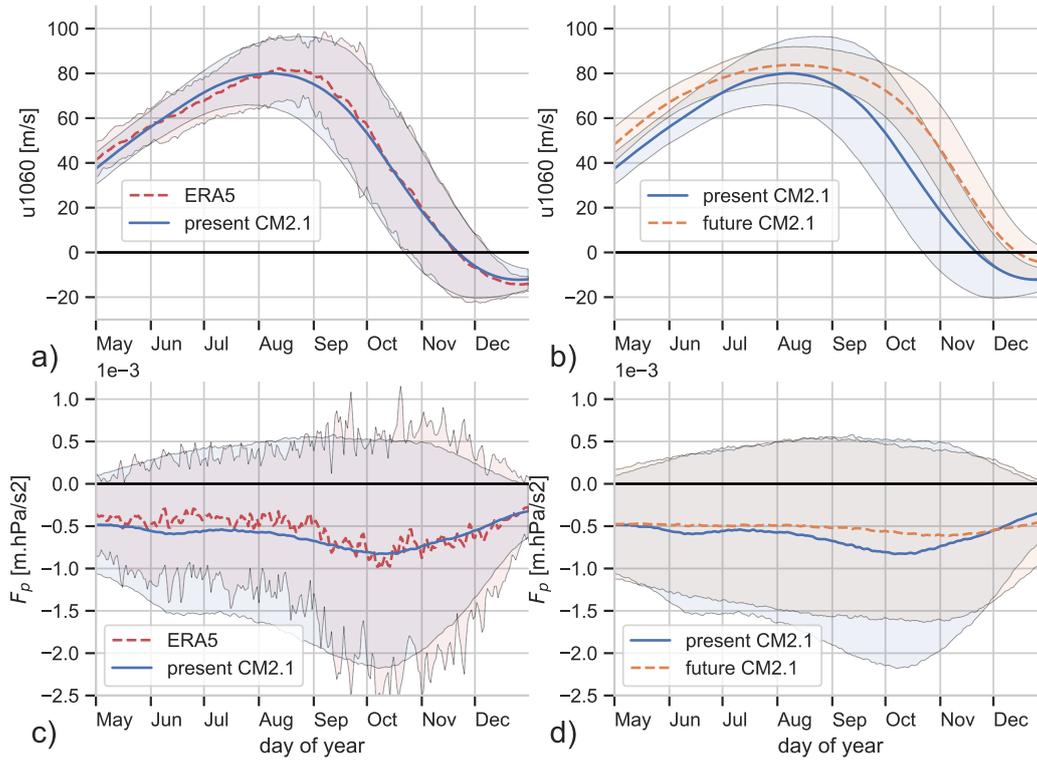


Figure 1. (top) Climatological mean (solid) and two interannual standard deviations (shaded) of zonal mean zonal wind at 60°S and 10 hPa (u_{1060}) for (a) present-day CM2.1 and ERA5 and (b) present-day and future CM2.1. (bottom) same but for vertical EP flux. The present-day simulation (blue, solid) reproduces both mean and variability of the ERA5 reanalysis (1979-2019; red, dashed) in both u_{1060} (a) and vertical EP flux (c). The future simulation (orange, dashed) shows a clear strengthening of the polar vortex throughout the year (b) and a weakening of the vertical EP flux (d), in particular during the spring.

114 tional definition, allowing for a more general determination of SSW frequency and fu-
115 ture change.

116 We found that the simplest method to define SSWs in the SH which detects both
117 2002 and 2019 as the only events during the satellite era is that the zonal mean zonal
118 wind anomaly with respect to the day of the year at 60°S and 10 hPa passes below -40 m/s.
119 The onset date is then defined as the day when the zonal mean zonal wind anomaly crosses
120 -20 m/s for the last time before crossing -40 m/s. These ‘SSW-weak’ events follow the
121 common features of stratosphere-troposphere coupling in the SH in their significant sur-
122 face impact on monthly timescales (Fig. S3).

123 For both definitions, two events have to be separated by at least 20 days, and the
124 onset date has to be at least 20 days before the vortex breakdown, which is defined as
125 the last day of the year when u_{1060} becomes negative.

