

# A mechanistic model and experiments on bedrock incision and channelization by rockfall

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1      **A mechanistic model and experiments on bedrock incision and channelization by**  
2      **rockfall**  
3

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13      **Key Points:**

- 14      • Rockfall can erode rocky hillslopes even below the angle of repose  
15      • Grain size has a dominant effect on impact abrasion; slope is of minor importance  
16      • Topographic steering of grains results in self-formed bedrock channels

19      **Abstract**

20      Rockfall and rock avalanches are common in steep terrain on Earth and potentially on other  
21      planetary bodies such as the Moon and Mars. Since impacting rocks can damage exposed  
22      bedrock as they roll and bounce downhill, rockfall might be an important erosive agent in steep  
23      landscapes, even in the absence of water. We developed a new theory for rockfall-driven  
24      bedrock abrasion using the ballistic trajectories of rocks transported under gravity. We  
25      calibrated this theory using laboratory experiments of rockfall over an inclined bedrock  
26      simulant. Both the experiments and the model demonstrate that bedrock hillslopes can be  
27      abraded by dry rockfall, even at gradients below the angle of repose, depending on the bedrock  
28      roughness. Feedbacks between abrasion and topographic steering of rockfall can produce  
29      channel-like forms, such as bedrock chutes, in the absence of water. Particle size has a  
30      dominant influence on abrasion rates and runout distances, while hillslope angle is of  
31      comparatively minor influence. Rockfall transport is sensitive to bedrock roughness; terrain  
32      with high friction angles can trap rocks creating patches of rock cover that affect subsequent  
33      rockfall pathways. Our results suggest that dry rockfall can play an important role in eroding  
34      and channelizing steep, rocky terrain on Earth and other planets, such as crater degradation on  
35      the Moon and Mars.

37 **Plain language summary**

38 Rockfall is common on Earth and other planets. Falling rocks bounce down rocky slopes and  
39 likely also erode them. However, it has not been explored how erosive this process is, nor what  
40 landforms it might generate. We developed a numerical model for this erosion process and  
41 calibrated it with experiments of dry grains hopping down an inclined erodible surface. Both  
42 experiments and modeling showed that bedrock erosion from rockfall can happen, even on  
43 relatively low-gradient hills. Small hollows were carved by rockfall, which over time coalesced  
44 into larger troughs that captured the path of subsequent rocks. This process led to a self-  
45 enhancing feedback that produced a bumpy surface with rocky chutes. Rock size had a larger  
46 effect on erosion amounts than the steepness of the hill. Our work suggests that dry rockfall  
47 can play an important role in the evolution of mountain slopes on Earth and craters on the  
48 Moon and other planets.

49

50 **1 Introduction**

51 Rockfall is a ubiquitous, gravitational-driven process in steep terrain. There is evidence  
52 for dry rockfall and rock avalanches on Earth (e.g., Stock et al., 2013), as well as on the Moon  
53 and Mars (Bickel et al., 2020a, 2020b; Kumar et al., 2013; Vijayan et al., 2022; Ward et al., 2011;  
54 Figure 1A to D). There has been a wealth of research into its preconditioning and cause, both on  
55 vertical walls and on mountain slope topography (e.g., Benjamin et al., 2020; D'Amato et al.,  
56 2016; Frayssines and Hantz, 2006; Grenon and Hadjigeorgiou, 2008; Matasci et al., 2018;  
57 Messenzechl et al., 2017; Wieczorek et al., 1992; Williams et al., 2019). Generally, lithological  
58 and exhumation-induced rock fracture, climate, hydrology, and earthquakes are triggers for  
59 rock mass release (André, 1997; Collins and Stock, 2016; Guerin et al., 2013; Hales and Roering,  
60 2007; Leith et al., 2014; Mackey and Quigley, 2014; Moore et al., 2009). Also, sediment mass  
61 routing following rockfall on steep topography has been accounted for in terms of block-runout  
62 and rock avalanching (Dade and Huppert, 1998; Volkwein et al., 2011), in combination with  
63 debris flows and fluvial bedload transport (Mergili et al., 2020; Montgomery and Dietrich, 1994;  
64 Shugar et al., 2021). Dry rockfall and rock avalanches are typically studied due to their  
65 substantial hazard potential. However, they also can be significant agents of erosion, mass  
66 transport, and landscape change (Loye et al., 2012; Delannay et al., 2017; Sass and Krautblatter,  
67 2007). Yet, we currently lack mechanistic modeling and experimental constraints on bedrock  
68 erosion by rockfall.

69 Discrete, dry rockfall in physical experiments was shown to erode sloping bedrock  
70 surfaces even below the angle of repose (Mokudai et al., 2011; Sun et al., 2021). Rockfall  
71 erosion is also supported by observations of boulder tracks (Bickel et al., 2020a, 2020b; Kumar  
72 et al., 2013). The impact energy of large rocks that break free from cliffs is substantial  
73 (Blackwelder, 1942; Le Roy et al., 2019; Rapp, 1960), and their momentum leads to increased  
74 runout distances compared to the smaller grain sizes (Kokelaar et al., 2017; Volkwein et al.,  
75 2011). Bedrock abrasion theory suggests that bedrock erosion should scale linearly with impact  
76 energy and inversely with the square of rock tensile strength (Beer and Lamb, 2021; Sklar and  
77 Dietrich, 2004). So, given abundant rockfall sources in rocky topography, abrasion of bedrock

78 along the rockfall traverse could be an important process in the topographic evolution of steep  
79 terrain (Beer et al., 2019; Sun et al., 2021), despite not being included in most landscape  
80 evolution models.

81 Landforms developed by water-rich rivers and debris flows have received far better  
82 study owing to their importance on Earth and they are known to produce channels. In contrast,  
83 most previous work on dry granular flows has focused on flow over loose granular substrates,  
84 like rock avalanches over talus slopes or grain flows on the front of a sand dune (e.g., Delannay  
85 et al., 2017; Selby, 1982). Granular flows tend to spread laterally (Lajeunesse et al., 2004),  
86 creating relatively smooth convex lobes, such as grain flows on the avalanche face of a wind  
87 dune. The subtle levees and depressions in-between lobes tend to be filled in or diffused away  
88 by subsequent avalanches (McDonald and Anderson, 1996). Dry granular flows also tend to  
89 cease movement at relatively steep angles of repose, which is around 35–45° for most grains,  
90 forming a cone or planar talus slope (Figures 1E and G; Delannay et al., 2017; Kirkby and  
91 Statham, 1975; Sass and Krautblatter, 2007; Selby, 1982). Similar angles for dry granular flow  
92 deposits have been measured on Mars (Atwood-Stone and McEwen, 2013; Dickson et al.,  
93 2007). The generally smooth and steep topography from dry flows over loose substrates  
94 contrasts sharply with the channel-like landforms developed in some steep rocky terrain, such  
95 as bedrock chutes (Figures 1C and D; Ward et al., 2011). This contrast has fueled the idea that  
96 water is needed to develop channelized forms, particularly at slopes less than the angle of  
97 repose for dry avalanches (e.g., Howard, 2007).

98 The mechanics of how flowing water produces channels is relatively well understood.  
99 Water follows the steepest slope, such that the topography funnels the flow, causing erosion  
100 rates to increase, which, in turn, causes further channelization (Horton, 1945). It is unclear if a  
101 similar feedback can occur for dry rockfall. While dry granular flows can be focused down pre-  
102 existing topography (Pelletier et al., 2008), dispersive pressures due to grain-grain and grain-  
103 bed collisions cause granular flows to spread laterally (e.g., Lajeunesse et al., 2004; Figure 1G),  
104 rather than to focus and entrench. However, Sun et al. (2021) showed in an experiment that dry  
105 rockfall traversing a bedrock substrate can form channelized landforms. The strong substrate  
106 allowed for persistent topographic forms over many rockfall events, which steered rockfall into  
107 preferred pathways. Thus, similar to fluvial incision, rockfall was funneled into proto-channels,  
108 enhancing erosion there (Figure 1H) and allowing for further entrenchment. Another  
109 experiment that produced chutes by dry flows used a very light and fine-grained sediment  
110 substrate under high humidity, which provided cohesive strength between grains (Shinbrot et  
111 al., 2004). Rockfall over a relatively smooth bedrock substrate can traverse relatively low-  
112 sloping terrain due to low friction angles (DiBiase et al., 2017; Sun et al., 2021), which may help  
113 explain channel-like landforms below the angle of repose on the Moon and Mars in the absence  
114 of water (Conway et al., 2015; Dickson et al., 2007; Heldmann and Mellon, 2004).

115 Here we develop theory and a numerical model for abrasion by rolling and bouncing  
116 rocks over a bedrock bed in order to better understand the role of rockfall in landscape  
117 denudation and landform development. We calibrated and evaluated the model against  
118 physical experiments of dry rockfall traversing a planar and tilted bedrock slope, where we used  
119 polyurethane foam as a bedrock simulant, similar to Sun et al. (2021). We used the model to

120 answer whether rockfall can form channelized landforms, and the effect that hillslope angles  
 121 and rockfall sizes have on rockfall erosion rates.

## 122 **2 Dry Grain Abrasion Model (DGAM)**

### 123 **2.1 Grain trajectories**

124 We develop a dry grain abrasion model (DGAM), which tracks discrete rockfall events,  
 125 including grain trajectories and abrasion over a gridded 2.5D digital elevation model (DEM with  
 126 no 3D overhangs), built out of  $X, Y, Z$ -coordinates with slopes  $\theta_{cell}$  [ $^\circ$ ] and cellsize  $d_{cell}$  [m]  
 127 (setup in Figure 2A; scheme and workflow in the Supplemental Information and Table S1; see  
 128 notation section). For simplicity, we model only one rockfall grain size and set  $d_{cell}$  equal to the  
 129 grain diameter  $d_{grain}$  [m]. Each grain of mass  $m_{grain}$  [kg] is released from the upstream  
 130 boundary of the model domain with initial variable values (grain deflection velocity,  $v_{out,0}$   
 131 [m/s], absolute grain deflection angle,  $\alpha_{out,0}$  [ $^\circ$ ], and grain hop length,  $l_{hop,0}$  [m]), in one of the  
 132 D8 grid directions (parameter  $\xi_0$ ; i.e. deflection to all adjacent neighbor cells). By having these  
 133 variables drawn from an intended distribution, this procedure ensures controllable randomness  
 134 to the first impacts.

135 Inside the model domain, grains hop over multiple cells following classical mechanics  
 136 (ballistics) along a tilted plane. For a defined grain impact cell  $i$  (with coordinates  $X, Y, Z_i$ ), we  
 137 calculate the incoming grain's trajectory from its original deflecting cell  $(X, Y, Z_{i-1})$  as grain hop  
 138 length  $l_{hop,i-1}$  along the direction  $\xi_{i-1}$ , as the distance between both cell's coordinates ( $\Delta XY$  is  
 139 the horizontal distance, and  $\Delta Z$  is the vertical distance between both cells)

$$140 \quad l_{hop,i-1} = \sqrt{\Delta XY^2 + \Delta Z^2} \quad (1)$$

141 The hop time,  $t_{hop,i-1}$  [s], of this trajectory is based on the grain's original deflection variables  
 142  $v_{out,i-1}$  and  $\sin \alpha_{out,i-1}$ , as well as on gravitational acceleration,  $a_{grav}$  [m/s $^2$ ]:

$$143 \quad t_{hop,i-1} = v_{out,i-1} \sin \alpha_{out,i-1} + \frac{\sqrt{(v_{out,i-1} \sin \alpha_{out,i-1})^2 + 2a_{grav}\Delta Z}}{a_{grav}} \quad (2)$$

144 Grain velocity at the cell impact,  $v_{in,i}$  [m/s], and grain impact angle,  $\alpha_{in,i}$  [ $^\circ$ ], then are:

$$145 \quad v_{in,i} = \sqrt{(v_{out,i-1} \cos \alpha_{out,i-1})^2 + (v_{out,i-1} \sin \alpha_{out,i-1} - a_{grav}t_{hop,i-1})^2} \quad (3)$$

$$146 \quad \alpha_{in,i} = \arcsin \frac{v_{out,i-1} \cos \alpha_{out,i-1}}{v_{in,i}} \quad (4)$$

147 After an impact (i.e., along the next trajectory direction,  $\xi_{i+1}$ ), the grain trajectory  
 148 follows a probabilistic direction-sampling based on weighted downslope gradients in the  
 149 proximity of the impact cell (DiBiase et al., 2017; Dorren et al., 2004). This procedure is  
 150 intended to account for natural stochasticity of the rebounds due to grain inertia, grain shape,  
 151 and surface roughness (cf. Volkwein et al., 2011).

