

Effects of geomagnetic storm on equatorial ionization during 27 February-1 March, 2014

DIBYENDU SUR¹, Omar Hammou Ali², Idahwati Binti Sarudin³, Joanna Rupiewicz⁴, Manuel Bravo⁵, Lekso Toriashvili⁶, and Xingxin Sun⁷

¹Institute of Radio Physics and Electronics, University of Calcutta

²University of Sciences and Technology Houari Boumediene

³Universiti Kebangsaan Malaysia

⁴European Satellite Service Provider Service Provision Unit Mission

⁵Universidad de Santiago de Chile Libertador General Bernardo O'Higgins

⁶Ilia State University, Abastumani Astrophysical Observatory

⁷China Research Institute of Radiowave Propagation

November 21, 2022

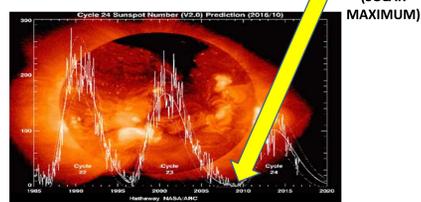
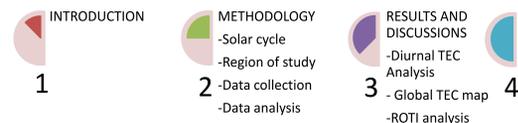
Abstract

The paper inspects the effects of a G2 class geomagnetic storm that occurred during 27 February- 1 March, 2014 on the equatorial ionization. This storm is observed following a Coronal Mass Ejection (CME) from a sunspot AR1967 on 26 February. An enhancement of solar wind speed is observed on 27 February, 2014 (483 km/sec). The maximum southward component of Interplanetary Magnetic Field (IMF) is observed around 21 UT of 27 February (12 nT). This interconnects with Earth's magnetic field and develops the main phase of a geomagnetic storm on the same day. The storm continues through 28 February and quiet-time ionospheric condition is recovered on 1 March. The effects of the storm on equatorial ionization is observed at Brasilia (15.95°S, 47.88°W geographic; 9.40°N, 21.13°E geomagnetic), Addis Ababa (9.04°N, 38.77°E geographic; 0.18°N, 110.47°E geomagnetic) and Colombo (6.89°N, 79.87°E geographic; 1.57°S, 151.57°E geomagnetic). An enhancement of TEC is observed during main phase of the geomagnetic storm at these stations. Increment in diurnal peak is observed on 28 February (14 TECU at 10 UT at Addis Ababa) while a decrement of diurnal peak TEC is observed during the recovery phase of the storm (15 TECU at 10 UT at Addis Ababa). Post-sunset ionospheric scintillation is inhibited at Brasilia on 28 February.

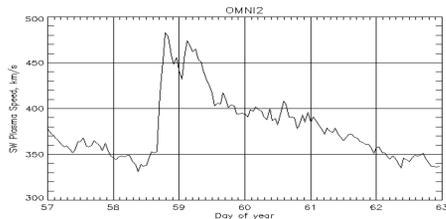
EFFECTS OF GEOMAGNETIC STORM ON EQUATORIAL IONIZATION DURING 27 FEBRUARY-1 MARCH, 2014

Dibyendu Sur, Institute of Radio Physics and Electronics, University of Calcutta, India, dibyendumalay@gmail.com
 Omar Hammou Ali, University of Sciences and Technology Houari Boumediene, Algeria
 Idahwati Binti Sarudin, Universiti Kebangsaan Malaysia, Malaysia
 Joanna Rupiewicz, European Satellite Service Provider Service Provision Unit Mission
 Manuel Bravo Sepulveda, Universidad de Santiago de Chile Libertador General Bernardo O'Higgins, Chile
 Xingxin Sun, China Research Institute of Radiowave Propagation, China
 Lekso Toriashvili, Iliia State University, Abastumani Astrophysical Observatory, Georgia