126 Finally, we follow Lim et al. (2018) who showed that weaker events can also have
127 an impact at the surface, and we will also report results from their detection method based
128 on the yearly timeseries of the first Principal Component of de-seasonalized monthly mean
129 zonal mean zonal wind between 55 and 65°S. The corresponding Empirical Orthogonal
130 Function is two-dimensional but in month of the year–pressure space (instead of the con-
131 ventional longitude–latitude space) and is centered around the vortex breakdown in spring
132 (the ‘L18’ method). This method does not provide onset dates, as there is only one value
133 per year, and L18 is closely related to variations in the date of the vortex breakdown (pos-
134 itive for earlier breakdown; the correlation coefficient between the first Principal Com-
135 ponent and the vortex breakdown date is $r = 0.79$ in ERA5 data, not shown). Follow-
136 ing Lim et al. (2019), we apply a threshold of 0.8 standard deviations, which detects many
137 more events than the other two definitions.

138 3 Occurrence of SSWs in the Southern Hemisphere

139 The present-day 9,900-year simulation produces 458 SSW-weak and 159 SSW-reversal
140 events, corresponding to an average frequency of about one SSW-weak every 22 years and
141 one SSW-reversal every 59 years. This compares well with the single SSW-reversal and
142 only two SSW-weak events in the 42-year long satellite observation record (and the 63-
143 year long non-satellite observational record since 1957 (Roscoe et al., 2005; Naujokat &
144 Roscoe, 2005)), as well as Wang et al. (2020). In addition to yearly occurrence, we also
145 analyze the seasonal occurrence of SSWs and find that the SSW-weak criterion detects
146 events during the entire winter, with a peak occurrence in late August to September (Fig. 2d)
147 and a mean occurrence of 27 August (note that early events in June and July have a sim-
148 ilar impact to later events, not shown). The 2002 SSW occurred in late September, a
149 time of the year when we estimate the mean return time of SSW-weak events to be 113 years,
150 and the 2019 SSW occurred in early September, when the mean return time is estimated
151 to be 102 years (Fig. 2a). Irrespective of time of the year, our present-day simulations
152 indicate that we should expect between 0 and 6 SSW-reversals and between 0 and 12
153 SSW-weak events per century, with most likely numbers of 0-2 SSW-reversal and 3-6 SSW-
154 weak events per century (25th and 75th percentiles, Figs. 2b,e). As indicated before, L18
155 events are much more abundant, with an occurrence of 7-36 events per century and a
156 mean return time of one in 5 years (Fig. 2h).

157 To get an estimate of when the next SSW might occur, we perform a return time
158 analysis where we produce a histogram of the number of SSWs which occur within a given
159 time interval (Fig. 2c,f,i). If SSWs are independent and random events, we can compare
160 the observed return time distribution to a theoretical distribution (Text S6). The return
161 time histogram follows closely the theoretical distribution for all methods, suggesting that
162 in the SH, SSWs are independent and random, with a mean return time of about 59 years
163 for SSW-reversal and 22 years for SSW-weak, or an annual probability of occurrence of