152 The model can be operated in two modes to assess frictional losses due to impacts with  
 153 the bed. In the pure grain hop mode (mode I), grain kinetic energy loss from impacts is  
 154 expressed in grain velocity reduction by means of a shock term,  $\kappa_{shock}$  [1/m] (Quartier et al.,  
 155 2000),

$$v_{out,i} = v_{in,i} - \kappa_{shock} v_{in,i}^2 \Delta t \quad (5a)$$

with  $v_{out,i}$  [m/s] is the deflection grain velocity, and  $\Delta t$  [s] is an impact time, which we assume to be 0.1 s (DiBiase et al., 2017; Gabet and Mendoza, 2012). In mode II, impact energy loss also includes sliding and rolling friction based on a modified Coulomb friction law (DiBiase et al., 2017; Gabet and Mendoza, 2012),

$$v_{out,i} = v_{in,i} - a_{grav}(\sin \theta_{cell,i} - \tan \Phi_{surf,i} \cos \theta_{cell,i}) - \kappa_{shock} v_{in,i}^2 \Delta t \quad (5b)$$

which includes the dynamic surface friction angle,  $\Phi_{surf,i}$  [ $^\circ$ ] between grains and the surface, accounting for microtopography. Following previous work (DiBiase et al., 2017; Gabet and Mendoza, 2012), we treat this slope as an exponential probability distribution

$$\Phi_{surf,i} = \arctan\left(\frac{1}{\tan \bar{\mu}_{surf}} e^{-\frac{\tan \mu_{eff}}{\tan \bar{\mu}_{eff}}}\right) \quad (6)$$

of the effective friction angle  $\mu_{eff}$  [°]. The grain's deflection angle  $\alpha_{out,i}$  [°] is assumed to be the reflection angle of  $\alpha_{in,i}$  on the local cell slope  $\theta_{cell,i}$  [°] in direction  $\xi_i$

$$\alpha_{out,i} = \alpha_{in,i} - 2\theta_{cell,i} \quad (7)$$

The location of the next impact ( $X, Y, Z_{i+1}$ ) then is determined by an iterative process. The grain's trajectory heights  $Z_{traj}$  [m] are calculated relative to the traversed cell boundaries  $Z_{cell.boundary}$  [m] along the grain's trajectory direction  $\xi_i$

$$Z_{traj} = \tan\alpha_{out}(n_{cells} + 0.5) \frac{d_{cell}}{2} - \frac{a_{grav}}{2v_{out,i}^2 \cos^2\alpha_{out,i}} \left( \frac{d_{cell}}{2} \right)^2 \quad (8)$$

for a number of  $n_{cells}$  until  $Z_{traj} < Z_{cell.boundary}$ . Then  $(X, Y, Z_{i+1})$  is defined as the last cell that could not be traversed by the grain, and  $l_{hop,i}$  is calculated in between  $(X, Y, Z_i)$  and  $(X, Y, Z_{i+1})$  (Equation 1). The grain's hop height  $h_{hop,i}$  [m] is the maximum of the vertical distances between traversed cell boundaries and trajectory heights, i.e.  $\max(Z_{traj} - Z_{cell.boundary})$ . The trajectory procedure is repeated until the grain leaves the model domain or comes to rest (Table S1). If a grain has too low deflection velocity,  $v_{out,}$  or too low deflection angle,  $\alpha_{out,}$  to cross the next cell boundary, it is deposited at the current cell. We assume a resting particle is subsequently set in motion from being hit by a mobile grain, drawing randomly from the grain entrance variable values discussed above. However, if a grain is in a depression with neighboring cells higher than two grain sizes, we assume the grain stays there and acts as cover that protects the bedrock from abrasion.

## 2.2 Bedrock abrasion and morphodynamics

The amount of bedrock abrasion of a cell,  $w_{cell,i}$  [m], due to a single grain impact is calculated as:

$$w_{cell,i} = k_{ero} \frac{0.5 m_{grain} v_{in,n,i}^2}{d_{cell}^2} = k_{ero} \frac{\varepsilon_{kin,n,i}}{d_{cell}^2} \quad (9)$$

$$v_{in,n,i} = \cos \theta_{cell,i} v_{in,i} \sin \alpha_{in,i} \quad (10)$$

191 based on a bedrock erosion efficiency factor,  $k_{ero}$  (i.e., grain erosivity, [ $m^3/J$ ]). To conserve  
192 mass, the erosion amount,  $w_{cell,i}$ , is assessed in the vertical direction since the cell area in our  
193 calculations is measured on a horizontal grid. The surface-normal component of the kinetic  
194 impact energy,  $\varepsilon_{kin,n}$  [J], results from the surface-normal component of the grain impact  
195 velocity,  $v_{in,n}$  [m/s]. The velocity component accounts for impact-induced fracturing causing  
196 wear, instead of surface-parallel gouging (cf. Sun et al., 2021). For steeper slopes, its value  
197 decreases compared to the approximation of a vertical impact velocity that is commonly used  
198 in fluvial abrasion theory (Beer and Lamb, 2021; Beer and Turowski, 2021; Engel, 1978; Sklar  
199 and Dietrich, 2004). While the actual geometry of the impact event depends on local  
200 parameters like grain shape and bedrock roughness that are not explicitly included in the  
201 model, the model is calibrated with experiments (below) and thus these local geometric effects  
202 are incorporated into the empirical model parameters.

203 The DGAM model allows the user to switch off abrasion (Equation 9), since it is  
204 decoupled from frictional losses (Equation 5). This option enables process-independent model  
205 assessment like varying grain sizes or hillslope angle. Accounting for abrasion (Equation 9)  
206 results in evolving hillslope topography that influences grain impact energy  $\varepsilon_{kin}$  (via modified  
207 hop time  $t_{hop}$ , which drives impact velocity  $v_{in}$ ; Equation 3), and alters the local slope gradient  
208 around each cell. This again affects the subsequent direction of deflecting grains (parameter  $\xi$ ),  
209 which can result in topographic steering feedback. Application of the model requires inputs of  
210 initial grain entrance variables ( $v_{out,0}$ ,  $\alpha_{out,0}$  and  $l_{hop,0}$ ) and the bedrock erosion factor,  $k_{ero}$ ,  
211 per model cell. The grain impact shock term,  $\kappa_{shock}$ , is calibrated as described below, but could  
212 be adjusted for specific situations (e.g., varying grain shape).

213

### 214 **3. Experimental Setup and Model Application**

215 We conducted two sets of experiments (Table 1) to generate grain trajectory and  
216 abrasion data, and used this data to calibrate the DGAM model. The first set consisted of five  
217 large-scale experiments with an erodible foam substrate that evolved during the experiments  
218 due to abrasion from dry rockfall. We refer to these as *erodible-bed experiments* (EB). The  
219 erodible-bed experiments had different inlet conditions, hillslope gradients, and particle sizes  
220 to test the model performance relative to these variables. The experiments of the second set  
221 were of smaller scale, did not vary the inlet nor erode the bed, and were used to evaluate grain  
222 trajectories as a function of bed slope. We refer to these as *fixed-bed experiments* (FB).

223

#### 224 **3.1. Erodible-bed experiments**

225 We ran five erodible-bed experiments (EB, Table 1; more details in Table S2). These  
226 experiments were not designed to replicate or reproduce particular rockfall and hillslope  
227 topography, but to provide data on grain trajectories, bedrock abrasion, and morphodynamic  
228 feedback for model comparison. Erodible-bed experiment 1 (i.e., EB1) was conducted using a  
229 2.2 m long, 0.76 m wide test section using large river cobbles on a relatively shallow sloping  
230 bed ( $\theta_{slope} = 16.7^\circ$ ). The detailed experimental setup and some results from experiment EB1

were previously described in Sun et al. (2021). These observations include the ability of rockfall to run out over low gradients and to focus, resulting in channelized landforms through topographic steering. Here, we use data from *EB1* to help evaluate DGAM and to compare results from four additional erodible-bed experiments and six fixed-bed experiments, as detailed below.

The four new erodible bed experiments (*EB2 – EB5*) were conducted in a different but comparable facility as *EB1*. We used a tilting flume, 4.5 m long and 0.65 m wide, filled with a block of smooth, homogeneous polyurethane (PU) foam, which acted as a highly-erodible substitute for bedrock (Scheingross et al., 2014; Figures 2B, C). Each experiment (including *EB1*) used the same type of foam with a density of  $0.06\text{t/m}^3$ , a tensile strength of  $\sigma_{foam} = 0.32\text{MPa}$ , and a Young's Modulus of  $3.92\text{MPa}$ . This foam has been shown previously to produce realistic erosional morphologies through abrasion by grain impacts in both air and water (Scheingross et al., 2014; Sun et al., 2021). Moreover, the foam erodibility follows the same scaling law with tensile strength as bedrock, supporting it as an experimental analog to natural rock (Beer and Lamb, 2021; Lamb et al., 2015). The erodibility framework holds over several orders of magnitude both in impactor energy and impact abrasion (Beer and Lamb, 2021), indicating that these laboratory experiments can be scaled to natural cases of larger impact energies and real bedrock using the relative erodibilities in a scaling factor.

The variables that changed between our experiments were the inlet design for the grains to enter the flume, grain size/shape properties, and the flume slope (Table 1). Experiment *EB1* used rounded granitic grains (density of  $2.75\text{ t/m}^3$ ) with a median grain diameter of  $d_{grain} = 0.061\text{m}$ ; experiments *EB2-EB4* used medium-sized and rounded andesitic grains ( $d_{grain} = 0.023\text{m}$  and  $0.03\text{m}$ , respectively; grain density of  $2.33\text{t/m}^3$ ; Figure 2D); and *EB5* used  $0.015\text{m}$  angular granite grains. The initial slope of the planar foam bed was  $\theta_{slope} = 16.7^\circ$  for *EB1*,  $19.5^\circ$  for *EB2-EB4* and  $35.0^\circ$  for *EB5*. The inlet for rockfall spanned the width of the flume for experiments *EB1*, *EB2* and *EB5*, but was constricted to  $0.2\text{m}$  width in the flume-center for experiments *EB3* and *EB4* (Table S2).

The experiments were designed such that the surface friction angle of the grains on the foam  $\Phi_{surf} [^\circ]$  was similar to the slope  $\theta_{slope}$  of the planar foam bed at the beginning of the experiment (Table 1). This design was intended to allow grains to be intermittently mobile even when patches of static grains were deposited on the bed. The grain's pocket friction angle,  $\Phi_{pocket} [^\circ]$  (corresponding to the angle of repose of a grain pile) was measured following previous work (Prancevic and Lamb, 2015; Sun et al., 2021), whereby we glued grains of like size and angularity on a planar board. Then a loose grain was placed on this surface, the board was slowly tilted until the grain was mobilized, and the tilting angle was reported as the pocket friction angle. The process was repeated for  $\sim 100$  different grains selected at random and placed at random on the board. We also repeated this process for grains placed on the planar foam board, which we report as the mean surface friction angle,  $\Phi_{surf} [^\circ]$ . Grains should be highly mobile when their friction angle is lower than the topographic slope (DiBiase et al., 2017), which was the case for all of our experiments with grains traversing the smooth foam bedrock. However, this mobility transiently changed during the experiments due to the growth of topographic bedrock roughness and due to static patches of grains that were more difficult

273 to traverse (i.e.,  $\Phi_{pocket} > \Phi_{surf}$ ; Table 1). Although we achieved high mobility in the  
274 experiments through relatively round grains and smooth foam topography, low surface friction  
275 angles are also expected in natural settings with angular rockfall grains that are much larger  
276 than the bedrock topographic roughness (DiBiase et al., 2017; Sun et al., 2021). In other words,  
277 modeling multi-meter scale boulders in the laboratory is not feasible, so we created similar  
278 particle dynamics by lowering the surface friction angle through particle roundness rather than  
279 by larger grain size.