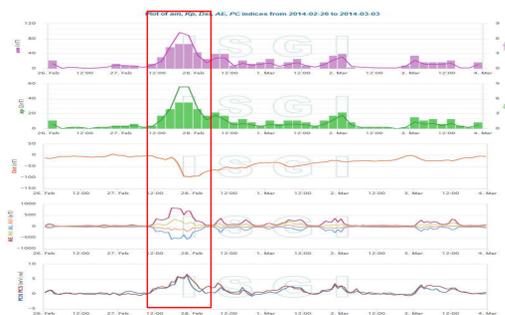
The Work is conducted at Workshop on Space Weather Effects on GNSS Operations at Low Latitudes 2018, 23 April – 4 May, 2018 at International Center for Theoretical Physics (ICTP), Italy



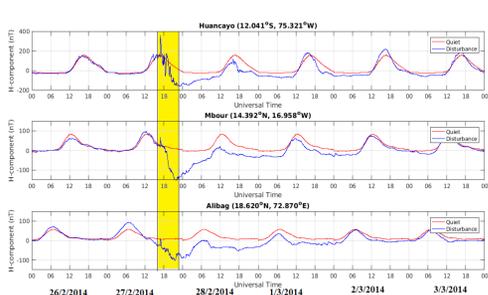
Variation of solar wind speed during 26 Feb – 3 Mar 2014
 Source: <https://omniweb.gsfc.nasa.gov>



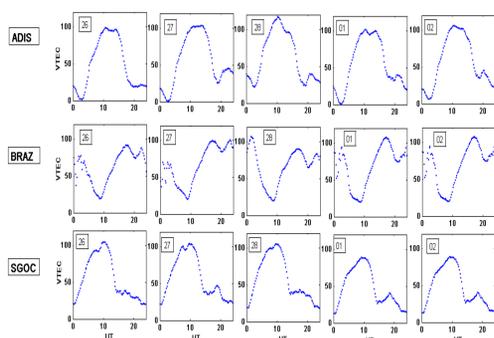
Magnetic indices during 26 February – 3 March, 2014
 Source: http://isgi.unistra.fr/geomagnetic_indices.php



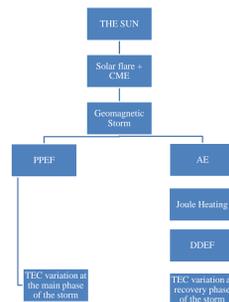
Variation of H-component of earth's magnetic field during 26 February-3 March 2014



Comparison of diurnal VTEC over the duration of 26 February - 2 March, 2014 from 1) Adis Ababa (Ethiopia) 2) Brasilia (Brazil) 3) Colombo (Sri Lanka)



Outline



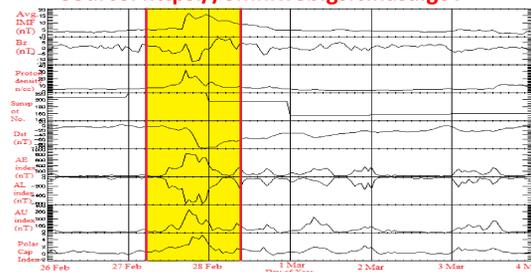
Abstract Number: SA31B-3418
 Abstract ID: 413612

CME from sunspot ar1967 observed on 26 February 2014 and G2 class magnetic storm on 27 February 2014



Source: www.spaceweather.com

Solar and magnetic indices during 26 February – 3 March, 2014
 Source: <https://omniweb.gsfc.nasa.gov>



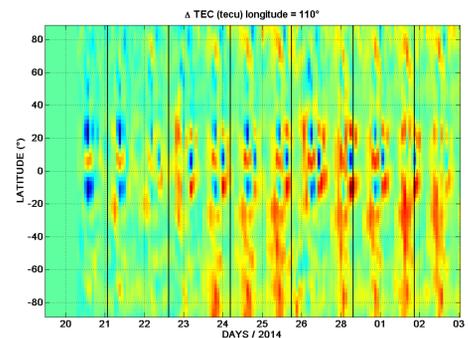
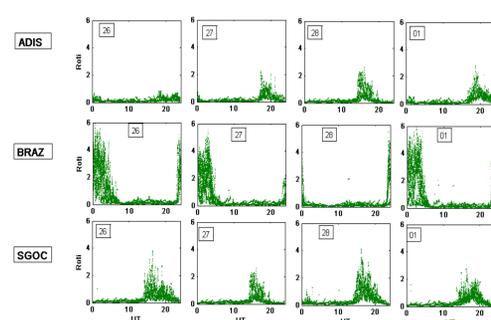
Region of Study