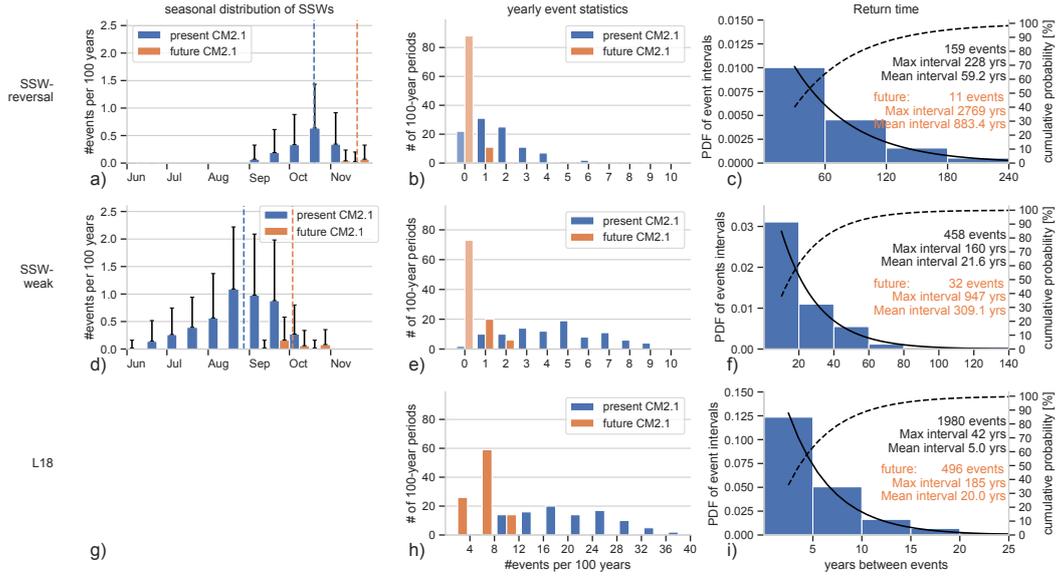


Figure 2. Event statistics: (left) Seasonal distribution, (middle) histogram of number of events per century, and (right) return time distribution histograms (bars) and theoretical distribution (black lines) for probability (solid) and cumulative distribution functions (dashed). Statistics are shown for (top) SSW-reversal, (middle) SSW-weak and (bottom) L18. For all plots, the present-day simulation is in blue and increased CO₂ (‘future’) in orange. On the left panels, statistics are shown for half-monthly intervals, the black whiskers show the standard deviation, and the vertical dashed lines indicate the mean date of occurrence. Panel (g) is empty as there is no seasonal information for L18. Note the differences in scales between rows. In panels (b) and (e), bars are drawn for each year, whereas in panel (h), the bars are drawn within intervals designated by the tick marks. Bars showing the number of centuries without event are pale.

164 1.6% for SSW-reversal and 4.6% for SSW-weak. Using the theoretical survival function,
 165 we can then compute the probabilities of various scenarios (reported in Table 1). All of
 166 these probabilities are consistent with the observational record of one SSW-reversal and
 167 two SSW-weak events during the satellite era. Finally, neglecting any changes in climate
 168 from further greenhouse gas forcing since 1990, we estimate from the present-day sim-
 169 ulation that the probability of at least one SSW by the end of the century (next 80 years)
 170 would be 74% for SSW-reversals and 98% for SSW-weak events. Of course, this is only
 171 hypothetical as greenhouse gas concentrations have already risen since 1990 and are pro-
 172 jected to further increase in the future.

173 **4 Enhanced greenhouse gas forcing**

174 To estimate the impact of enhanced greenhouse gas forcing on the occurrence of
 175 SSWs in the SH, we conducted a second 9,900 year long simulation using increased CO₂
 176 corresponding to the end of the century (1120 ppm instead of 353 ppm, henceforth called
 177 ‘future’). The occurrence of SSWs in this simulation decreases drastically. The number
 178 of SSW-reversals reduces from 159 SSWs for present-day to only 11 in the future sim-
 179 ulation, while SSW-weak events decrease from 458 to only 32 (Fig. 2). This translates
 180 into a return time of one SSW-reversal every 883 and one SSW-weak every 309 years,
 181 and a maximum of 1 SSW-reversal and 2 SSW-weak events per century. Note how the
 182 most probable outcome by far for any given 100-year period is zero SSWs (median is zero
 183 for both SSW-reversal and SSW-weak; Fig. 2b and e, orange). From the theoretical fit,
 184 the probability of occurrence of at least one SSW-weak event in 80 years is now about
 185 23% (2.8% for at least two SSWs; Table 1). The analysis also suggests that SSW-reversals
 186 become very rare (probability of 8.7% within 80 years). SSWs not only become much rarer,
 187 but are also occurring later in the year, with a mean date of 3 October for SSW-weak,
 188 i.e. more than one month later than in the present-day simulation. For all definitions,
 189 there is a strong tendency for fewer SSWs in the future—including L18, which reduce to
 190 0-11 events per century. Thus, while the 2019 event is consistent with the occurrence rate
 191 in our present-day simulation, it is inconsistent with the rate seen in our future simu-
 192 lation. Given the trend in SSW frequency, and that we are already one-third of the way
 193 towards the year 2080 (when the greenhouse gas concentrations are projected to reach
 194 the levels of our future simulation), we conclude that this latest event should not be at-
 195 tributed to increased CO₂ forcing, and might indeed be the last observed event this cen-
 196 tury.