280 Each experiment started with a new block of planar, smooth foam (Figure 2C). Dry  
281 grains were introduced at the upslope end of the flume at a steady rate from an auger  
282 sediment feeder. The feed rate was slow enough (250 – 1'550 grains/minute), so grains  
283 entered and traversed the flume individually, with minimal grain-grain interactions. Particles  
284 traversed a board with pegs spaced at 0.05m to spread the grains across the inlet. The flume  
285 had rigid vertical walls that reflected grains towards the center of the test section, mimicking  
286 grains exiting and entering the domain under an infinitely wide scenario. Each experiment  
287 lasted for several hours of runtime, in which 5 – 22 t of sediment traversed the test section  
288 (Table 1; details in Table S2).

289 Grain trajectories were recorded using high-speed cameras with fisheye lenses  
290 (GRASHPER, set to 160 frames/s) at three lateral positions along the flume (Figure 2B) and  
291 one camera from top. We rectified and cut the distorted fisheye-lens pictures, converted them  
292 to black and white, and scaled their dimensions by scale bars attached to the flume walls in the  
293 photos. Then we applied particle imaging velocimetry (PIV) to measure grain trajectories using  
294 Python-based software packages (OpenCV and TrackPy; Python, 2021) and calculated grain  
295 trajectory metrics. For the side-view cameras, these metrics were grain impact and deflection  
296 velocities ( $v_{in}$  and  $v_{out}$ ), impact and deflection angles ( $\alpha_{in}$  and  $\alpha_{out}$ ), hop heights, and hop  
297 lengths ( $h_{hop}$  and  $l_{hop}$ ; Figure 2B; Table 1; Table S2), which we calculated perpendicular to the  
298 foam surface from grain traces through subsequent pictures. We only used complete grain  
299 trajectories showing several hops, but discarded incomplete trajectories, photos with unclear  
300 grain detection from the black-white conversion, and photos comprising several grains. For the  
301 top camera, we only calculated the lateral and downslope ( $X, Y$ ) coordinates of the trajectories,  
302 as we could not detect the actual impact positions.

303 We surveyed the evolving foam bed topography approximately every one to two hours  
304 in each experiment. During this time, we stopped the particle feed and removed any  
305 accumulated foam dust using compressed air. The foam surface was surveyed from two  
306 positions above the flume using a terrestrial laser scanner, TLS (FARO FOCUS 3D), which  
307 delivered 3D pointclouds (i.e.,  $X, Y, Z$ -coordinates) with a mean spatial resolution of ~1 mm.  
308 The individual, subsequent TLS-measured pointclouds were co-registered on the initial smooth  
309 surface using twelve fixed target points along the flume walls (0.1m-diameter wooden spheres,  
310 which allowed for calculating their centers; Figure 2B and 2C). Vertically differencing the co-  
311 registered pointclouds using the M3C2 algorithm in cloudcompare software (CloudCompare,  
312 2022; Lague et al., 2013), we calculated transient spatial foam abrasion and also noted the total  
313 abrasion volume (i.e., total abrasion amount over the whole flume surface;  $V_{flume}$  [ $m^3$ ]).

315 **3.2 Fixed-Bed Experiments**

316 The fixed-bed experiments (*FB*) were designed to gain more data on grain hop  
 317 trajectories but using a simpler setup than the erodible-bed experiments. The experiments used  
 318 a tilting chute that was 1.1 m long and 0.1 m wide. Six experiments were conducted (*FB1-FB6*),  
 319 each with identical parameters except that the flume bed slope,  $\theta_{slope}$ , was varied between 20°  
 320 and 45° (Table 1; lower part). The experiments used the same rounded andesite gravel as  
 321 experiments *EB3* and *EB4*. The flume bed consisted of the same foam as in the erodible-bed  
 322 experiments, but since it was only traversed by a hundred grains over time, abrasion was  
 323 negligible, and the topography remained planar. Grains were fed into the chute individually by  
 324 hand. A high-speed lateral-view camera (the same as described above) was used to capture  
 325 grain trajectories, and grain trajectory analysis was the same as in the erodible bed  
 326 experiments.

327

328 **3.3. Comparing the Model to Experiments and Natural Cases**

329 As we want to verify the dry rockfall abrasion theory to represent a feasible hillslope  
 330 erosion process, we (i) calibrate the DGAM model to reproduce the experimental observations  
 331 of the EB and FB, then (ii) explore grain trajectories and abrasion varying hillslope angle and  
 332 rockfall grain size, and finally (iii) scale the model to predict natural hillslope topography.

333 To run the model for the experimental setups, grain reflection from the flume walls was  
 334 accounted for by stopping a grain's trajectory on the last cell in front of the wall. From there, it  
 335 starts a new trajectory with its given variables but in a new direction  $\xi$ . Mean foam abrasion  
 336 per grain impact,  $V_{cell}$  [m<sup>3</sup>], for an experiment of a given flume slope and grain type (Table 1)  
 337 was calculated as:

338

$$339 V_{cell} = \frac{V_{flume}}{n_{grains,tot} n_{imp,tot}} = \frac{V_{flume} m_{grain} l_{hop}}{m_{grains,tot} l_{flume}} \quad (11)$$

340

341 by estimation of the number of grains used,  $n_{grains,tot}$  [-] (i.e., the total sediment mass fed into  
 342 the experiment,  $m_{grains,tot}$  [kg], divided by a single grain's mass,  $m_{grain}$ ), and the mean  
 343 number of impacts per grain along the flume,  $n_{imps,tot}$  [-] (i.e., flume length,  $l_{flume}$  [m], divided  
 344 by mean grain hop length,  $l_{hop}$ ). To convert the experimental results into the grid world of the  
 345 DGAM model, we assumed this abrasion volume is equally distributed over a model cell that is  
 346 impacted by a grain (i.e.  $V_{cell} = w_{cell} d_{cell}^2$ , with  $d_{cell} = d_{grain}$ , as defined above). This  
 347 assumption is reasonable, given the observation of generally platelet-shaped bedrock  
 348 fragments abraded from grain impacts (Beer and Lamb, 2021). We then scaled the grain  
 349 erosivity factor,  $k_{ero}$ , as the fraction between  $V_{cell}$  and the surface-normal component of the  
 350 grain's mean kinetic impact energy,  $\varepsilon_{kin,n}$  (Equation 9).

351 Using the calculated  $k_{ero}$  values and the measured initial grain entrance variables  
 352 ( $v_{out,0}$ ,  $\alpha_{out,0}$  and  $l_{hop,0}$ ) for each erodible-bed experiment *EB* (Table 1, upper part), we  
 353 iteratively fit the DGA model (mode I, i.e. pure grain hopping) shock term coefficient,  $\kappa_{shock}$ , to

354 best reproduce the means of the observed grain trajectory variables and foam surface abrasion  
 355 rates of the experiments.

356 Having calibrated the model, we used it to explore the rockfall transport and impact  
 357 abrasion over a range of natural hillslope angles ( $5 < \theta_{slope} < 45$ ) and grain sizes ( $0.1\text{m} < d_{grain}$   
 358  $< 1\text{m}$ ). To model dry rockfall abrasion on rocky hillslopes under natural scenarios of hillslope  
 359 angle, grain sizes and lithology, we scaled the bedrock abrasion rate ( $V_{cell}$ ) according to the  
 360 rock tensile strength following  
 361

$$362 V_{rock} = V_{cell} \left( \frac{\sigma_{foam}}{\sigma_{rock}} \right)^2 \quad (12)$$

363 where  $V_{rock}$  [ $\text{m}^3$ ] is the volumetric abrasion for any bedrock cell of tensile strength  $\sigma_{rock}$  (Beer  
 364 and Lamb, 2021; Scheingross et al., 2014).  
 365

## 367 4. Results

### 368 4.1 Topographic evolution in the experiments

369 All five erodible-bed experiments (*EB*; Table 1, upper part) evolved in a similar pattern,  
 370 and the final bed topographies resembled each other (Figure 3B). Here, we describe the general  
 371 evolution of these experiments to document the dry abrasion process, using *EB5* as an example  
 372 (Figure 3A). Grains discretely hopped down the foam surface and abraded it by incremental  
 373 impact abrasion, resulting in tiny pit craters and abraded foam dust creating lasting topography  
 374 (cf. Figure 1H). Initial grain abrasion pits down the entrance transiently grew into larger hollows  
 375 from ongoing impacts of subsequent grains (Figure 3A, left panel, shown for *EB5*), although  
 376 separate hollows were less distinct for the largest grains used in *EB1*. Grains leaving these  
 377 hollows initiated faint (mm-deep), parallel rills down the slope. Reaching a depth of around one  
 378 grain diameter, these hollows laterally coalesced into a trough, and the rills further evolved  
 379 (Figure 3A, central panel). This process is portrayed by the temporal evolution of the lateral  
 380 profiles through the hillslope (Figure 4A, upper panel). Over time, the rills extended in depth  
 381 and converged downslope into a central main channel (Figure 3A, right panel; Figure 4A, lower  
 382 panel). This channel's long profile maintained a slight bumpiness over time (Figure 4D), arising  
 383 from the subsequent evolution of new troughs, whose rims transiently traversed downslope  
 384 (Figure 3A, right panel). This pattern emerged in all erodible bed experiments (Figure 3B).

385 Throughout the experiments, some grains came to rest, though they were soon hit by  
 386 mobile grains and remobilized. So, permanent spatial cover generally did not occur on the main  
 387 foam board, even at the lowest experimental slope of  $\theta_{slope} = 17^\circ$  (*EB1*; Table 1). However,  
 388 when a topographic depression (as the upper trough) reached a depth of two grain sizes  
 389 relative to its downslope rim, it gradually got clogged by resting grains, which formed a  
 390 stationary cover in the depression. Subsequent grains laterally traversed this patch of grains  
 391 and funneled into the evolving main central channel, uniting the former rills downslope  
 392 (Figure 3A, right panel); a similar sequence was described for *EB1* by Sun et al. (2021). Due to  
 393 the focusing of the grains into the central trough, lateral parts of the foam surface experienced

394 a decreasing number of grain impacts over time (Figure 4B), and they gradually abraded slower  
 395 (Figure 4C; shown here is vertical abrasion equivalent to  $w_{cell}$ , for comparison). This  
 396 morphodynamic feedback resulted in a channelized hillslope for all erodible-bed experiments,  
 397 independent of hillslope angle or grain size (Figure 3B). The current pattern of the abrasion  
 398 measurements therein reflected the current surface topography, e.g., the eroded rills (Figure  
 399 4C upper panel vs. Figure 3A central panel).

400 When the patch of static grains in the upper trough initiated, it grew laterally and in  
 401 height due to the higher pocket friction angle of the grain patch relative to the foam board (cf.  
 402 Table 1). Once this grain pile backed up onto the peg board, the experiment was terminated  
 403 (Figure 2B). Without the upslope limitation of the experimental facility, the grains probably  
 404 would have continued piling until reaching their pocket friction angle, resulting in a grain  
 405 avalanche, followed by a subsequent pile-up, and so on. Final bedrock topographies typically  
 406 consisted of an upslope trough filled with a static grain patch, with a channel that extended and  
 407 became less defined downslope (Figure 3B and Figure 4D; shown after sediment cover patch  
 408 removal).