Station	Country	Lat_Geogr	Long_Geogr	Lat_Geomag	Long_Geomag
Braz	Ethiopia	-15,95	-47,88	9,40	21,13
Adis	Brazil	9,04	38,77	0,18	110,47
Sgoc	Sri Lanka	6,89	79,87	-1,57	151,57

Comparison of diurnal ROTI Index over the duration of 26 February - 2 March, 2014 from 1) Adis Ababa (Ethiopia) 2) Brasilia (Brazil) 3) Colombo (Sri Lanka)

Latitudinal variation of δ TEC (TECU) during 20 February – 3 March, 2014 along 110° longitude

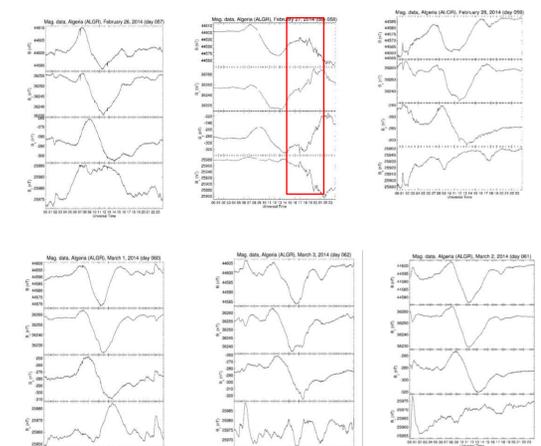


Methodology

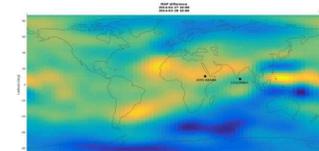
Data	Resources
- MAGNETIC INDICES	https://omniweb.gsfc.nasa.gov http://isgi.unistra.fr/geomagnetic_indices.php http://magnetometers.bc.edu/index.php/78-magnetometers/78-home
- SOLAR INDICES	https://solarscience.msfc.nasa.gov/SunspotCycle.shtml https://omniweb.gsfc.nasa.gov
- TEC DATA	http://www.igs.org/about/data-centers ftp://cddis.gsfc.nasa.gov/gnss/data/daily/

Variation of geomagnetic field during 26 February - 2 March, 2014

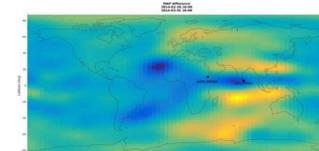
Samba-amber Magnetometers Data center
 Algeria (28°n 3°e geographic 15.23°n, 76.58°e Geomagnetic)
 Variation of b, bz, Bx, by indices with Universal time



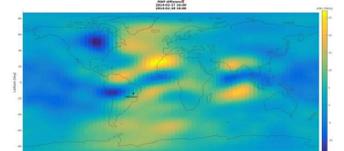
Global TEC difference on 10 UT (between 27 and 28 February, 2014)



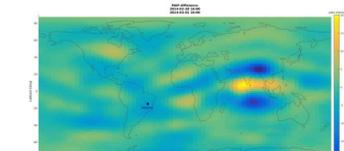
Global TEC difference on 10 UT (between 28 February, and 1 March 2014)



Global TEC difference on 16UT (between 27 and 28 February, 2014)



Global TEC difference on 16 UT (between 28 February, and 1 March 2014)



Conclusion

A geomagnetic storm is observed (main phase is on 17-23 UT, 27 February, 2014). TEC Enhancement is observed at the main phase of geomagnetic storm (starting 17 UT, 27 February, 2014). Increment of diurnal peak TEC is observed on 28 February, 2014. Diurnal TEC peak decreases on 1 March 2014, at the recovery phase of geomagnetic storm.