197 The decrease in SSW frequency in the future is accompanied by a strengthening
 198 of the SH polar vortex (Fig. 1b), which can be linked to stronger radiative cooling un-
 199 der increased greenhouse gas concentrations (Thompson et al., 2012; Santer et al., 2013).
 200 In addition, our simulations suggest a decrease in wave forcing, more so during spring
 201 than other times of the year (Fig. 1d). Together with an earlier study, which found a di-
 202 rect link between the SSW-reversal frequency and polar vortex strength (Jucker et al.,
 203 2014), our results suggest that the projected strengthening of the polar vortex along with
 204 a decrease in wave forcing are responsible for a substantial decrease in the probability
 205 of occurrence of SSWs.

206 **5 Comparison to NH**

207 The occurrence of SSWs in the NH is very different from the SH, not just because
 208 of the much higher SSW frequency at present, but also in terms of future projections of
 209 both polar vortex strength and SSW frequency. As discussed in detail by Horan and Re-
 210 ichler (2017), our model climatology and variability in the NH compares well to reanal-
 211 ysis products (Fig. 3), and it produces about five SSWs per decade in the NH, in accor-
 212 dance with observations (Jucker & Reichler, 2018). Therefore, we perform the same anal-
 213 ysis for the NH and briefly report our findings here.

		SSW-weak	SSW-reversal
present	Yearly probability	4.6%	1.6%
	Probability of less than observation	43%	52%
	Probability of exact observation	28%	35%
	Probability of more than observation	30%	15%
	Probability > 50% after	15 years	41 years
future	Yearly probability	0.3%	0.1%
	Probability > 50% after	214 years	612 years
	Probability of at least one SSW in 80 years	23%	8.7%
	Probability of at least two SSWs in 80 years	2.8%	0.4%

Table 1. Results from the theoretical fitting of the return times (Figs. 2c and f). Yearly probability is the probability of an event occurring during any given year ($1/\text{mean return time}$), probability of exact observation is computed for 2 SSW-weak and 1 SSW-reversal in 41 years. Time periods give the interval after which an SSW is more probable than not (probability of one or more events $> 50\%$). The labels ‘present’ and ‘future’ refer to the relevant CM2.1 simulations, and we use an 80-year period to compare to the time span 2021-2100 in the future simulation, but noting that this has CO_2 concentrations that are more representative of the end of the 21. century. Note that the observation percentages in the present simulation add to 101 instead of 100 due to rounding errors.