409 Changing the grain inlet width for the erodible-bed experiments (flume-wide for *EB1*,  
 410 *EB2*, and *EB5*, central for *EB3* and *EB4*; Table S2) dictated the lateral extend of the upper trough  
 411 (Figure 3B). The larger the grain size of the experiment, the farther the trough extended  
 412 downslope (cf. *EB1* vs. *EB5*). Regardless of inlet width or particle size, all experiments showed a  
 413 smooth rim at the trough outlet, followed by rills and a subsequent emerging trough, which  
 414 initiated a channel (Figure 3A and Figure 4D).

#### 415 4.2. Grain trajectories and model calibration

416 On average, for the erodible-bed experiments *EB1-EB3* and *EB5* (for *EB4*, there were too  
 417 few measurements available for robust statistics), grains hopped by  $\bar{l}_{hop} = 0.19 \pm 0.11$  m at  
 418  $\bar{h}_{hop} = 0.02 \pm 0.02$  m height (mean and standard deviation) (Figure 5, grey boxplots; Table S2).  
 419 They impacted at angles of  $14 \pm 14^\circ$  above the respective foam surface (i.e.,  $\bar{\alpha}_{in} - \theta_{slope}$ ).  
 420 The grain's hop lengths, hop heights, and impact angles were insensitive to the hillslope angle.  
 421 However, the mean impact velocities,  $1 < \bar{v}_{in} < 2$  m/s, increased with steeper hillslope  
 422 angles. Deflection angles and deflection velocities generally equaled their impact pendants, so  
 423 little kinetic energy was lost by the impacts. Mean initial grain entrance velocity from the peg  
 424 board was around  $v_{in,0} = 1.1$  m/s in the erodible-bed experiments. In the fixed-bed  
 425 experiments, these velocities were higher ( $\sim 1.5$  m/s), resulting in increased hop lengths, impact  
 426 angles, and impact velocities (Figure 5, white boxplots).

427 Derived vertical grain impact abrasion volumes per cell area,  $w_{cell}$  (i.e.  $V_{cell}/d_{cell}^2$ ; cf.  
 428 Equation 11) in the order of  $\mu\text{m}$  decreased with increasing slope angle (Figure 6A). This pattern  
 429 is consistent with impacting grain's grain erosivity,  $k_{ero}$  (i.e.  $V_{cell}/\varepsilon_{in,z}$ ), with some uncertainty  
 430 for smaller, rounded grains of low erosivity, while even smaller but angular grains maintained  
 431 their erosivity even for low impact energies (*EB5*; Figure 6B; Table 1). Normalizing  $k_{ero}$  values  
 432 by grain cross-sectional area or cell size,  $d_{grain}^2$ , resulted in an erosivity measure that collapsed  
 433 the data of the fixed-bed experiments with round grains (*FB1-FB4*) around 0.001, while for the

435 angular grains, it remained higher ( $k_{ero}/d_{grain}^2 = 0.003\text{m/J}$ ; Figure 6C; Table 1). These values can  
436 be used to calculate the DGA model's  $k_{ero}$  factor for a given grain size.

437 To calibrate the model using the experiments, we set  $k_{ero}$  based on the observed  
438 erosion amounts (Figure 6B). Next, we kept the observed initial grain entrance variables ( $v_{in,0}$ ,  
439  $\alpha_{out,0}$  and  $l_{hop,0}$ ) fixed in the model and varied the shock term coefficient,  $\kappa_{shock}$ , to best  
440 reproduce the suite of the mean trajectory parameters for each erodible-bed and fixed-bed  
441 experiment. Comparing the predicted versus the modeled means of the grain trajectory  
442 parameters  $l_{hop}$ ,  $v_{in,z}$ ,  $\alpha_{in}$ , and  $V_{cell}$  (the latter parameter only for the erodible-bed  
443 experiments), we identified experiment-specific  $\kappa_{shock}$  values with the first closest general  
444 agreement (Figure 7 for EB2; cf. Figure S1 for all experiments). All these identified values fell in  
445 a narrow range around  $\kappa_{shock} = 0.8 \text{ m}^{-1}$ .  
446

#### 447 **4.3. Model and experimental comparison**

448 Predictions from the  $\kappa_{shock}$ -calibrated model generally fit the pattern of the measured  
449 PIV-derived trajectory parameters along the flume, though the range of the predictions was  
450 much lower (mean deviation -15% and range -60% for EB2 in Figure 8). The largest deviations  
451 existed for the predicted grain hop length,  $l_{hop}$  (-65%, Figure 8A), and the deflection velocities  
452 (-15%, Figure 8D), both mainly further downslope of the flume. Both grain impact and  
453 deflection angles were overpredicted at the flume's entrance. The impact angle,  $\alpha_{in}$ , soon  
454 matched the observations, but the deflection angle,  $\alpha_{out}$ , remained increased (6%, Figure 8E-F).  
455 Overall the grain impact velocities were met (-5% deviation, Figure 8B, C) and thus also the  
456 initially relatively increased impact abrasion fit the calculated values along the flume (-5%,  
457 Figure 8G).

458 All grain trajectory parameters for a fixed grain size increased with a steeper hillslope  
459 angle (Figure 9A; for  $d_{grain} = 0.03\text{m}$ ). Over the range of  $\theta_{slope} = 20^\circ$  to  $45^\circ$  hop length and  
460 impact velocity doubled, while impact angles remained more constant relative to the surface  
461 slope (Figure 9A; upper three panels). The resulting abrasion volume remained within one order  
462 of magnitude for volumetric impact abrasion,  $V_{cell}$ , and also for local erosivity (i.e., abrasion per  
463 meter downslope,  $V_{cell} / l_{hop}$ ; Figure 9, two lower panels). DGAM-predictions over  $\theta_{slope} =$   
464  $5^\circ$  to  $45^\circ$  for both round grains ( $d_{grain} = 0.03\text{m}$ , representative for EB3, EB4, and FB1-FB6) and  
465 for angular grains (EB5) followed the general trends, in which angular grains consistently  
466 underpredicted observed abrasion volume ( $V_{cell}$ ; Figure 9A, two lower panels).

467 In contrast to the influence of slope angle, grain trajectory parameters showed more  
468 sensitivity to increasing grain size when holding slope fixed (Figure 9B; for  $\theta_{slope} = 35^\circ$ ). Over  
469 the range of  $d_{grain} = 0.015$  to  $0.036\text{m}$  hop length, impact velocity, and impact angle all  
470 doubled (upper three panels of Figure 9B). Grain impact abrasion and local erosivity increased  
471 nonlinearly with grain size following a strong trend (two lower panels of Figure 9B; impact  
472 abrasion was not measured for FB4, but the predicted value from the measured impact  
473 energies was comparable to EB5). Accordingly, and from a general perspective of natural  
474 hillslopes, rocky surfaces with slope angles ranging from  $\theta_{slope} = 15^\circ$  to  $45^\circ$  and impacted by  
475 rockfall grains of  $d_{grain} = 0.01\text{m}$  to  $1.00\text{m}$  diameter may experience local impact abrasion

476 volumes spanning six orders of magnitude (Figure S2; calculated using the erosivity for angular  
 477 grains, as in *EB5*; Figure 6B, C). Herein, the influence of slope angle is inferior as compared to  
 478 grain size. The abrasion volumes predicted for laboratory foam can be scaled to abrasion  
 479 volumes of any (massive) bedrock by the inverse square of the material's tensile strengths  
 480 (Equation 13).

481

#### 482 **4.4 Model exploration**

483 Having calibrated the model, we sought to explore the impact of the upstream  
 484 boundary condition on bedrock landforms developed by rockfall. For this, we simulated  
 485 topography evolution from an initially smooth, sloping plain, similar to the experiments, with  
 486 rockfall fed in from the top of the domain. We set the DGAM parameters to be more realistic  
 487 for natural cases, including larger, angular rockfall grains ( $m_{grains,tot} = 800$ tons of  $d_{grain} =$   
 488 0.20m) on a steep granite hillslope ( $\theta_{slope} = 35^\circ$ ;  $\sigma_{rock} = 5$ MPa, cf. Equation 12;  $k_{ero} =$   
 489 0.003m/J, cf. Figure 6C, C;  $\kappa_{shock} = 0.8\text{m}^{-1}$ , cf. Figure 7). All other parameters were set as in *EB2*  
 490 ( $\phi_{surf}$ , grain density,  $v_{out,0}$ ,  $\alpha_{out,0}$  and  $l_{hop,0}$ ). We conducted two numerical experiments with  
 491 all parameters equal except for a change in the feed of rockfall: Uniform feed over the center of  
 492 the model domain (cross-sections in Figure 10A, long profile in Figure 10C) vs. rockfall dispersed  
 493 over three source areas (Figure 10B and D).

494 For the case of a uniform central rockfall entrance, the initially planar hillslope surface  
 495 developed a deepening trough at the entrance, which sourced into a channel with decreasing  
 496 depth further downslope (panels of Figure 10A; more panels in Figure S3A). This process was  
 497 driven by steering of grains into the channel center, increasing abrasion there (transient lateral  
 498 grain distribution in the third panel of Figure 10A; cf. grain trajectories and local impact  
 499 abrasion in Figure S3C and E). Down the hillslope, the hopping grains produced a sequence of  
 500 intermittent and downslope-wandering concave troughs and convex rims of decreasing size,  
 501 comparable to the topographic slope evolution during the experiments (Figure 10C vs. Figure  
 502 4D). The experiment ended when the upper trough reached a depth of one grain diameter  
 503 relative to its downslope rim, capturing all subsequent grains.

504 Modeling with the same number of grains as before, but fed onto the hillslope in three  
 505 separated inlets (Figure 10B; more cross-sections shown in Figure S3B), resulted in comparable,  
 506 but smaller concave-shaped channels downslope, i.e., in parallel rills that started coalescing.  
 507 This experiment also stopped due to over-deepening of the upper trough (Figure 10D), after a  
 508 remaining wider lateral grain distribution and abrasion than in the other experiment  
 509 (Figure 10B third panel; Figure S3D and F).

510

## 511 **5 Discussion**

### 512 **5.1 Model calibration and validation**

513 It is currently not possible to compare our model predictions to natural erosion rates  
 514 because the model requires specification of rockfall frequency, rock size and bedrock strength,  
 515 which are generally unknown. Ultimately, a complete model of landscape evolution by dry  
 516 rockfall will need to incorporate these rockfall generation processes, which can then be coupled

517 to model rockfall abrasion. Due to the lack of field constraints, we turned to scaled laboratory  
 518 experiments to test the model. By varying hillslope angle, grain size, and grain shape, we  
 519 calibrated a cellular, dry grain trajectory abrasion model by means of grain shape erosivity and  
 520 an impact shock term,  $\kappa_{\text{shock}}$  (DGAM; Figures 2, 6 and 7). The grain trajectory velocities, angles,  
 521 and hop length only varied within their magnitude in our *EB* flume experiments and they  
 522 showed a larger spread for the *FB* due to a small test population of some tens of grains (Table 1  
 523 and Table S2; Figure 5 and Figure 8). The calibrated model did not entirely reproduce these  
 524 measured grain trajectories (fewest the hop length; Figure 8), which may be attributable to the  
 525 larger range of the experimental trajectory variables due to uneven grain shape (Figure 5).  
 526 Varying the impact shock term,  $\kappa_{\text{shock}}$ , could account for this discrepancy by generating a wider  
 527 distribution of trajectories. Grain shape likely has a nonlinear influence both on grain mobility  
 528 (angular grains have large pocket and surface friction angles; Table 1; Figure 9) and on grain  
 529 impact erosivity (angular grains will be more erosive; Neilson and Gilchrist, 1968). Though,  
 530 summed impacts of a given grain shape mixture may cancel out varying abrasion volumes of  
 531 different grain shapes, as indicated by the general collapse of experimental abrasion data for  
 532 local impact abrasion (Figure S4D-F). This leveraging is also reflected in the deviation of model-  
 533 predicted lower deflection velocities but higher deflection angles that still led to acceptable  
 534 abrasion rates based on a fixed shock term (Figure 8D, F, and G).