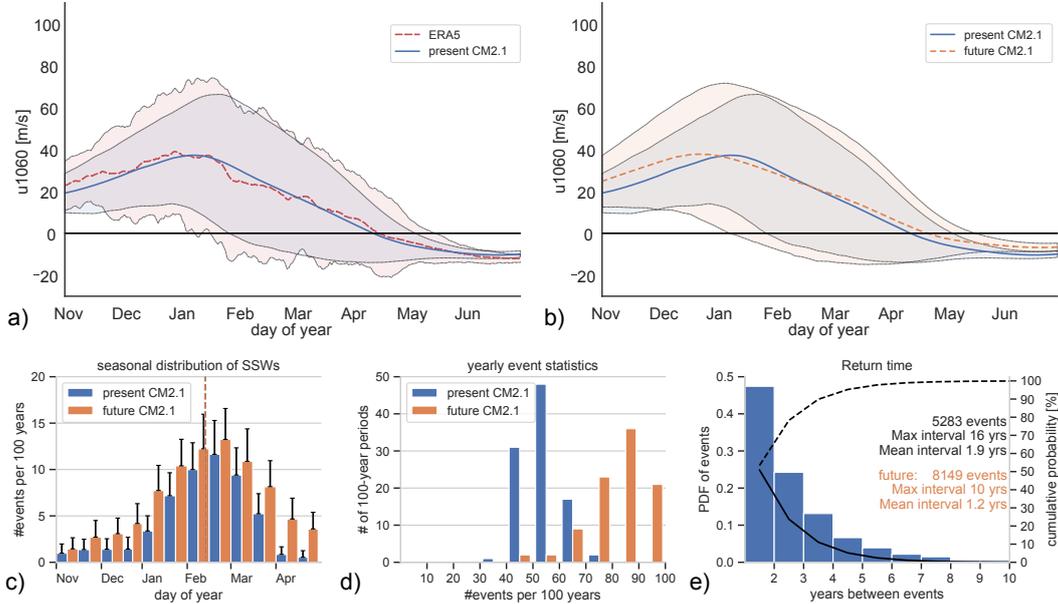


Figure 3. (top) u_{1060} for the Northern Hemisphere (NH) for (a) present-day and (b) increased CO_2 (‘future’), similar to Fig. 1. (bottom) NH SSW-reversal statistics for (c) seasonal distribution, (d) number of events per century and (e) return time, similar to Fig. 2. Note the differences in scale of the bottom row compared to Fig. 2, which is a result of the higher occurrence rate for the NH.

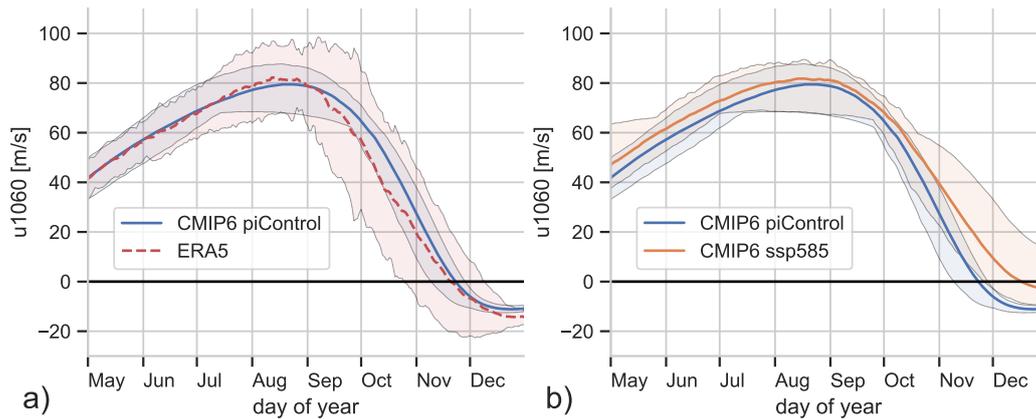


Figure 4. Analysis similar to Fig. 1, but using CMIP6 data. SSP585 data represents the climatology over 2080-2100. Shading corresponds to the range of model means (min to max) and the thick lines the multi model means. piControl is shown in blue, and SSP585 in orange, similar to Fig. 1.

214 The return time distribution shows that at intervals shorter than four years, NH
 215 SSWs are not independent and random (Fig. 3e), probably reflecting the influence of slowly
 216 evolving large scale climate modes, such as the El Niño Southern Oscillation or the Quasi-
 217 Biennial Oscillation, on the occurrence of SSWs (Holton & Tan, 1980; Taguchi & Hart-
 218 mann, 2006; Anstey & Shepherd, 2014). The NH polar vortex is also weaker and more
 219 influenced by upward propagating planetary waves from the troposphere, resulting in a
 220 more variable polar vortex than in the SH (Fig. 3, top). Our simulations suggest a slightly
 221 weaker polar vortex and more SSWs in the future NH (Fig. 3, bottom; SSW-reversal only).
 222 However, we have less confidence in this result because strong dynamical coupling be-
 223 tween the troposphere and the stratosphere in the NH complicates future projections,
 224 and also because several past studies were unable to reach a consensus on possible fu-
 225 ture changes of SSW occurrence rates over the NH (Manzini et al., 2014; Ayarzagüena
 226 et al., 2018; Wu et al., 2019; Ayarzagüena et al., 2020). There is also no consensus about
 227 the future strength of the polar vortex (Simpson et al., 2018), which is in agreement with
 228 our conclusion that the polar vortex strength is important for the frequency of SSWs.