535

## 536       **5.2 Effect of slope and grain size**

537 Constraining grain impact abrasion volume is a crucial factor in the process, and grain  
 538 size showed to be of dominant influence compared with hillslope angle (Figure S2). Modelled  
 539 trajectory parameters increased modestly with increasing slope angle, and abrasion volume  
 540 only rose by one order of magnitude from shallow to steep slopes (Figure 9A). All parameters  
 541 also increased with larger grain size (Figure 9B). Importantly, grain impact erosivity nonlinearly  
 542 rose, spanning six orders of magnitude from pebbles to 1m boulders (Figure 9B lower panels  
 543 and Figure S2) due the nonlinear impact energy-dependence on grain diameter cubed (cf.  
 544  $m_{\text{grain}} = \sigma_{\text{rock}} \frac{4}{3} \pi (\frac{d_{\text{grain}}}{2})^3$ ). This matches the high erosivity of large (meter-sized) rockfall  
 545 boulders analyzed in rockfall runout studies (Bickel et al., 2020a, 2020b; Volkwein et al., 2011)  
 546 and in previous abrasion experiments (Mokudai et al., 2011), and matches their importance in  
 547 fluvial abrasion (Beer and Lamb, 2021; Turowski et al., 2015).

548 Within a distribution of rockfall grain sizes, the largest grains will have an immediate  
 549 effect on surface morphology since both subsequent grain trajectories will be more influenced  
 550 by their erosive impact on surface roughness, and their momentum-dependent runout distance  
 551 is the largest (Kokelaar et al., 2017). Though, the actual/transient grain size distribution will  
 552 determine the representative grain size that may be applicable for average modeling. Field data  
 553 on individual (caprock) rockfall grain size distributions are lacking to our knowledge, though it  
 554 could, e.g., be derived from rocky hillslope's fracture-spacing (Neely and DiBiase, 2020) and  
 555 then allow assessment of the interplay between rockfall erosivity and slope erodibility.

556

## **5.3 Effect of substrate strength**

As grain impact erosivity depends on the surface-normal component of kinetic impact energy, independent of the actual medium through which the grain moves (e.g., air or water), it scales inversely with bedrock substrate tensile strength,  $\sigma_{rock}$  (Beer and Lamb, 2021; Scheingross et al., 2014). Thus, dry grain impact abrasion,  $V_{cell} = k_{ero} \varepsilon_{kin,n}$ , can be transformed to fit into the bedrock erodibility framework established for fluvial abrasion and grain drop experiments on rocks of different strengths,  $V_{rock} = c_{ero} \frac{\varepsilon_{kin,n}}{\sigma_{rock}^2}$  (with a bedrock erodibility conversion factor of  $c_{ero} = 3.8 \times 10^4 \text{ J/Pa}^2$ ; Beer and Lamb, 2021). Conversely, any massive bedrock as defined by its tensile strength can be applied within DGAM by multiplying grain impact erosivity (Equation 12). Compared to our used foam substrate,  $V_{rock}$  would shift to one order of magnitude higher abrasion rates for a weak sandstone ( $\sigma_{rock} = 0.1 \text{ MPa}$ ) or to four orders of magnitude lower abrasion rates for quartzite ( $\sigma_{rock} = 20 \text{ MPa}$ ; cf. the measured rock tensile strengths in Sklar and Dietrich, 2001).

There may be additional important tradeoffs between the erodibility of bedrock and the frequency and magnitude of rockfall events. For example, bedrock tends to be stronger in massive rock with low fracture density, like granite (cf. Figure 1E), which should slow rockfall erosion rates by reducing  $k_{ero}$ . In addition, granite also tends to weather into small grains, which would have low kinetic energy and therefore could reduce rockfall erosion rates further (Equation 9). In contrast, jointed rocks like sandstone or columnar basalt produce more intact rock blocks (cf. Figure 1A and B; Ward et al., 2011). Due to the more-than-linear dependence of abrasion on impactor size (Figure 9B, lower panels), fewer more massive rocks would produce more erosion than more frequent events with smaller rocks. These ideas could be incorporated in a future effort to describe the rockfall generation process, which is needed to drive the rockfall abrasion model.

#### 5.4. Rockfall erosion on low gradients

Our experiments and modelling confirm that bedrock hillslopes can be eroded by dry rockfall abrasion even below the angle of repose (Figure 3; DiBiase et al., 2017; Pelletier et al., 2008; Sun et al., 2021). Given energetic rockfall and low friction angles relative to the surface roughness (DiBiase et al., 2017), even small grain sizes are able to traverse rocky slopes (Figure 9B). As their impact energy is not diffused into granular debris like on granular substrate (Figure 1G), it contributes to rock fracturing and subsequent abrasion (Figure 1H). Thus, dry rockfall, as an endmember of dry granular avalanching (Howard, 1998), is an erosive process not restricted to steep alpine environments.

The abundance of rockfall on rocky slopes in both dry and humid areas permit to elucidate the absolute and relative contribution of rockfall-driven erosion to earthen and planetary surface evolution, so far generally ascribed to fluvial or aeolian erosion (e.g., Figure 1A to D). While in steeper areas rockfall may outpace other erosive processes and create indicative topographic features (cf. Howard and Selby, 2009), at the foot slopes of lower gradient, dry bedrock abrasion could set preferential routes for fluvial mass transport processes and this way enhance their channelization.

#### 5.5. Formation of rocky chutes

597 As shown, rockfall-prone hillslopes evolve into bumpy and channelized chute  
598 topography (Figure 3 and Figure 4A and D; Blackwelder, 1942), which steers grains into  
599 preferential pathways resulting in topographic feedback (Figure 4B and C; cf. Sun et al., 2021).  
600 This transient process was successfully reproduced by the dry grain abrasion model DGAM  
601 fitted with a fixed impact shock term coefficient,  $\kappa_{\text{shock}} = 0.8 \text{ m}^{-1}$  (Figure 10 and Figure S3).  
602 Improvement of this calibration could have been reached by better constraining the initial grain  
603 entrance conditions, though we took the approach of modeling the inlet conditions as random.  
604 The experimentally observed and modeled topographies generally resemble earthen and  
605 planetary rocky hillslope topography (Blackwelder, 1942), showing bedrock chutes and gully  
606 alcoves with downslope bumps and channels (Figure 3 and Figure 10 vs. Figure 1A to D). The  
607 lateral grain mobility (so far treated by probabilistic direction-sampling in DGAM, Table S1) was  
608 not retrievable from the vertical PIV camera in our experiments. Grain spread transience would  
609 help quantify the topographic steering feedback and its separation from diffusional processes  
610 (Jop et al., 2005; Williams and Furbish, 2021).

611 As long as rocky hillslopes remain free of cover (from regolith, saprolite, or vegetation),  
612 continuous and local dry grain abrasion will create rills that fuel a sequence of downhill-  
613 wandering troughs and rims (Figure 3B and 10A, B; Sun et al., 2021; cf. examples in Figure 1F  
614 and H), somehow an antipode to upstream-migrating knickpoints in (bedrock) rivers driven by  
615 fluvial sediment transport (Berlin and Anderson, 2007; Crosby and Whipple, 2006; DiBiase et al.,  
616 2015; Grimaud et al., 2016). Grain routing around sediment patches (Figures 1A and 3B) and  
617 grain deflection from elevated topography will enhance downhill channelization, which over  
618 time can lead to chutes (Figure 3A right panel, Figure 4C) or even gully channels (Figure 1D) by a  
619 self-enhancing process. Model-predicted topographies resembled both throughs (Figure 10A vs.  
620 Figure 3A right panel) and parallel rills (Figure 10B vs. Figure 3A central panel). The physical  
621 steering process of grains around resting sediment patches, as in the troughs of the  
622 experiments (Figure 3B), has not implicitly been implemented in the DGA model so far, and  
623 would require parameterizations of grain-grain interaction, grain piling (with varying angle of  
624 repose), and release mechanisms.

625 Talus cover from lower-sloping regions downhill reaching up onto the active abrasion  
626 area will suddenly terminate the process and seal the rocky surface due to rockfall grains  
627 starting to rest below their angle of repose, i.e., shielding a so-called sub-debris or Richter  
628 denudation slope below (a rectilinear,  $35.0^\circ$  thinly-covered rocky hillslope; Rapp, 1960).  
629 Termination will happen given short hillslope lengths, large amounts of simultaneous rockfall  
630 grains (i.e. dry grain avalanches), or low talus removal rates by other processes. Thus, there is  
631 potential that large talus cones or ramparts actually cover and hide channelized rocky slopes  
632 initially created by dry rockfall abrasion – a topic that could be verified by studying impact  
633 crater degeneration or escarpment retreat in dry planetary areas (Golombek et al., 2014; Ward  
634 et al., 2011).

### 635 5.5. Application to other planets

636 Dry grain abrasion modeling can generally be performed for any planetary body by  
637 adjusting gravitational acceleration. Though, there likely is no significant influence of this  
638 parameter on model mode II (grain rolling and sliding; not studied here; cf. Atwood-Stone and

639 McEwen, 2013), as there also is none on mode I (grain hopping) besides the influence on the  
640 acceleration of the grains during hopping. Air (or other gas) drag during the grain trajectories is  
641 neglected in the model since we deal with relatively low velocities and small, compact grains.  
642 Specifically, dry grain abrasion could be modeled in concert with other erosion processes (such  
643 as diffusion) to study the degradation of planetary crater walls, etc. (Golombek et al., 2014).  
644 This will help verify if dry bedrock abrasion is a reason why crater walls remain rocky or how  
645 low-sloping sinuous gully channels are maintained over time (Mangold et al., 2010).

646 Dry grain abrasion modeling on planetary surfaces is feasible considering rock or ice-  
647 cemented sediments using estimates of the substrate's tensile strengths (Beer et al., 2019). For  
648 example, low-fractured basaltic rock on Mars may have a tensile strength of  $\sigma_{rock} \sim 10$  MPa,  
649 which certainly is much lower at fractured impact craters (Wright et al., 2022; Figure 1D). Ice-  
650 cemented sediment near the melting point has tensile strengths similar to our applied foam  
651 ( $\sigma_{foam} = 0.1$  MPa), whereas colder permafrost can have tensile strengths again similar to basalt  
652 (Akagawa and Nishisato, 2009; Azmatch et al., 2010; Yuanlin and Carbee, 1987). Given dry  
653 regions on Earth, absolute dryness on the Moon, and current dry conditions on Mars (Figure 1A  
654 to D), together with abundant rocky hillslope areas and rockfall activity (Bickel et al., 2020a,  
655 2020b; Dickson and Head, 2009; Kumar et al., 2013; Vijayan et al., 2022; Xiao et al., 2013), the  
656 rockfall abrasion process has potential to be a local to regional sculptor of planetary hillslopes.  
657 Shattered rocky crater walls and caprock-topped badlands are ideal sites for the process to  
658 occur. The spatio-temporal imprint of dry rockfall abrasion, specifically its distinction from and  
659 interaction with fluvial processes (Figure 1G vs. H; Levin et al., 2022), remains to be studied in  
660 detail, both for Earth and planetary hillslopes.

661

## 662 **5 Conclusions**

663 Our experiments and modeling show that bedrock abrasion by dry, impacting rockfall can  
664 erode and in some cases channelize rocky hillslopes. The model captures the trends in the  
665 experiments to first order by including the physics of ballistic trajectories and a bedrock wear  
666 (abrasion) relation that depends on the surface-normal kinetic energy of the impactor. Erosive  
667 grains can hop on slopes even shallower than the angle of repose (at least down to  $20^\circ$ ), and  
668 thus contribute to landscape evolution in areas where fluvial and debris flow processes are  
669 thought to dominate. We found that increasing rockfall grain size has the most substantial  
670 effect to increase abrasion amounts due to a nonlinear relationship. Increasing hillslope  
671 gradient also caused faster erosion rates.

672 Hopping grains are routed around topographic highs, which steer grains trajectories in a  
673 self-enhancing feedback. First, a bumpy surface evolves with patches of immobile sediment  
674 collecting in lows (troughs) due to greater friction angles of grain piles. Around these piles and  
675 bedrock highs, shallow rills form, which coalesce into chutes and finally into emerging channels  
676 further downslope. These channels increasingly attract subsequent grains, focusing abrasion  
677 into their centers, and cause a sequence of troughs wandering downslope. The rockfall abrasion

678 process will terminate abruptly, where talus grows uphill from the toe of the hillslope or by  
 679 coalescence of local resting sediment patches.

680 Given abundant rocky hillslopes and rockfall sources from cliffs and outcrops (Blackwelder,  
 681 1942; Howard and Selby, 2009; Ward et al., 2011), dry impact-driven bedrock abrasion is a  
 682 conceivable contributor to Earth and planetary hillslope evolution. It could be important in high  
 683 mountain rockfall areas, dry climate scarpland retreat, and in planetary surface crater decay.  
 684 The model explicitly includes gravity and can be scaled to other planets.

685

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694

## 695 Data Accessibility

696 The experimental data and the dry gravel abrasion model DGAM code (in R language)  
 697 will be made publicly available at <https://data.caltech.edu> or at <https://fdat.uni-tuebingen.de/>.

698

## 699 Notation

$a_{grav}$	acceleration due to gravity [m/s <sup>2</sup> ]
$c_{ero}$	bedrock erodibility conversion factor [J/P <sup>2</sup> ]
$d_{cell}$	model cell size [m]
$d_{grain}$	rockfall grain diameter [m]
$h_{hop}$	grain hop height (trajectory maximum above crossed cell boundaries) [m]
$k_{ero}$	bedrock erosion factor (grain erosivity) [m <sup>3</sup> /J or ms <sup>2</sup> /kg]
$l_{flume}$	length of the laboratory flume [m]
$l_{hop}$	grain hop length [m]
$l_{hop,0}$	grain hop length at entrance of a grain into the model domain [m]
$m_{grain}$	rockfall grain mass [kg]
$m_{grains,tot}$	total mass of all grains in one experiment [kg]
$n_{cells}$	number of DEM cells traversed by a grain's trajectory [-]
$n_{grains,tot}$	total number of grains used in an experiment [-]
$n_{mps,tot}$	total number of grain impacts per grain down the laboratory flume [-]
$t_{hop}$	grain hop time [s]
$V_{flume}$	total volumetric foam abrasion of an experiment from grain impacts [m <sup>3</sup> ]
$V_{cell}$	volumetric abrasion of a cell by a grain impact [m <sup>3</sup> ]
$V_{rock}$	volumetric abrasion of a bedrock cell by a grain impact [m <sup>3</sup> ]
$v_{in}$	grain impact velocity [m/s]
$v_{in,0}$	grain impact velocity at entrance of a grain into the model domain [m/s]

720	$v_{in,n}$	surface-normal component of the grain impact velocity [m/s]
721	$v_{out}$	grain deflection velocity [m/s]
722	$v_{out,0}$	grain deflection velocity before entrance of a grain into the model domain [m/s]
723	$w_{cell}$	vertical cell abrasion or wear [m]
724	$X, Y, Z$	cell coordinate (X: downflume, Y: lateral, Z: vertical) [-]
725	$Z_{boundary}$	surface height at the boundary between two DEM cells [m]
726	$Z_{traj}$	grain trajectory height above a cell boundary [m]
727	$\alpha_{in}$	absolute grain impact angle [ $^{\circ}$ ]
728	$\alpha_{in,0}$	absolute grain impact angle at entrance of a grain into the model domain [ $^{\circ}$ ]
729	$\alpha_{out}$	absolute grain deflection angle [ $^{\circ}$ ]
730	$\alpha_{out,0}$	absolute grain deflection angle at entrance of a grain into model domain [ $^{\circ}$ ]
731	$\varepsilon_{kin}$	grain kinetic impact energy [J]
732	$\varepsilon_{kin,n}$	surface-normal component of the grain kinetic impact energy [J]
733	$\Delta t$	grain impact time [s]
734	$\Delta XY$	horizontal distance between two cells [m]
735	$\Delta Z$	vertical distance between two cells [m]
736	$\xi$	grain hop direction in D8 [-]
737	$\kappa_{shock}$	impact shock term [1/m]
738	$\sigma_{foam}$	tensile strength of the polyurethane foam [MPa]
739	$\sigma_{rock}$	tensile strength of bedrock [MPa]
740	$\theta_{cell}$	cell slope angle [ $^{\circ}$ ]
741	$\theta_{slope}$	hillslope angle or flume slope angle [ $^{\circ}$ ]
742	$\Phi_{surf}$	dynamic friction angle between grain and (bedrock) surface [ $^{\circ}$ ]
743	$\Phi_{pocket}$	grain pocket friction angle [ $^{\circ}$ ]
744	$\mu_{eff}$	effective grain friction angle [ $^{\circ}$ ]

745

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946   **Table 1: Outline of the dry grain abrasion experiments**, ordered by flume slope angle. Each experiment's data  
 947   symbol (as used in Figure 5, Figure 6, and Figure S4) refers to the relative size and shape of the used grains (large  
 948   vs. small, and round vs. angular). Erodible-bed experiments (*EB*) are denoted with grey background shading, and  
 949   fixed-bed experiments (*FB*) are of white background. More detailed measurements of the erodible bed experiments  
 950   *EB* are given in Table S2.

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 952   **Figure 1: Concept of rocky hillslope abrasion by dry rockfall:** Exemplary erosional rocky hillslope topography:  
 953   (A) plinth bedrock below a sandstone cap (Marble Canyon, AZ, USA), (B) chute channel in a basaltic lava flow  
 954   (Pan de Azúcar National Park, Chile), (C) basaltic bedrock gullies on Dawes crater walls on the Moon (Kumar et al.,  
 955   2013), and (D) furrowed Basalt bedrock gullies on Endurance Crater wall on Mars (google Mars). Exemplary sites  
 956   of dry grain transport over underlaying (E) granular substrate and (F) over bedrock substrate, San Gabriel  
 957   Mountains, CA. Conceptual sketches illustrate hillslope morphologies resulting from dry grain transport and erosion  
 958   over (G) gravel substrate and (H) over bedrock substrate.  
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960   **Figure 2: Dry grain abrasion model setup and experimental scheme:** (A) definitions of grain trajectory variables  
 961   in the dry grain abrasion model (DGAM; model scheme in Table S1; see notation section), (B) schematic of the  
 962   tilted flume filled with PU foam, sediment feeding and collection, terrestrial laser scanner (TLS) positions and  
 963   visual fields of particle imaging velocimetry (PIV), (C) picture in horizontal view on an initial smooth flume foam  
 964   surface, and (D) sample set of used dry, rounded rhyolite grains of  $d_{grain} = 0.03\text{m}$ .  
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966   **Figure 3: Abraded surface patterns of the erodible-bed experiments (EB):** (A) evolution of the foam surface  
 967   during *EB5* given at three temporal states, as indicated by the total grain mass run through until then ( $m_{grains,tot}$ ).  
 968   Color code is for vertical surface abrasion (note different range per panel). Contours denote abrasion depths in steps  
 969   of grain size ( $d_{grain} = 1.5\text{cm}$ ). The cleft to the bottom left in the central panel is an artifact due to missing surface  
 970   data. Three lateral (cross sections, cs) and one central long profile through the evolving surface of *EB5* are shown in  
 971   Figure 4A to C. (B) Grey-shaded surface meshes of the grain entrance area at the final experimental states, resulting  
 972   from different flume slope angles, grain sizes, and grain feed configurations: equal feed over the whole flume width  
 973   (*EB1* and *EB2*), central feed (*EB3* and *EB4*), and pointwise feed (*EB5*; cf. Table S2). Grain feed entrance directions  
 974   are indicated by the arrows, and flume constrictions for *EB3* and *EB4* are visible by the vertical black boards,  
 975   respectively. The upper flume bed section visible in (B) consisted of a fixed (non-abradable) board. Parallel blue  
 976   lines are horizontal (lateral) contours in 0.05m spacing, and yellow lines are vertical contours in 0.01m spacing.  
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978   **Figure 4: Transient topographic evolution of an erodible-bed experiment (EB5):** (A) cross sections through the  
 979   flume showing bed elevation below the initial surface for three experimental times and at three positions down the  
 980   flume ( $cs1=0.3\text{m}$ ,  $cs2=0.65\text{m}$ , and  $cs3=1.25\text{m}$ ; see Figure 3A left panel), (B) relative distribution of grains passing  
 981   through these cross sections around the three experimental times, (C) mean abrasion depth per grain impact on a  
 982   quadratic grain footprint (equals  $w_{cell}$  in the DGAM model), and (D) central long profile evolution down the flume  
 983   shown for several experimental times with indicated evolving topographic features (cf. Figure 3A right panel).  
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985   **Figure 5: Grain trajectory statistics of the experiments:** (A) a grain hop length  $l_{hop}$ , (B) grain hop height  $h_{hop}$ ,  
 986   (C) grain impact angle  $\alpha_{in}$ , and (D) grain impact velocity  $v_{in}$  against flume slope angle  $\theta_{slope}$ . Boxplots show  
 987   statistics given as described in the inset in (C). Data from the erodible-bed experiments (*EB1-EB3* and *FB5*) is  
 988   shown with grey shading, and data from fixed-bed experiments are of white background (*FB1-FB6*). Grain shapes  
 989   and relative grain sizes of the experiments are indicated as symbols above (A), (cf. symbol assignments in Table 1,  
 990   upper part). The mean values indicated by the dotted lines ( $\bar{l}_{hop}$ ,  $\bar{h}_{hop}$ ,  $\bar{\alpha}_{in}$ , and  $\bar{v}_{in}$ ) refer to the erodible-bed  
 991   experiments only, since the fixed-bed experiments likely started with higher initial grain velocities.  
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993   **Figure 6: Mean impact abrasion and grain erosivity of the erodible-bed experiments (EB) varying flume slope  
 994   angle:** (A) mean vertical impact abrasion,  $w_{cell}$  (i.e., abrasion volume of an impacting grain,  $V_{cell}$ , divided by cell  
 995   area  $d_{cell}^2 = d_{grain}^2$ ; Equation 9), (B) abrasion volume,  $V_{cell}$ , divided by the surface-normal component of the grain's  
 996   impact energy,  $\varepsilon_{in,n}$ , called grain erosivity,  $k_{ero}$ , and (C) these values further divided by the impacting grain's cross-  
 997   sectional area  $d_{cell}^2 = d_{grain}^2$  with two labeled values for rounded and angular grains, respectively. Relative symbol  
 998   size and shape are defined by used grain size and grain shape of the erodible-bed experiments (*EB*, symbols assigned  
 999   in Table 1, upper part). The grey-shaded area in the background denotes the common span for the angle of repose  
 1000   for grain piles.

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**Figure 7: Calibration of the DGAM model by selecting the shock term:** Predicted mean grain trajectory variables divided by measured mean trajectory variables for erodible-bed experiment 2 (*EB2*), plotted against the shock term ( $\kappa_{shock}$ ), which was varied in the modeling in steps of  $0.1\text{m}^{-1}$ . Unity on the y-axis means ideal model-reproduction of the measurements. The selected shock term (indicated by the vertical dotted line) was chosen at the first closest general agreement.