229 6 CMIP6

230 To check the robustness of our single model simulations, we repeat our analysis with
 231 CMIP6 data (see supplementary text S4 for details). We find that these models show
 232 a positive polar vortex strength bias (Fig. 4) and generally struggle to produce the ob-
 233 served frequency of SSWs, with a range of 0.3-2.4 SSW-weak events on average in 80 years
 234 for piControl (Table S1). The low SSW frequency in CMIP6 was also briefly noted in
 235 recent work (Ayarzagüena et al., 2020). However, the statistical analysis again suggests
 236 a decrease in SSWs in the future, with three models producing one single and two mod-
 237 els producing no SSW-weak event in SSP585 between 2021 and 2100 (Table S1b). Sim-
 238 ilar to our CM2.1 simulations, the CMIP6 models consistently project a strengthening
 239 of the SH polar vortex (Fig. 4), suggesting that our main conclusion that SSWs will be-
 240 come much rarer in the future is robust.

241 Our enhanced CO₂ CM2.1 simulation only considers future increases in CO₂. Changes
 242 in other radiatively active gases, in particular the expected recovery of the ozone hole
 243 by 2080 (Dhomse et al., 2018), are not included. However, our 1120 ppm CO₂ concen-

244 tration is equal to the CO₂ concentration at the end of the century following the SSP585
 245 scenario (which in addition to CO₂ also increases other greenhouse gases such as methane
 246 and nitrous oxide (O'Neill et al., 2016; Meinshausen & Nicholls, 2020)). Consequently,
 247 u₁₀₆₀ of our future simulation compares well to the end of the 21st century in CMIP6
 248 SSP585 model data (Fig. 4b). This is consistent with previous findings that over the long
 249 term, the greenhouse effect from increasing CO₂ concentrations dominates the effect of
 250 the ozone hole recovery (Barnes & Polvani, 2013). The similarities in u₁₀₆₀ and CO₂ con-
 251 centrations between our CM2.1 simulations and CMIP6 models gives us confidence that
 252 our enhanced CO₂ simulation is relevant for end-of-century projections.

253 7 Conclusions

254 The 2002 and 2019 SSWs both resulted in exceptionally small ozone holes as have
 255 not been observed since the 1980s. They were also followed by extended periods of neg-
 256 ative Southern Annular Mode at the surface, and 2019 in particular was linked to the
 257 catastrophic fire season in South Eastern Australia. While possibly predictable on the
 258 seasonal time scale, it has been difficult to determine how often SSWs should be expected
 259 in the southern hemisphere, due to a relatively short observational record on one hand
 260 and large model biases in the southern hemisphere stratosphere in most comprehensive
 261 climate models on the other hand. Using a pair of exceptionally long and low bias cli-
 262 mate model runs, we found that while SSWs in the SH have significant impacts on strato-
 263 spheric ozone and surface weather, such events are rare and will become even rarer as
 264 CO₂ concentrations increase. In our simulation based on 1990 conditions, the mean re-
 265 turn time for events similar to the 2002 and 2019 SSWs is about 22 years, with a 57%
 266 chance of at least two and a 30% chance of three or more SSW-weak events happening
 267 within the time period spanned by the satellite record. Thus, it is no surprise that two
 268 events have been observed, and there would be a fair chance of another SSW (of either
 269 flavor) in the near future if CO₂ levels were kept constant. However, we show that one
 270 should not make predictions of future occurrence from past data; given that the world
 271 follows a high emissions pathway, our projections suggest that events similar to 2002 and
 272 2019 will become extremely rare, with a mean return time of one in 309 years (or 0.3%
 273 each year) by the end of the century.

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 281 & Hamman, 2017) and aostools (Jucker, 2021) packages. The zonal mean zonal wind time-
 282 series from our simulations are available at [https://data.mendeley.com/datasets/
 283 hknv82tz7v/draft?a=84a23625-6306-440f-ae69-19011c620c7a](https://data.mendeley.com/datasets/hknv82tz7v/draft?a=84a23625-6306-440f-ae69-19011c620c7a). All authors contributed
 284 to conceptualization, methodology and writing of the original draft. T.R. provided the
 285 GCM simulations and M.J. performed the formal analysis. Authors declare no compet-
 286 ing interests.

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