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**Figure 8: Reproduction of erodible-bed experiment (EB) grain trajectories with DGAM:** PIV-measured trajectory data and calibrated DGAM predictions along the flume for (A) grain hop length,  $l_{hop}$ , (B) grain impact velocities,  $v_{in}$ , (C) surface-normal grain impact velocity components,  $v_{in,n}$ , (D) grain deflection velocities,  $v_{out}$ , (E) absolute grain impact angles,  $\alpha_{in}$ , (F) absolute grain deflection angles,  $\alpha_{out}$ , and (G) grain impact cell abrasion volume,  $V_{cell}$ . Shown are measured PIV data for the fields of view of the three lateral PIV cameras (boxplots with median in grey, box-size of 50% interquartile range and whisker length of 1.5 times thereof, for data over 5cm bins downslope the flume) and mean DGAM predictions for 280 grains, equally sourced across the modeled flume width (black triangles). Data from erodible-bed experiment 2 (*EB2*; Table 1), PIV-camera positions are shown in Figure 2B.

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**Figure 9: Parameter space exploration for DGAM:** Model predictions vs. experimental data for (A) varying slope angle,  $\theta_{slope}$ , and keeping grain size constant ( $d_{grain} = 0.03\text{m}$ ), and (B) varying grain size,  $d_{grain}$ , and keeping slope angle constant ( $\theta_{slope} = 35^\circ$ ), respectively. Experimental PIV data is from the erodible-bed experiments (*EB*, boxplots with whiskers extending to 1.5 times the interquartile range from the box), and mean PIV data is from the fixed-bed experiments (*FB*, diamonds; Table 1). DGAM-predictions in (A) based on the erosivity-calibration for round grains (following the normalized grain erosivity  $0.001\text{m/J}$  in Figure 6C; bold blue lines), representative for *EB3*, *EB4*, and *FB1-FB6*, while prediction for angular grains measured on another grain size is shown for comparison (normalized erosivity  $0.003\text{m/J}$ ; *EB5* and thin yellow lines). DGAM-predictions in (B) based on erosivity-calibration for angular grains, representative for *EB5* and *FB4* (bold yellow lines), while prediction for round grains measured on another slope angle is shown for comparison (*EB3* and thin blue lines). The panels per row show grain hop length,  $l_{hop}$ , grain impact velocity,  $v_{in}$ , grain impact angles,  $\alpha_{in}$ , volumetric grain impact abrasion,  $V_{cell}$ , and local impact abrasion (i.e.  $V_{cell}$  divided by mean hop length,  $l_{hop}$ ), respectively. Abrasion for the fixed-bed experiments (*FB*, diamonds) was not measured but predicted based on the erodibility of experiment *EB3*.

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**Figure 10: Predicted transient hillslope channelization varying rockfall grain feed:** Simulations used varied grain feed patterns: (A) uniform central feed over 10 model cells, and (B) uniform feed in three inlets of 3, 4 and 3 cell widths (as indicated by the blue arrows on top of the third panels). DGAM-calibration was for erodible-bed experiment 2 (*EB5*; Table 1) with fixed  $m_{grains,tot} = 800\text{tons}$  of angular  $d_{grain} = 0.20\text{m}$  grains (normalized erosivity  $0.003\text{m/J}$ ; cf. Figure 6C) and a hillslope angle of  $\theta_{slope} = 35^\circ$ . Shown are stacked cross-sections (cs) through the transiently abraded hillslopes in a horizontal perspective, with initial (dotted), intermediate (i.e., half-time; grey), and final topography (black), respectively (more cross-sections are given in Figure S3). The lowest panels of (A) and (B) additionally show the transient lateral distribution of passing grains down the whole slope for the three experimental times (normalized number of transported grains; bin width is  $0.2\text{m}$ ). (C) and (D) show the central long profiles (lp) for both simulations with evolving troughs and rims; the position of the cross-sections of (A) and (B) are also indicated. Modeled topographies in panels (A, B) are comparable to the experimental topographies in Figure 4A and B, long profiles in panels (C, D) are comparable to Figure 4D.

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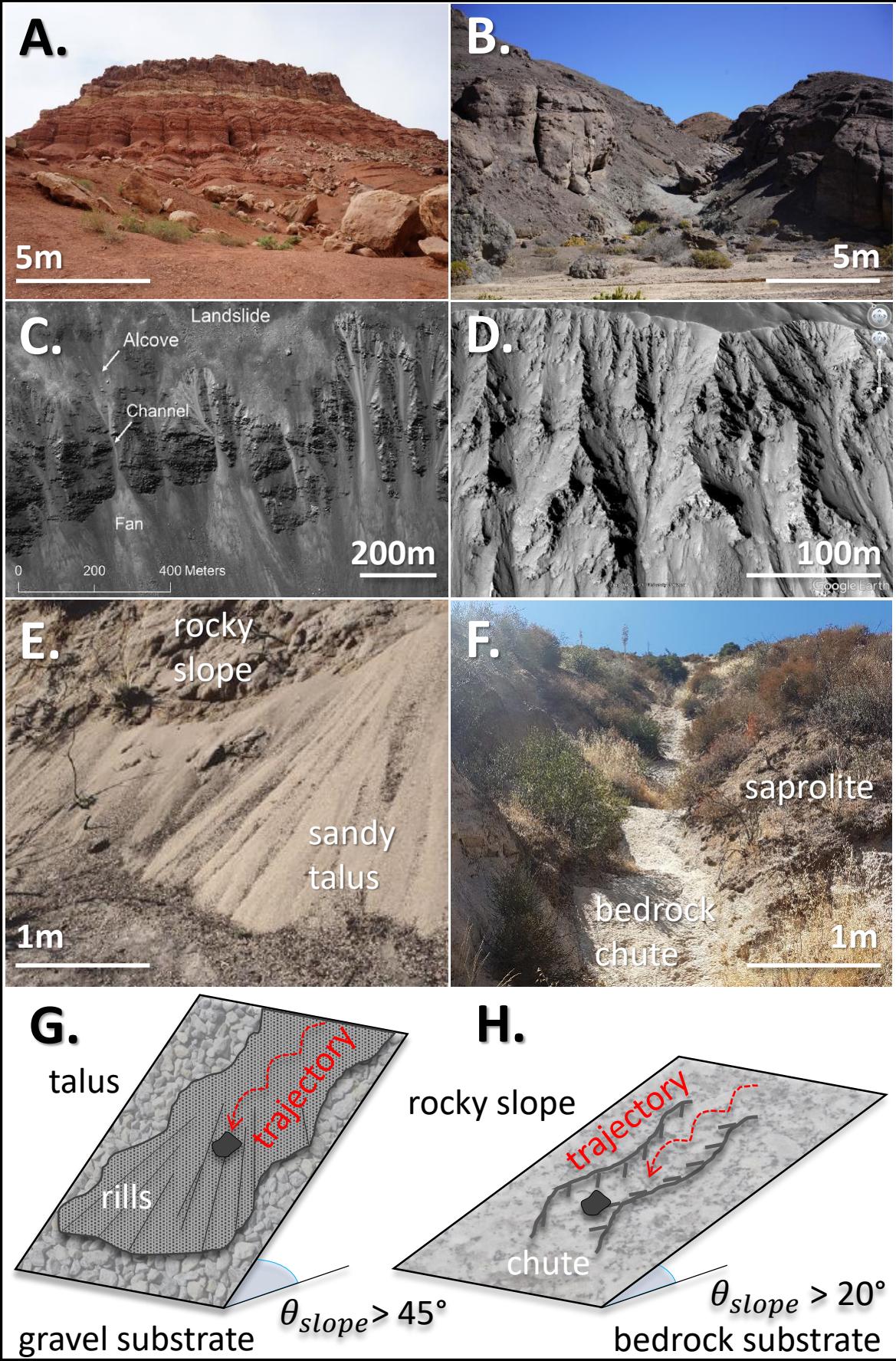
experiments		flume	dry grain properties						grain trajectory statistics <sup>†</sup>			foam abrasion		
erodible-bed	fixed-bed	symbol in Figures	slope, $\Theta_{slope}$ [°]	type	shape	surface friction angle, $\Phi_{surf}$ [°]	pocket friction angle, $\Phi_{pocket}$ [°]	size, $d_{grain}$ [m]	total grain mass, $m_{grains,tot}$ [tons]	relative impact angle <sup>#</sup> , $\alpha_{in}$ [°]	impact velocity, $v_{in}$ [m/s]	hop distance, $I_{hop}$ [m]	total abrasion volume, $V_{flume}$ [m <sup>3</sup> ]	grain impact erosivity, $k_{ero}$ [cm <sup>3</sup> /J]
EB1*	●	16.7	Granite	20	rounded	34	34.7	0.061	22.5	9.5	1.1	0.16	0.090	2.65
EB2	○	19.5	Andesite (tumbled)	23.5	40.3	23.5	44.4	0.023	8.3	9.0	1.2	0.11	0.046	1.00
EB3	○	35.0	Granite	angular	33.4	55.6	44.4	0.030	2.7	9.8	1.8	0.09	0.013	0.30
EB4	○	35.0	Andesite	(tumbled)	35.0	35.0	35.0	0.036	5.1	10,5 <sup>^</sup>	0,8 <sup>^</sup>	0,07 <sup>^</sup>	0.017	1.38
EB5	□	35.0	Granite	angular	33.4	55.6	55.6	0.015	16.5	10.5	1.6	0.16	0.040	0.68
FB1	○	20	Andesite (tumbled)	rounded	23.5	44.4	44.4	0.036	2.6e-4	6.0	1.9	0.13	-	-
FB2	○	25								16.9	2.1	0.20	-	-
FB3	○	30								2,7 <sup>^</sup>	2,2 <sup>^</sup>	0,22 <sup>^</sup>	-	-
FB4	○	35								19,1 <sup>^</sup>	2,5 <sup>^</sup>	0,32 <sup>^</sup>	-	-
FB5	○	40								29.8	2.4	0.23	-	-
FB6	○	45								35.5	2.7	0.24	-	-

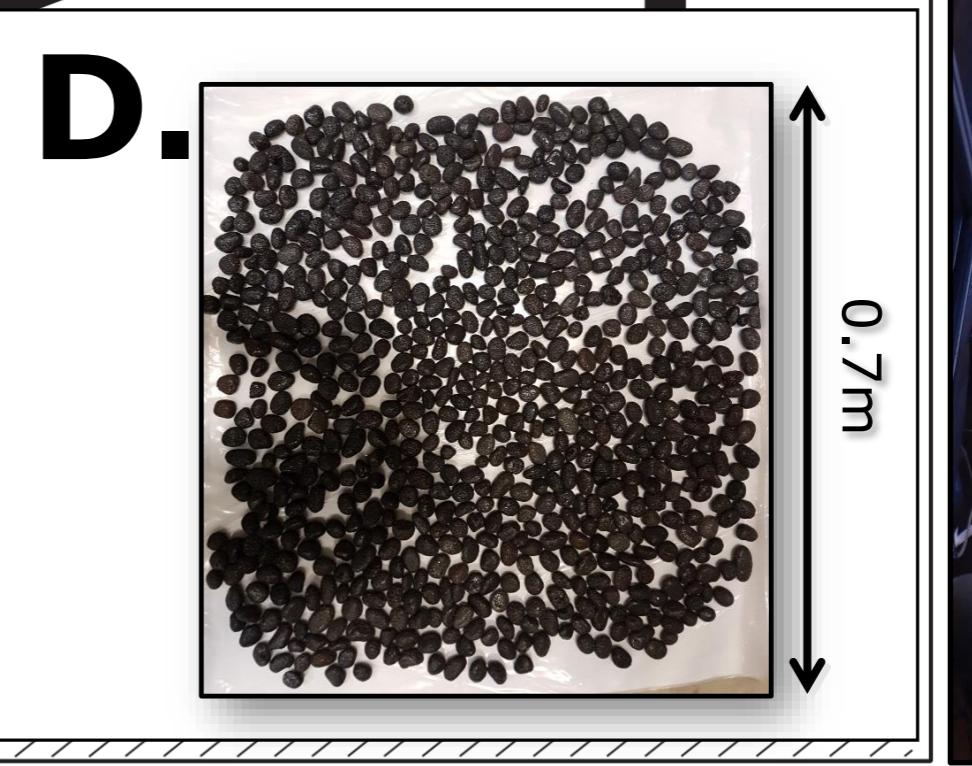
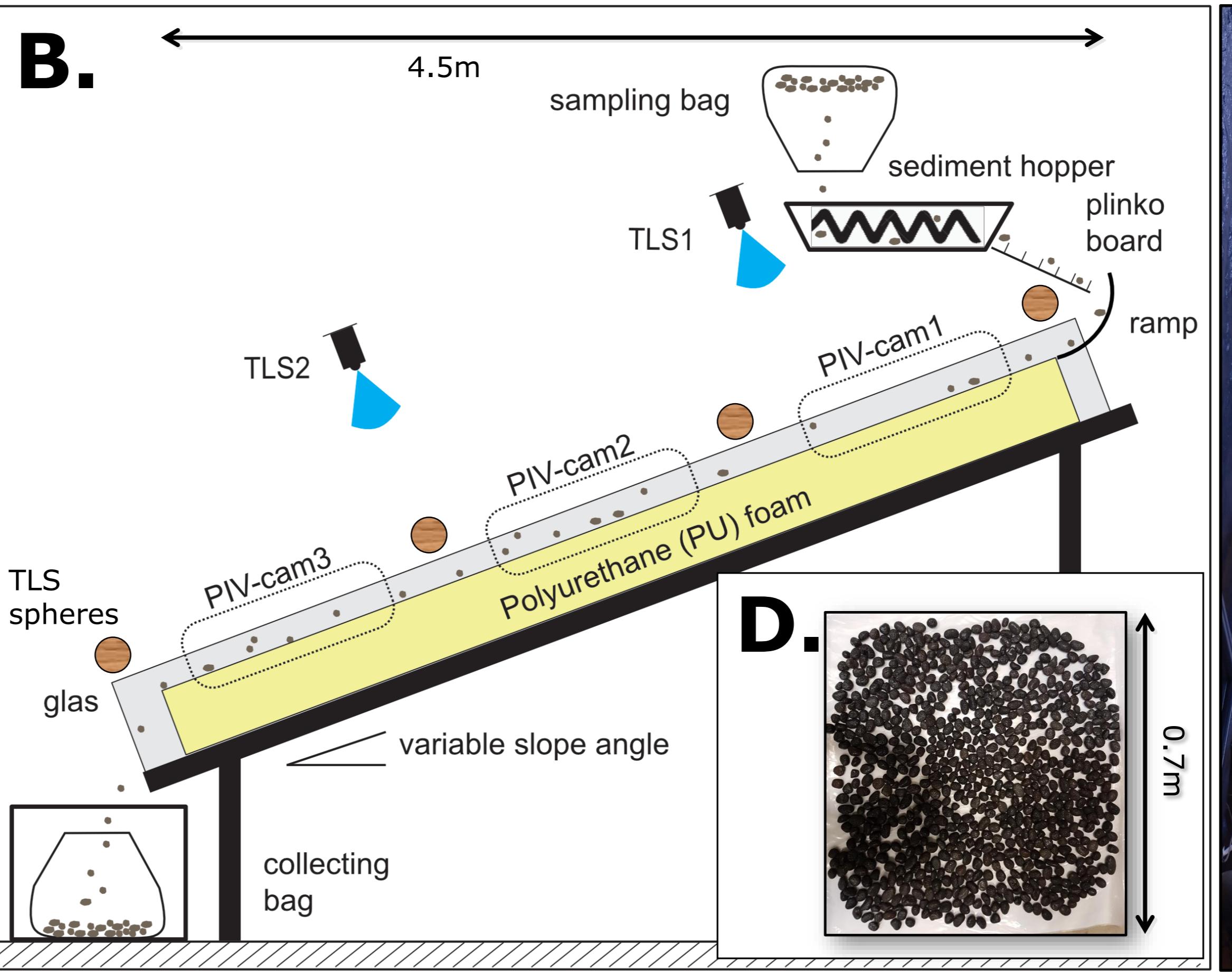
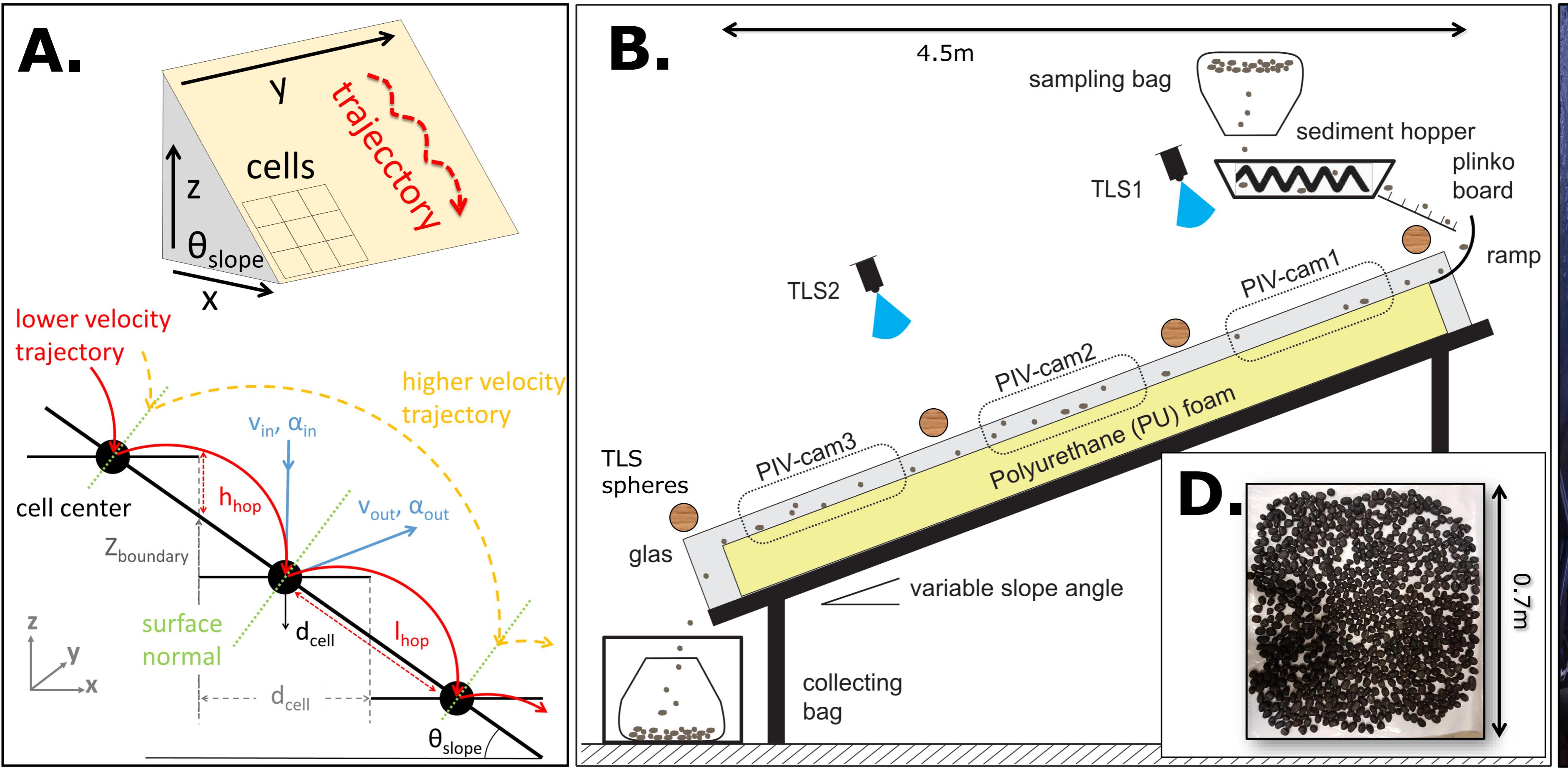
\* data from Sun et al., 2021

<sup>†</sup> data from the second lateral camera (i.e. PIV-cam2, central along the flume, neither at the inlet nor at the outlet; Figure 2B)

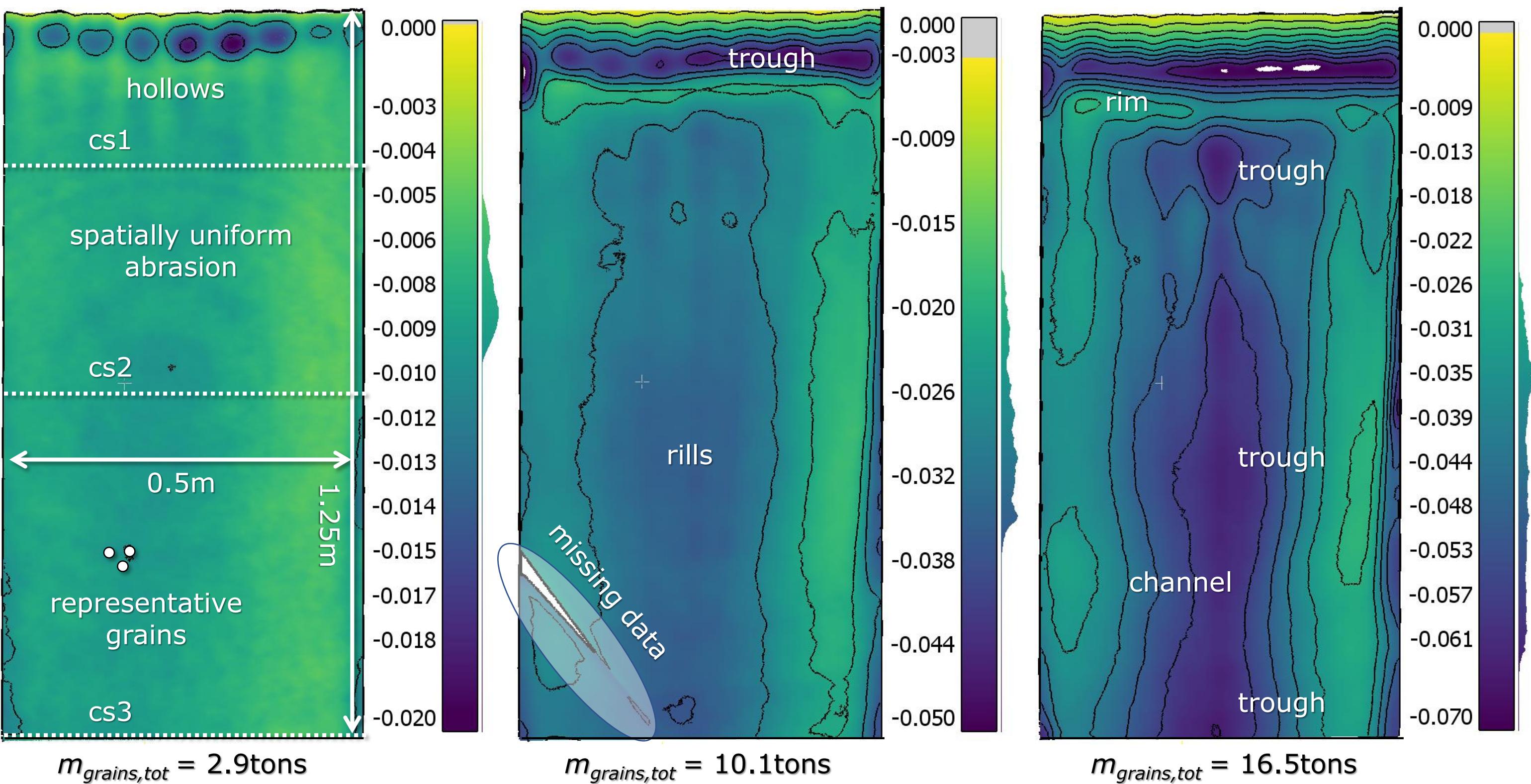
<sup>#</sup> angles are relative to the flume surface (i.e. they are not corrected for the flume slope  $\Theta_{slope}$ )

<sup>^</sup> uncertain data (few measurements)





# A. transient topography of EB5



# B. final erodible-bed topographies

